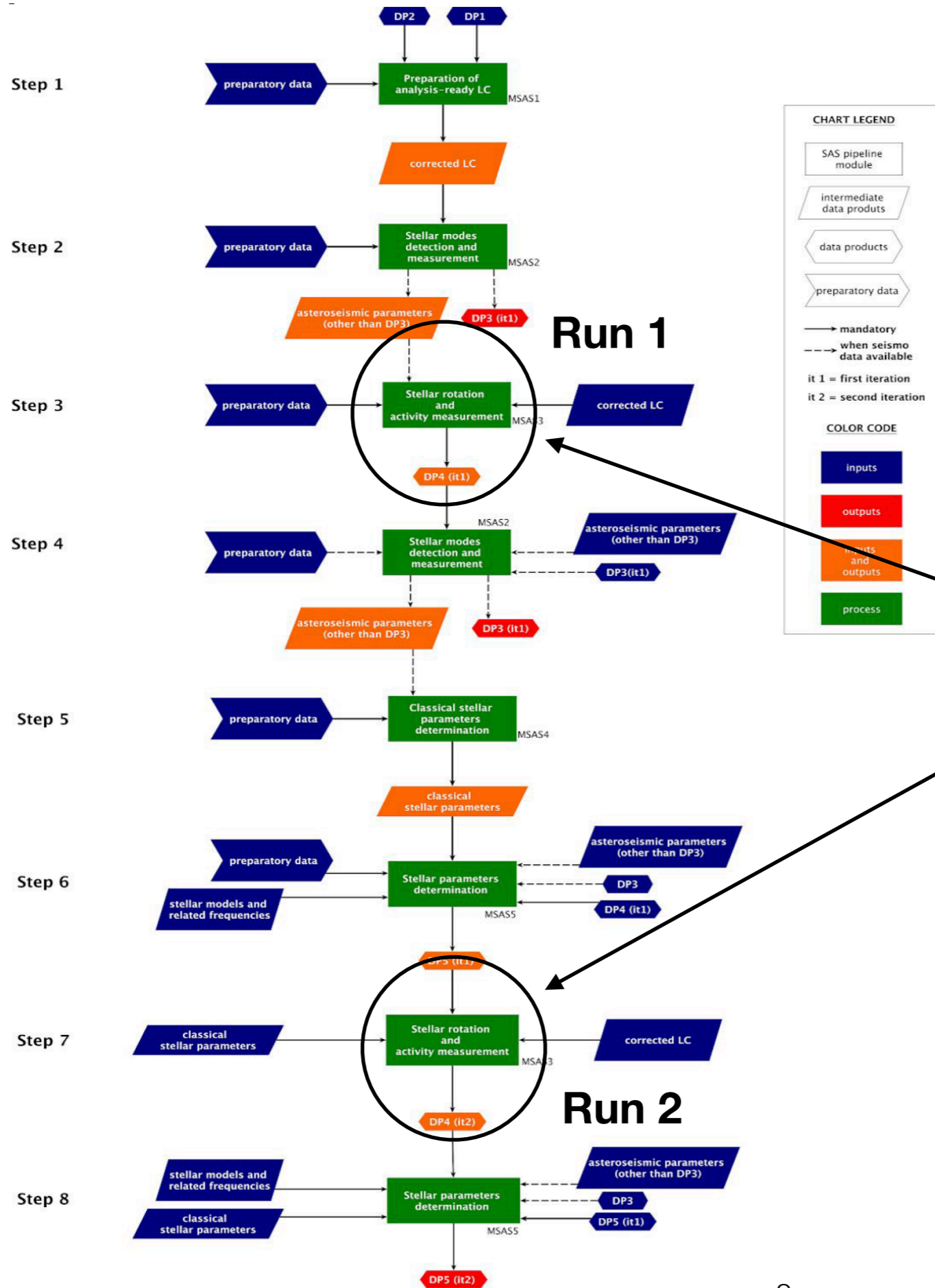


Overview of the MSCOPP3 for DP4

[Overview of the pipeline for rotation and activity measurements]

