

Lessons learned from Gaia data processing & analysis

Jordi Portell (IEEC/ICCUB)

LISA Spain Meeting
ICE-CSIC-IEEC, 15-October-2024

Finançat per



GOBIERNO
DE ESPAÑA



Next Generation
Catalunya



Generalitat
de Catalunya



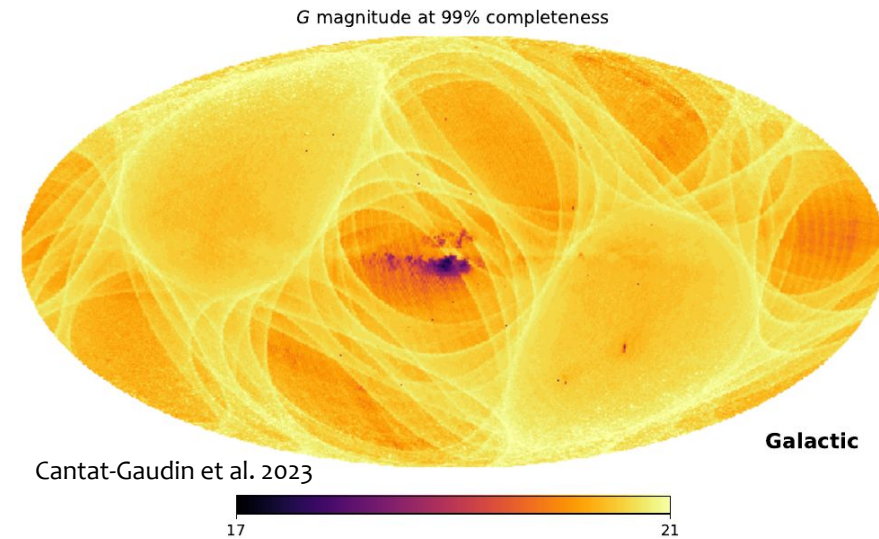
Selection function

- **Discovery machine**

- Observing any source within the *Gaia Selection Function*:
Scan law, brightness, source extension, colour, multiplicity, onboard events, downlink or ground segment issues, ...

- **LISA will also be a discovery machine!**

- *Listening to any source within the LISA Selection Function*
→ good to **start defining it**: envisaged limitations and capabilities: frequency band, distance to source, duration of event, sampling, kind (waveform) of event, noise features, orientation/localization, duty cycle, ...
- **Prepare for possible failures**, malfunctions or under-performance of some onboard systems



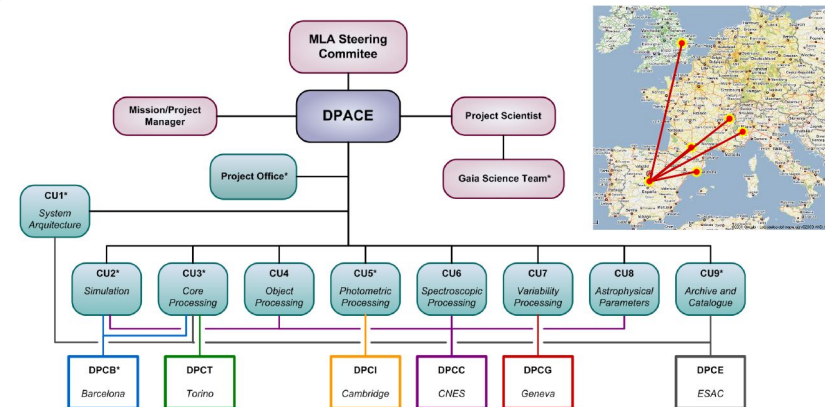
Tons of data

- Quite large **data volume**, yes:
 - 136 TB raw data from the spacecraft
 - About 1 PB of reduced data in the Main DataBase (MDB)
- But in Gaia, the main challenge is the **number of records**
 - 258 billion transits (star measurements)
 - **2.5 trillion** astrometric measurements
 - 50 billion high-resolution spectra
 - About 2 billion sources
 - (We don't really deal with *images* in the “classical” sense)
- Furthermore:
 - Strongly interrelated data
 - Complex calibration models
 - Iterative (and cyclic) data reduction approach
- Maybe not really comparable to LISA:
 - Not “autonomous observations/measurements” onboard
 - LISA will be more “pure number-crunching” than Gaia → **estimate resources** (Gaia computing is quite dominated by input/output)

CURRENT DATE AND TIME	2024-10-14T20:39:05 (TCB)
MISSION STATUS	
Satellite distance from Earth (in km)	1,595,107
Number of days having passed since 25 July 2014	3734
Number of days in mission extension	1917
OPERATIONS DATA (collected since 2014/07/25)	
Volume of science data collected (in GB)	136,287
Number of object transits through the focal plane	258,876,153,148
Number of astrometric CCD measurements	2,551,779,223,881
Number of photometric CCD measurements	512,748,606,080
Number of spectroscopic CCD measurements	50,123,773,560
Number of object transits through the RVS instrument	16,840,769,964