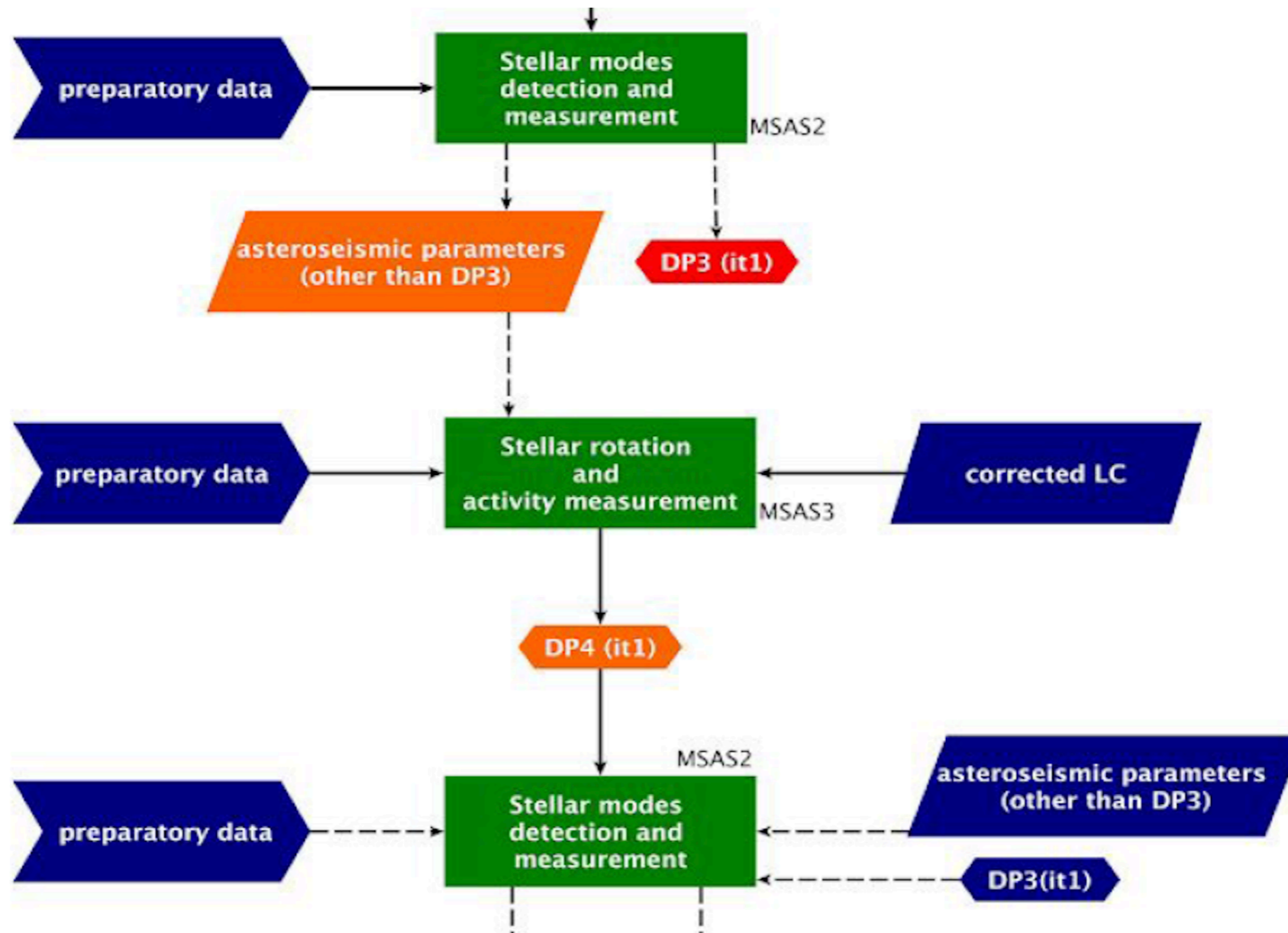
F. Baudin, N. Lanza and the WP120 team

Context



You are here

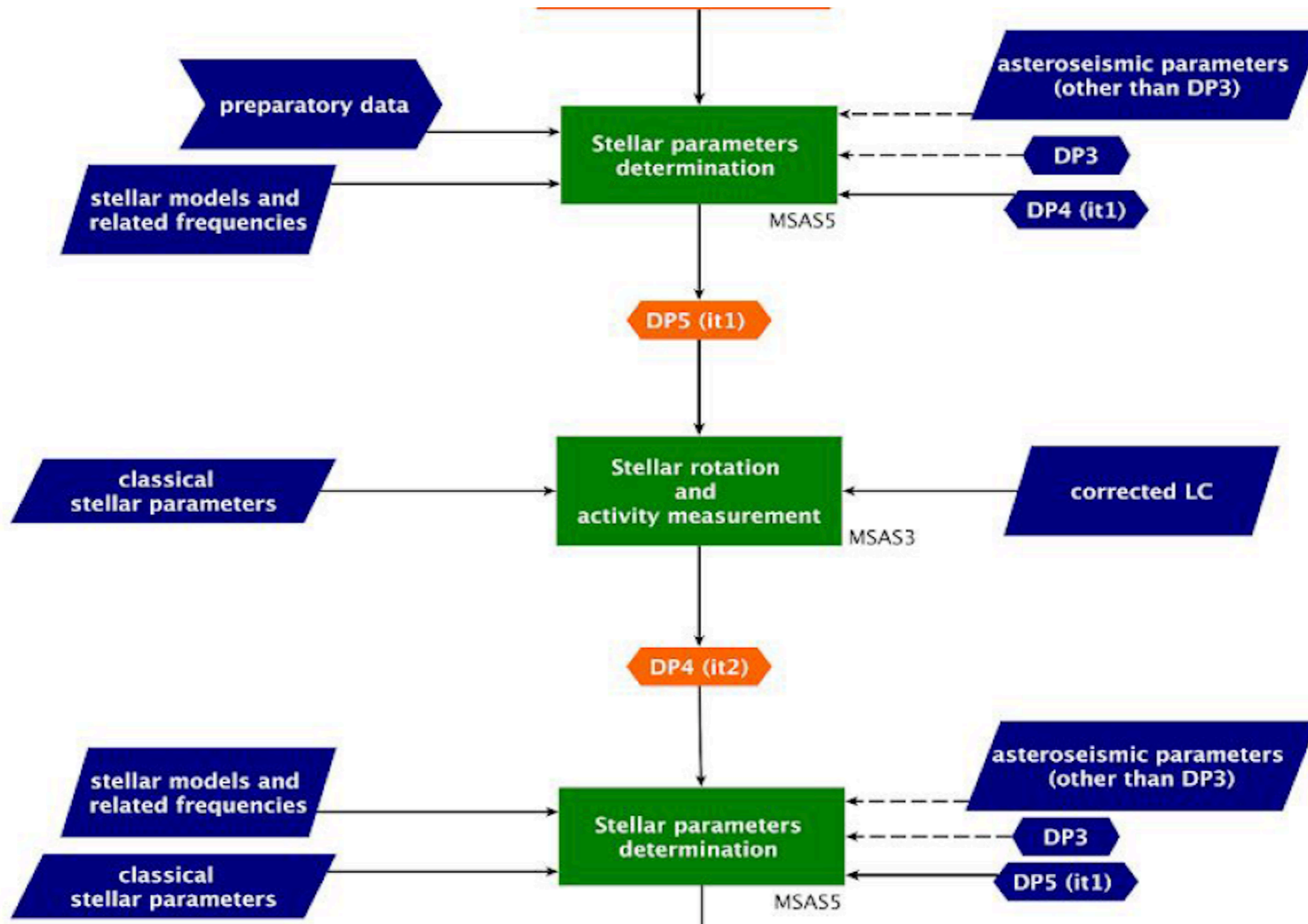
Context: Run 1



Step 3

Context: Run 2

Step 7



Let's start with the end: the Data Products 4

DP4_123_PROT	Rotation period	
DP4_123_DELTAPROT	Differential rotation	Amplitude of the surface differential rotation inferred from spot modelling
DP4_123_PCYCLE	Cycle period	Inferred from spot modeling (?)
DP4_123_HARVEYN_AMPLITUDE DP4_123_HARVEYN_TIME DP4_123_HARVEYN_EXPONENT	Harvey profiles characteristics (amplitude, width,	(fitted) Profiles modelling the activity in the Fourier background
DP4_123_WHITE-NOISE-FOURIER	White noise in the Fourier spectrum	Fitted in the spectrum

DP come with error bars or PDF

Let's start with the end: the Intermediate Data Products 4

IDP_123_VARLC

(Non-seismic) Variability Analysis-ready LC

IDP_123_MOD-LC-SPOT

Modelled LC reconstructing the variations due to spots

IDP_123_MOD-TAU-SPOT

Mean spots lifetime estimates from spot modelling
(+ other characteristics? position, etc...)

IDP_123_LOGG-GRANULATION

log g determined from the granulation profile

Let's start with the end: the Additional Data Products 4

ADP_123_INCLINATION-ANGLE-SPOT

Estimate from spot modelling

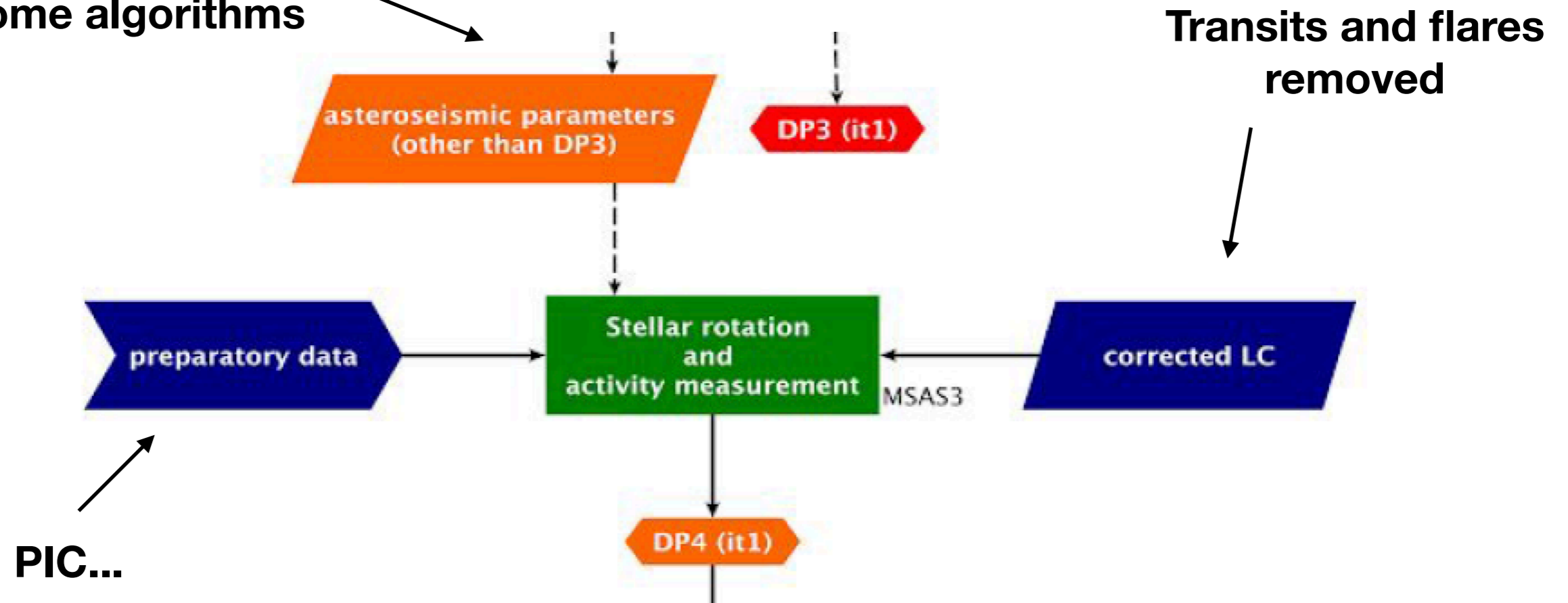
ADP_123_AGE-GYROCHRONOLOGY

Age determined from gyro chronology

+ list of flares? (talk by B. Stelzer)

Run 1: the inputs

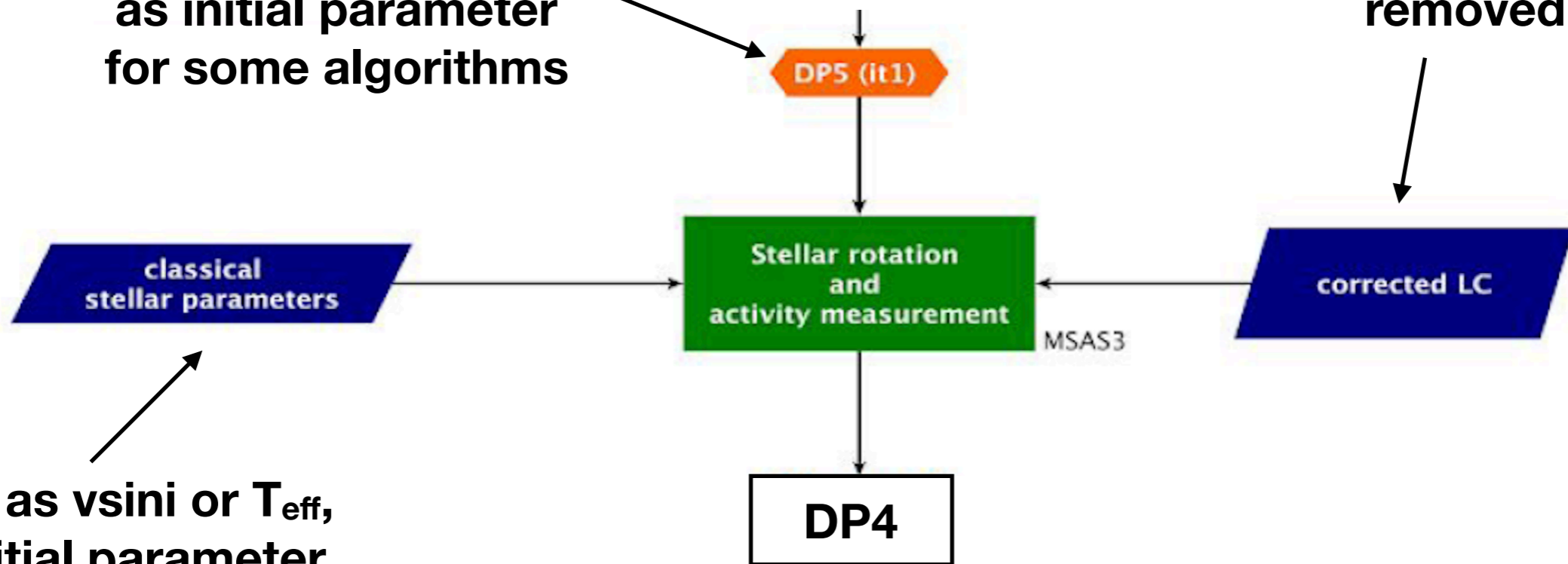
Such as mass or radius
(or from Input Catalog),
as initial parameter
for some algorithms



Run 2: the inputs

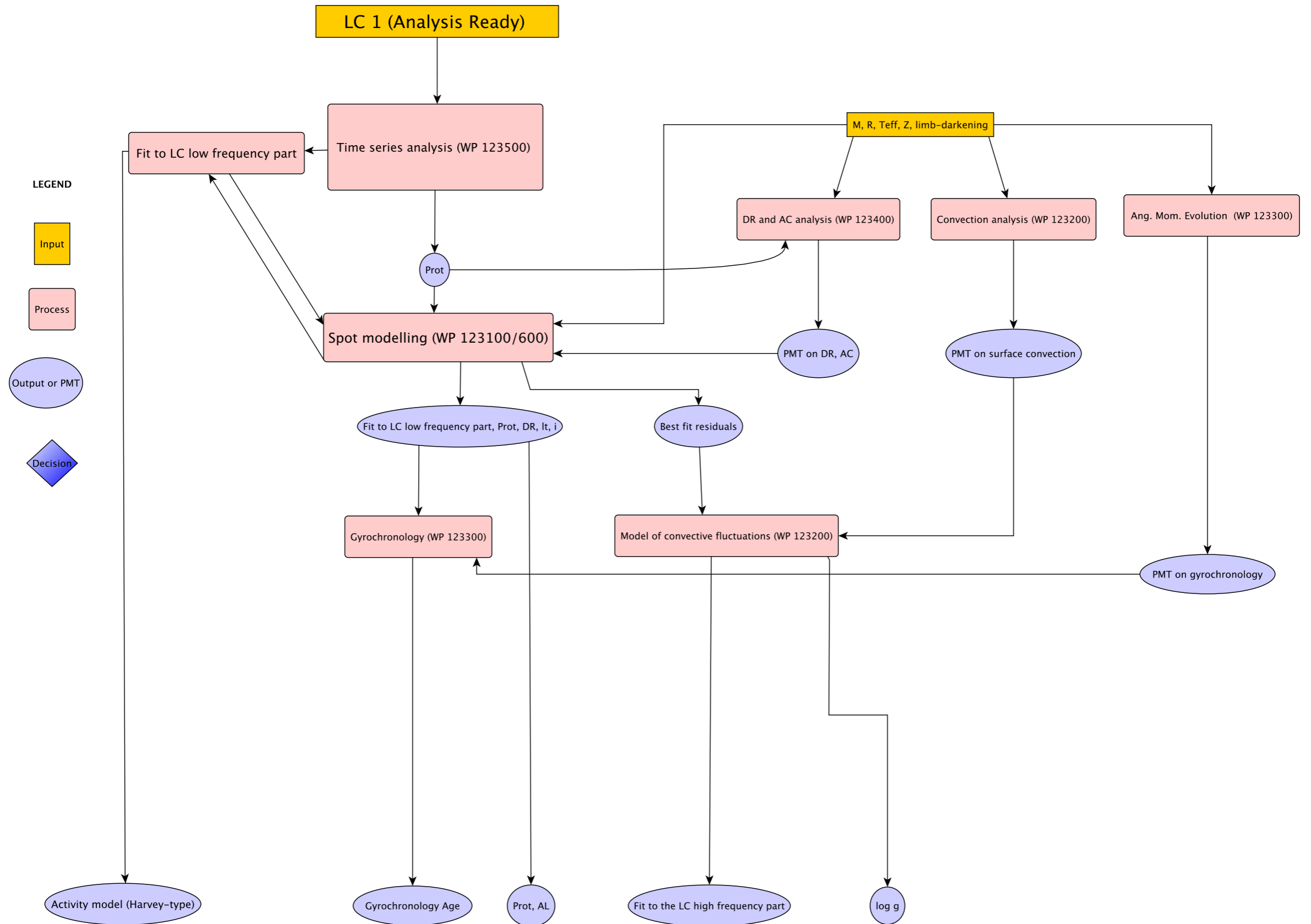
Such as mass, radius,
inclination angle,
as initial parameter
for some algorithms

Transits and flares
removed

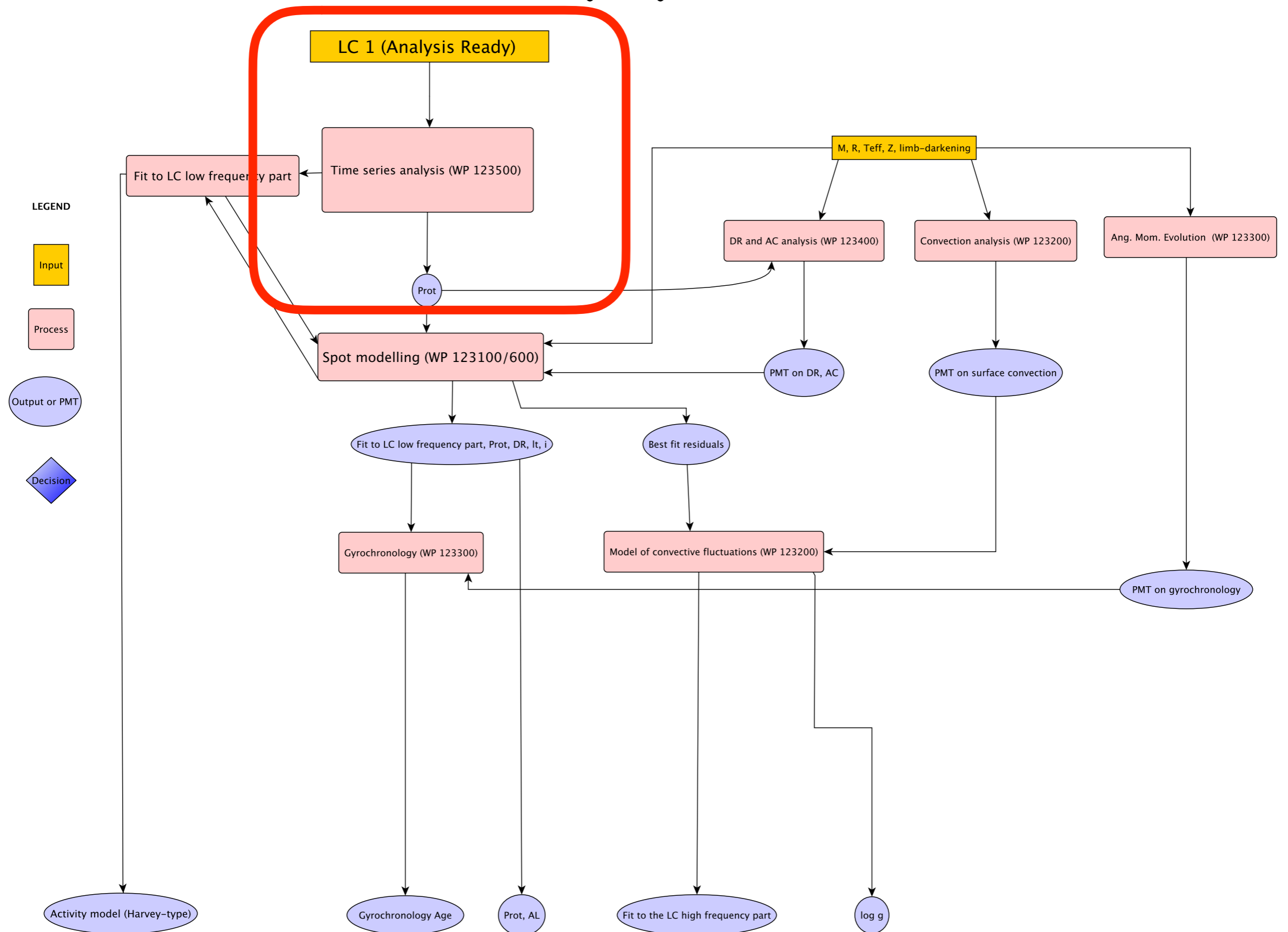


Such as $v \sin i$ or T_{eff} ,
as initial parameter
for some algorithms

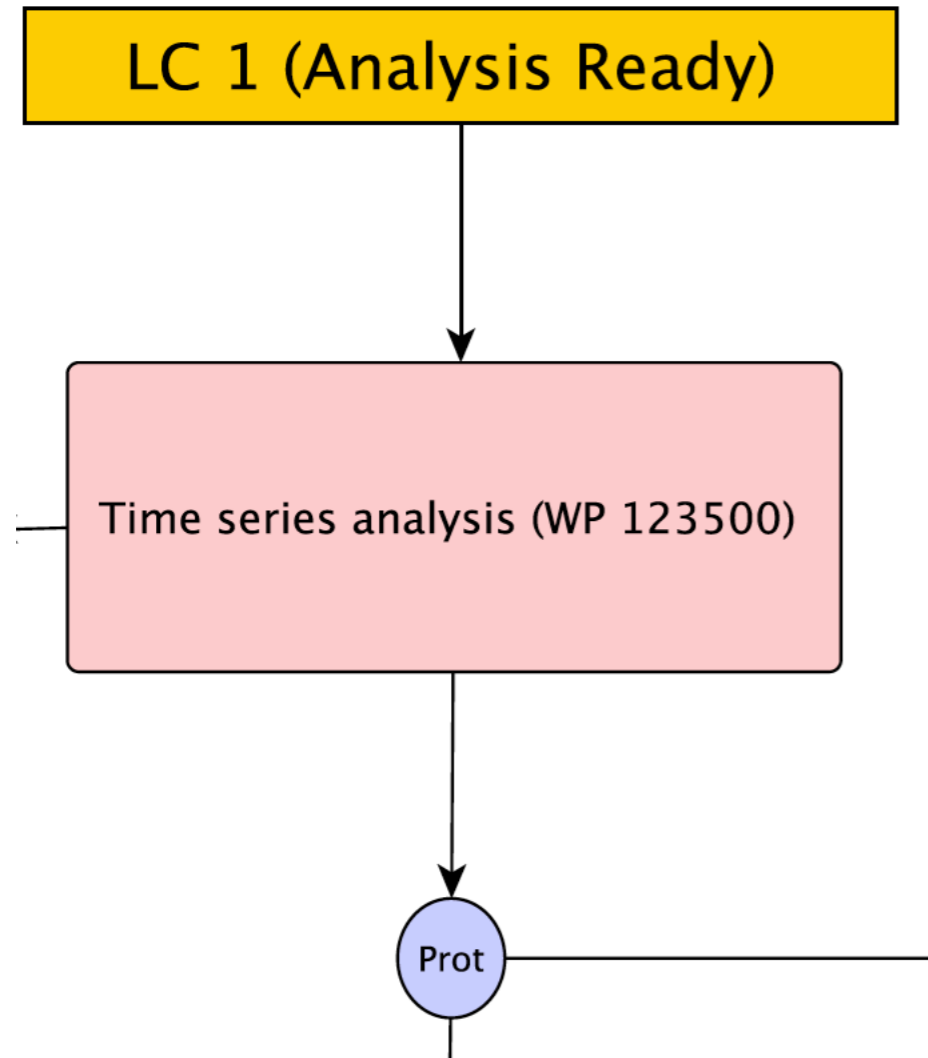
The pipeline



The pipeline



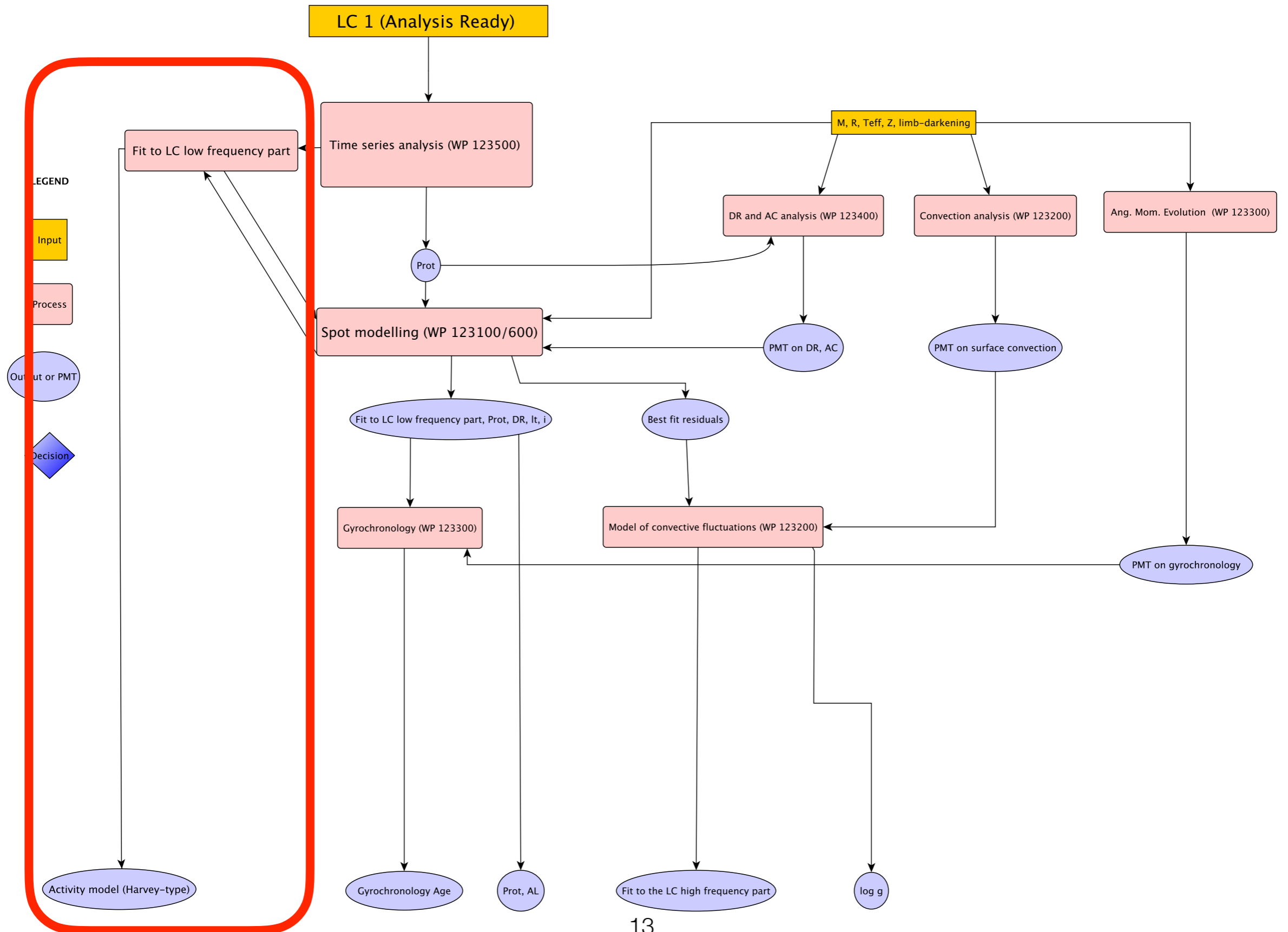
The pipeline



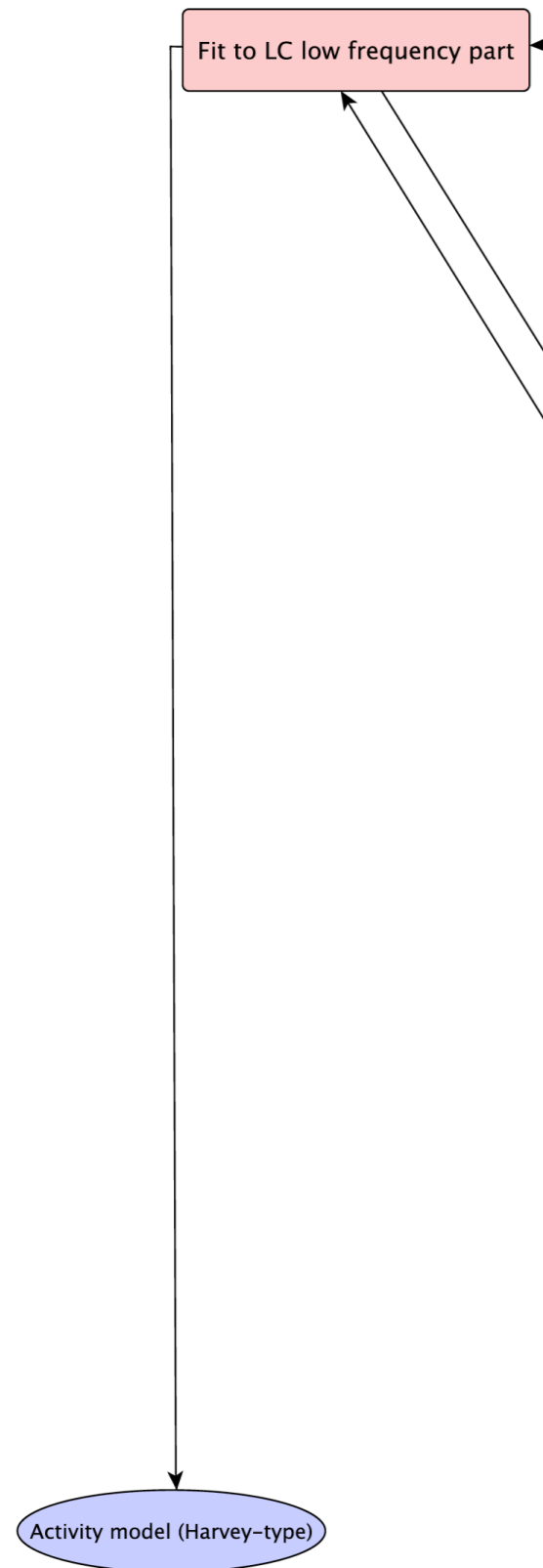
WP123500
Several existing and reliable methods
(from spectrum, ACF...)
See TN by Nuccio and talk by S. Messina

=> NMW
(No Much Worries)

The pipeline



The pipeline

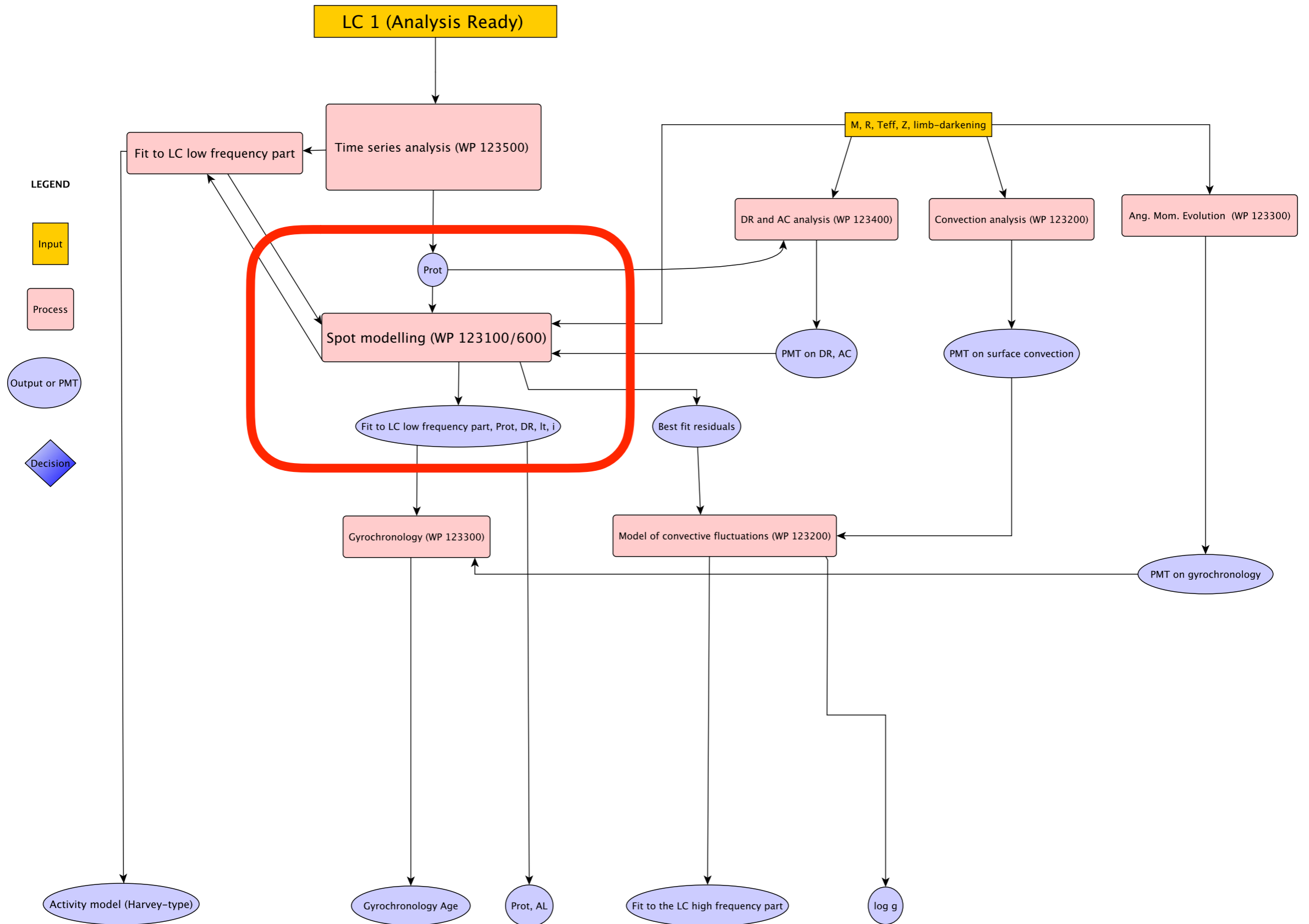


WP123500
Simple fit, performed without too many problems (?)

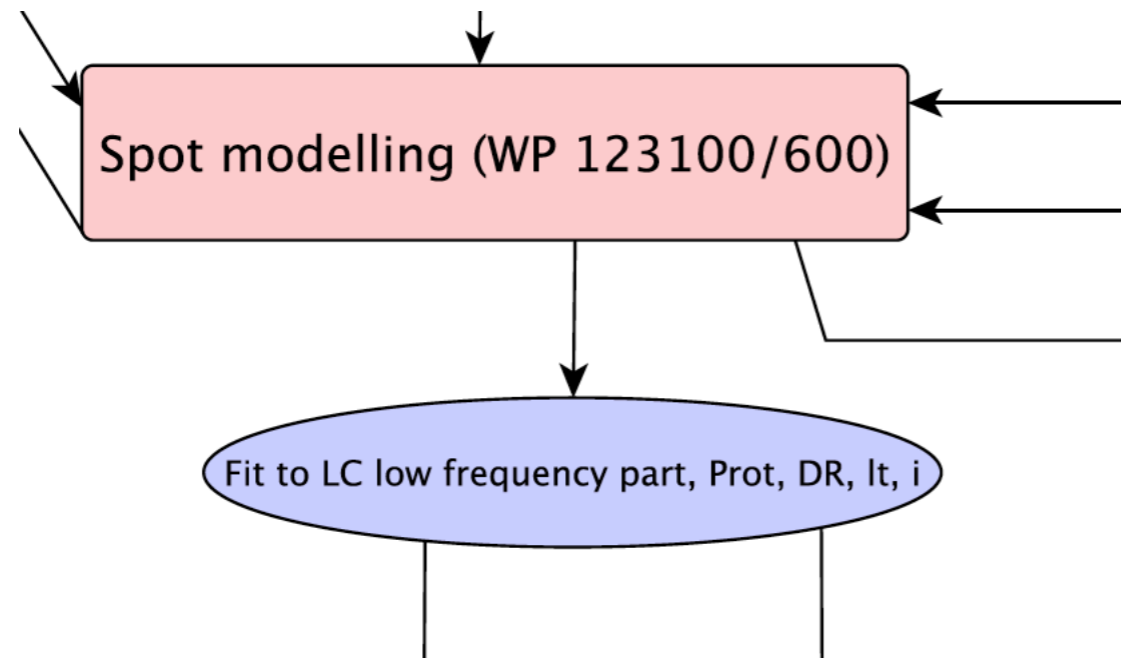
Useful in Run 1 (?)
Possible comparison with spot modeling in Run 2

=> NW
(No Worries)

The pipeline



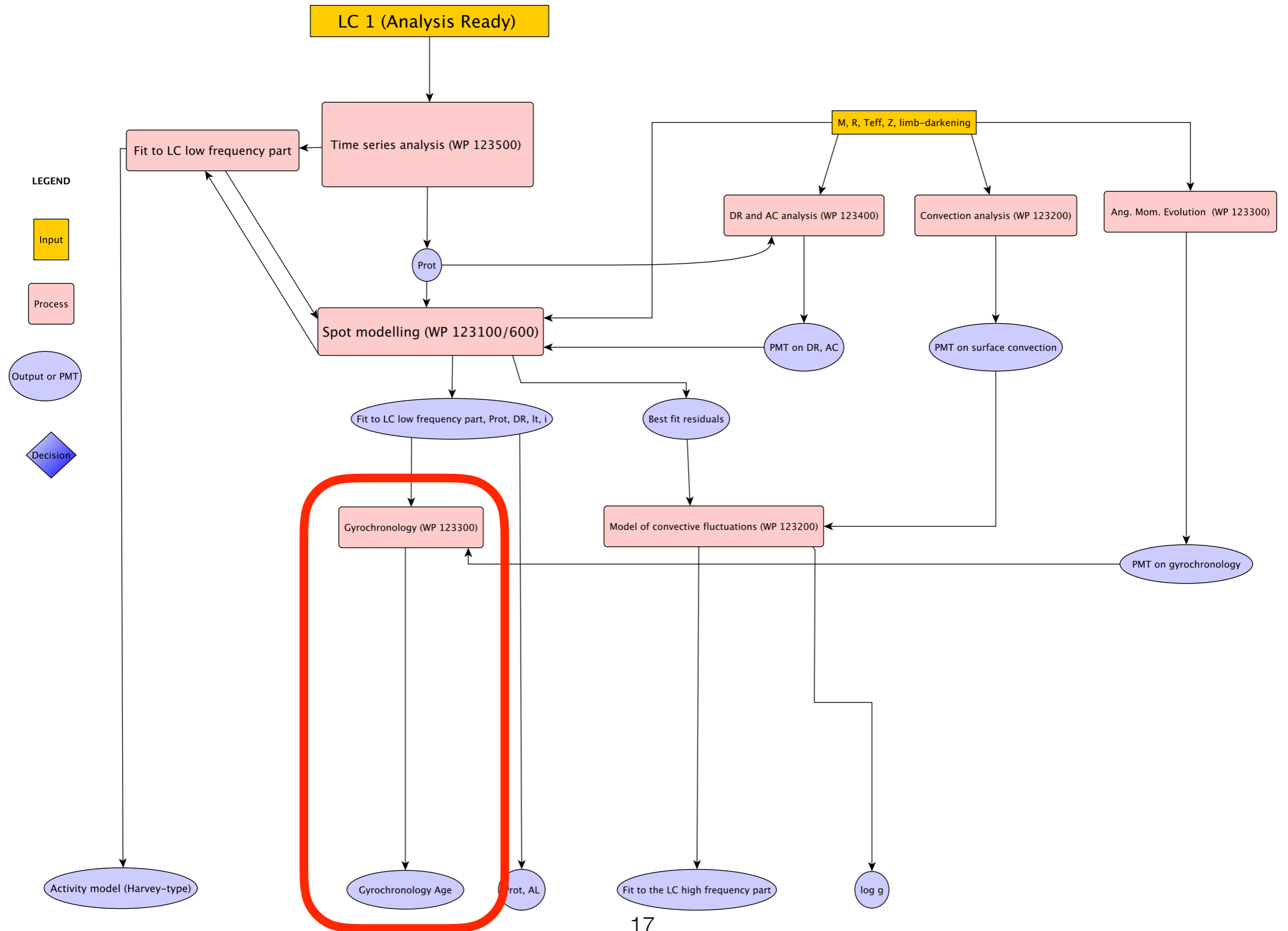
The pipeline



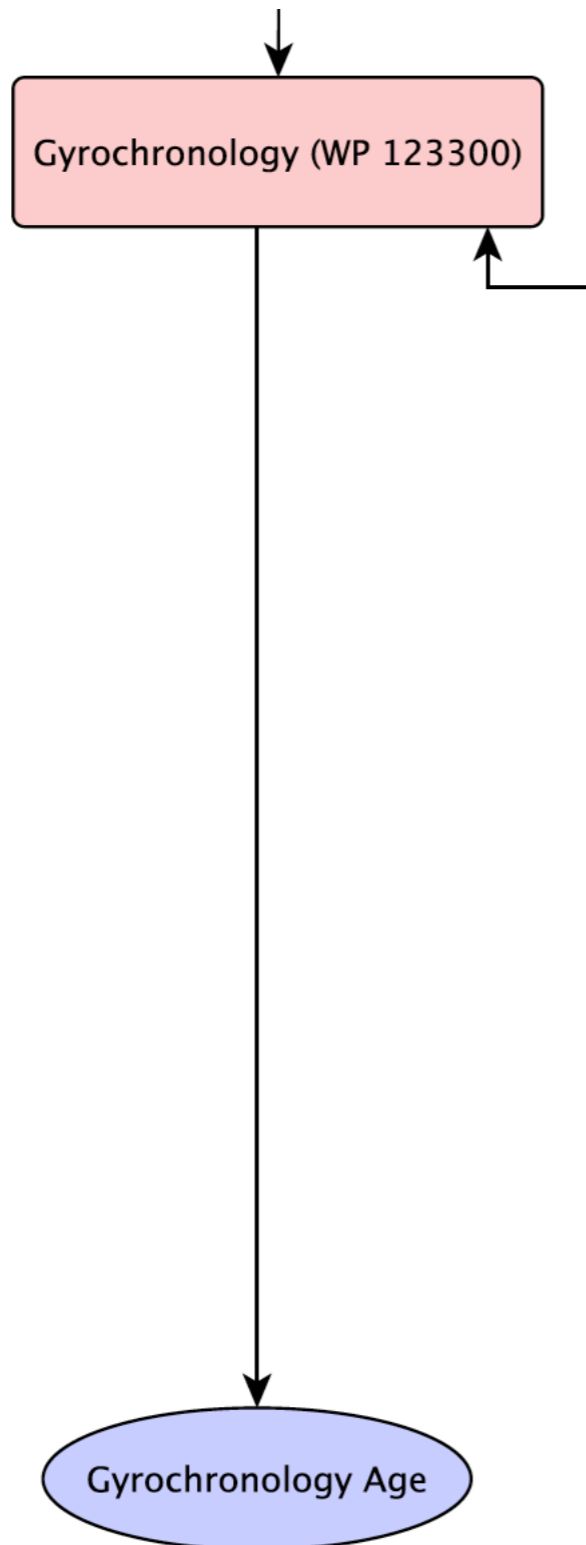
WP123100/600
Very tricky
No existing reliable algorithm
See Nadège's talk

=> SW
(Serious Worries)

The pipeline



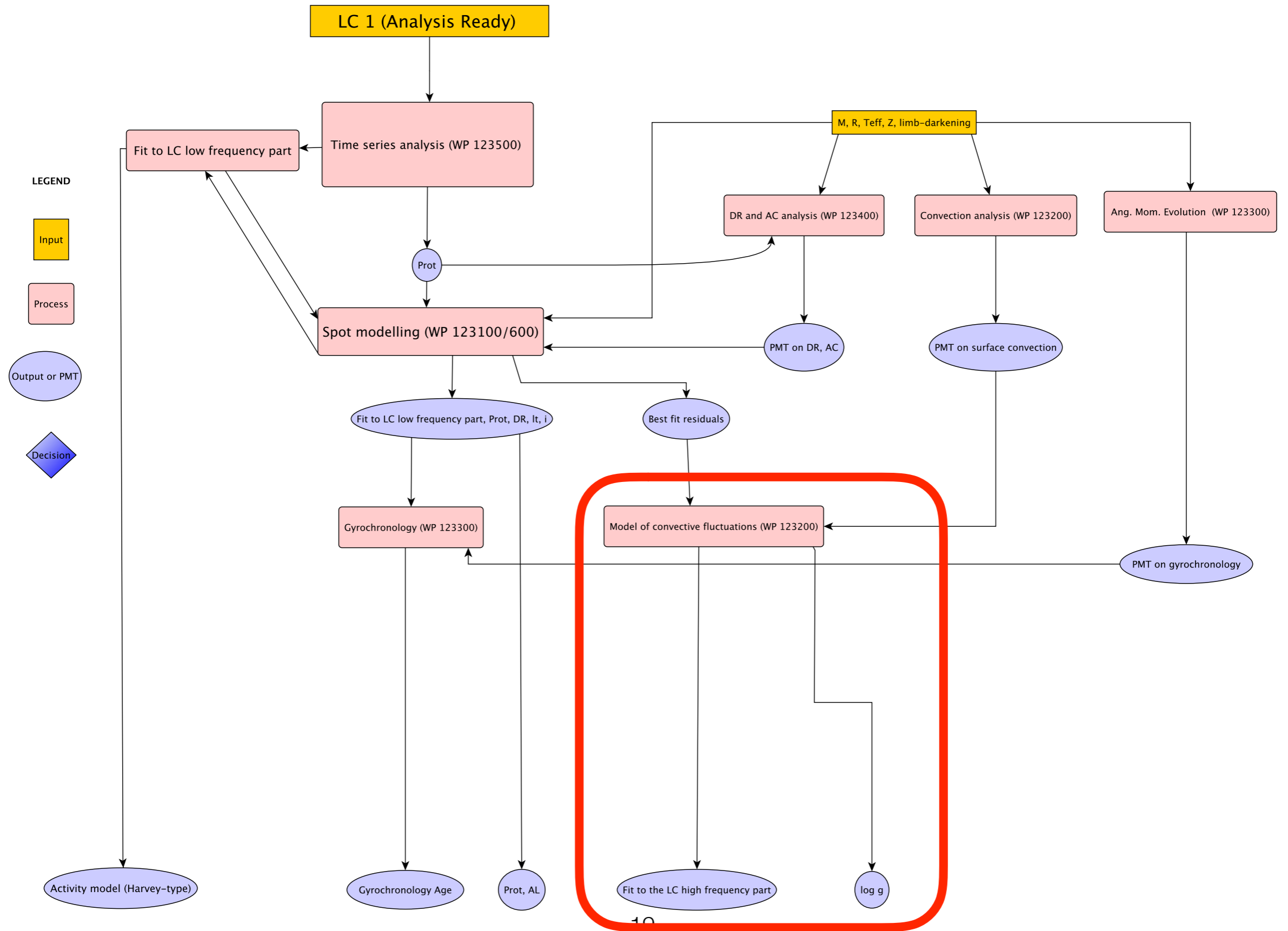
The pipeline



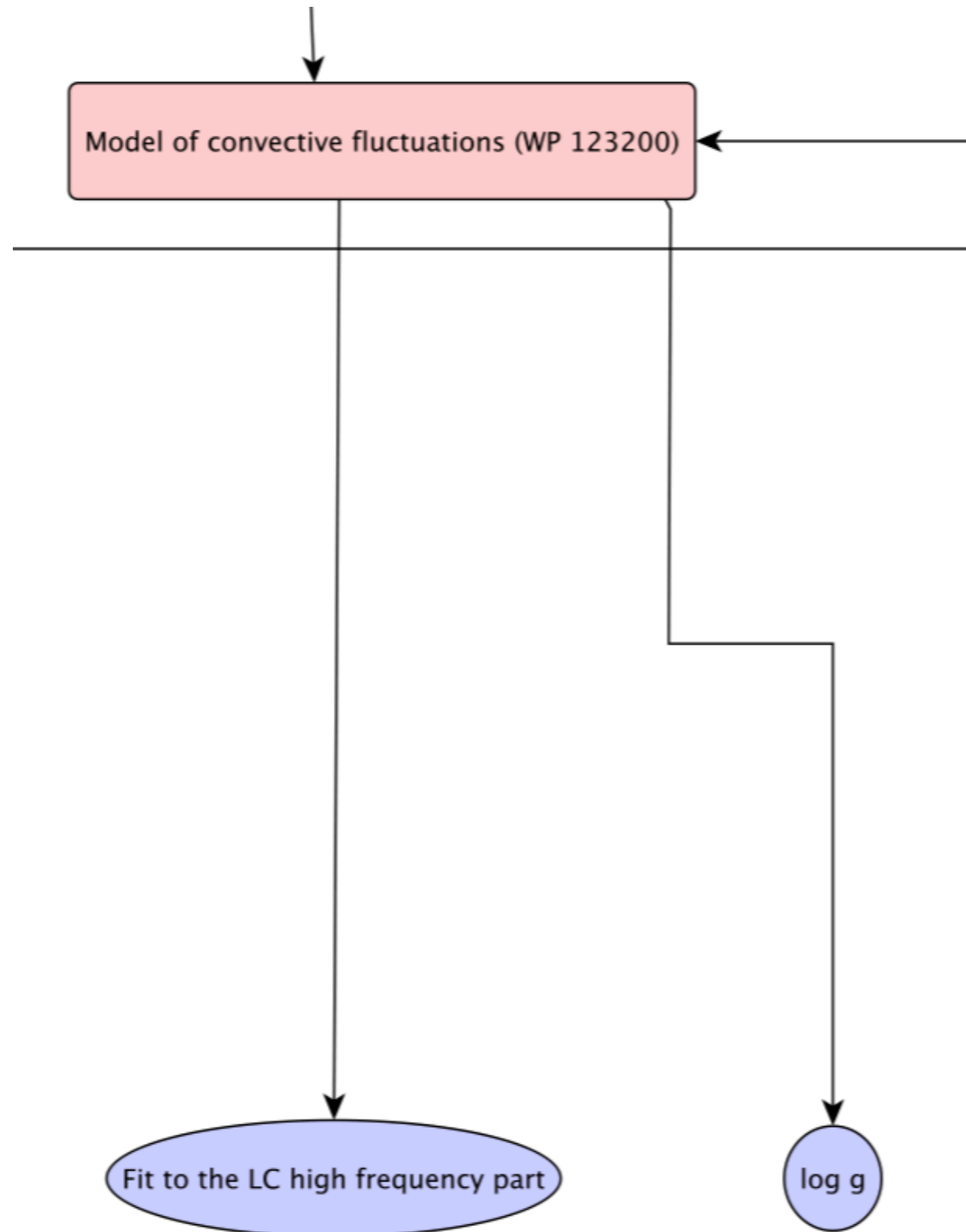
WP123300
Existing laws
To be compared with DP5

=> NW
(No Worries)

The pipeline



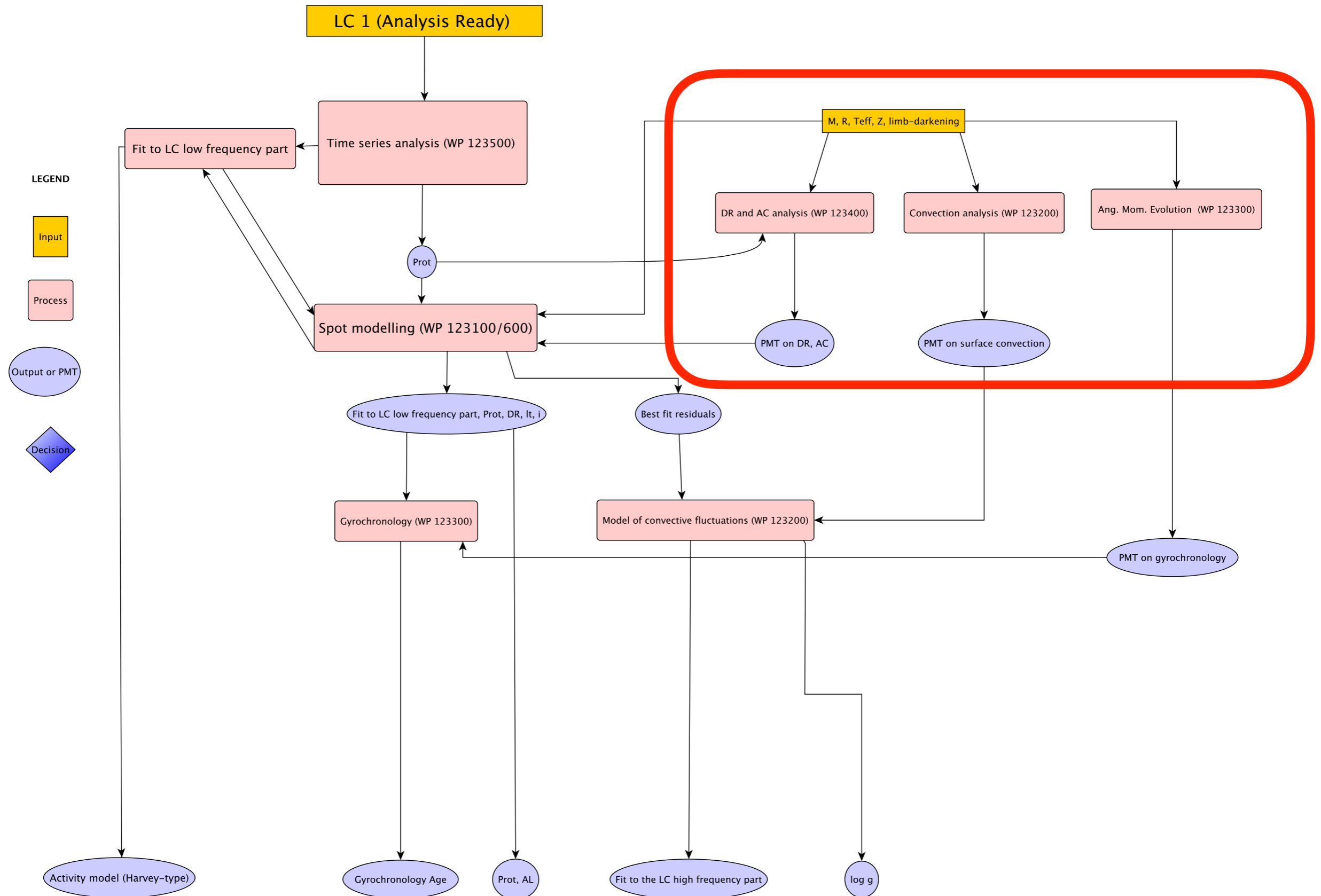
The pipeline



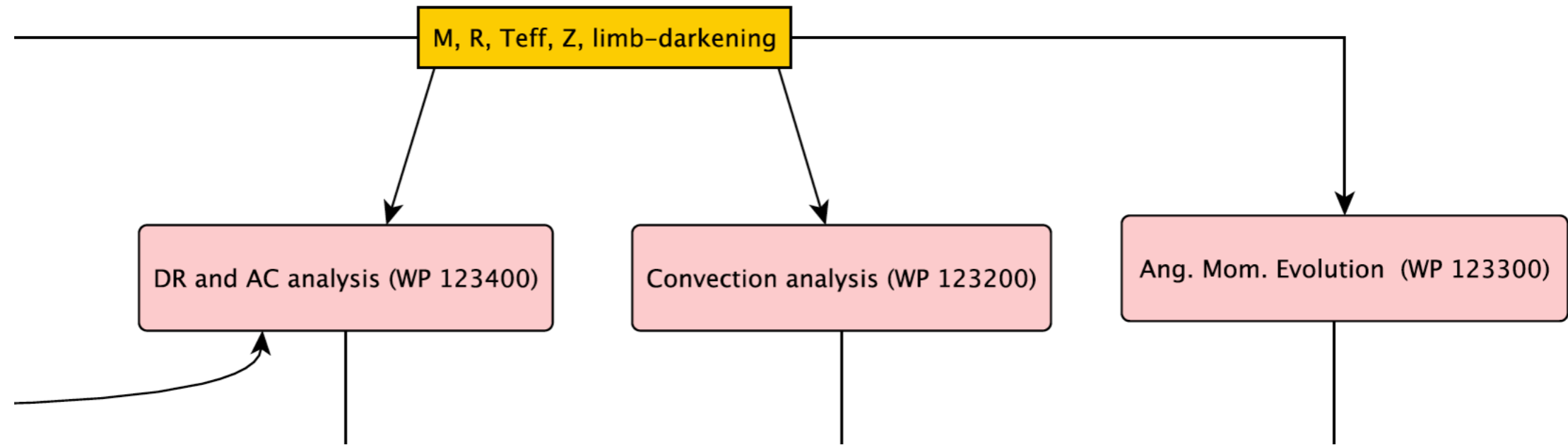
WP123200
Existing models and simulations
log(g) to be compared with DP5

=> NW
(No Worries)

The pipeline



The pipeline



WP123200/300/400

**Provide theoretical inputs/guess parameters
for other algorithms (differential rotation, cycles)
See Antoine Strugarek's presentation (WP123400)**

**=> NW
(No Worries)**

Interactions

- Inclination angle: spot modeling \Leftrightarrow mode fitting
 - Age: gyrochronology \Leftrightarrow stellar modeling
 - $\log(g)$: convection model \Leftrightarrow stellar modeling
 - Background fitting \Leftrightarrow spectrum fitting
- + interactions with EAS: to be explored

Conclusion

- Some tasks/DP are tricky (spot modeling)
- Some other look OK
- Interactions to be precised