Complex data reduction challenge

- Data Processing and Analysis Consortium, organized in
 - Coordination+Development Units (CUs/DUs) → **algorithms & software**
 - Data Processing Centers (DPCs) → **hardware & operators**
- Each CU, DU and DPC with a (quite) clear goal
 - Not much overlap between them, but strong inter-dependencies
→ **strict schedule required**
(avoid propagation of delays downstream)
 - Sometimes “politics” may affect in the CU/DU/DPC definition
→ try to **harmonize politics with science+technology capabilities/expertise**



- **Embrace the Bubble**
 - Some core systems may need to “live in a bubble”:
Nearly-ideal conditions, data, models, etc.
(the conceptual/scientific/technical challenge may be extremely demanding)
 - Some/many systems basically exist to “create this bubble”
(minimize instrumental effects, issues in data, ...)

Technical considerations (I)

- **Define/recommend the main programming language**
 - Good for reusing “general” tools between units (algorithms, data access...)
 - Otherwise: code duplication in different languages, wrappers (overhead), ...
- **Test-oriented development**
 - Define manageable **Validation DataSets** for continuous regression testing
 - Ensure determinism: beware of multithreading/parallelism race conditions
- **Cyclic development**
 - **Define essential and realistic features, implement first version ASAP**
 - Improve with “not-so-essential” ones progressively
- **Implement realistic simulators ASAP**
 - In Gaia we started it ~14 years before launch
 - Universe + Instrument models; data simulation at different levels
- **Define an adequate file format**
 - If possible, fulfilling (1) data processing in a center, (2) data exchange between centers, and even (3) bulk data publication
 - Nowadays, Parquet looks great for this
 - If you really have to use more formats, make sure that bidirectional conversion can easily and reliably be done

Technical considerations (II)

- Choose the right computing framework/approach, incl. I/O approach
 - E.g. in Gaia, huge number of records
 - a *database* may not be the right approach for number crunching
 - ⇒ processing based on *files* (with some supporting DB for metadata)
 - But if you need a DB, choose very carefully the right approach for you. Nowadays, *columnar* DBs perform great - even for public/massive archives
- Ensure consistency in data, interfaces and software
 - Traceability + reproducibility
 - Tag and track versions of solutions, data model, software...
 - Consider adding a “**Solution Identifier**” to each and every table/file
- Define and clearly inform/describe **central repositories** for:
 - Software
 - Reference/test data
 - Working + released documents. BTW, **LaTeX** is *really* great.
 - If all these are organized following the Units and Centers, much better
 - Have a **central point** (wiki? Confluence?) with a **clear list and basic description** of all these services

Behold, the Holy Grail

- This is our 8th Wonder. Please, use or implement some tool to clearly **define and document your data model** in a centralized manner:

- Systems, tables, fields, units, multiplicity, descriptions...
- Transfers: consumers, periodicity
- Size estimation
- Reports
- Automatic code
- Sync with SVN/Git

- Seriously, this is one of the best things done in DPAC

The screenshot displays the Gaia Dictionary Tool interface. The main window shows a table of data model elements with columns for Name, Description, Det. Desc., Type, Multiplicity, Units, Minimum, Maximum, Default value, and Consumers. A 'Table Description' dialog is open, showing a detailed description of the 'Along Scan Coordinate of transmitted SM and AF windows wrt ref Detailed Description' table. The description explains that the table contains AL coordinates of transmitted SM windows and AF windows, used for double star work and 2D imaging. It also includes a warning about absolute AL coordinates and a code snippet for calculating corrected coordinates.

#	Name	Description	Det. Desc.	Type	Multiplicity	Units	Minimum	Maximum	Default value	Consumers
1	soluionId	Solution identifier	View	long						
2	fileId	PDHU file identifier, associated to a priority	View	short			0	255		
3	ccdRow	CCD Row	View	byte			1	7		
4	directTime	Residual (high-resolution) reference time	View	short		Time(OBMT50)	0	4095		
5	objectType	Object type	View	byte			0	6		
6	gClass	Window class used in SM, AF and XP	View	byte			0	3		IOAC(Daily/CU5)
7	gPriority	Priority of the observation used by the VPU	View	short			0	255		
8	gMag	G magnitude estimated by the VPU	View	short		Magnitude(mag)	0	1023		
9	gAlignCentring	Centring information for AF1-9	View	byte	[5]					
10	shapeAf	Shape (truncation) information for AF2-9 windows	View	byte	[18]					
11	aocsUpdate	Information on AOC5 updates occurred in SM + AF1-9	View	byte	[10]					
12	transitAccFlags	Transit-level information on the on-board acquisition	View	short						
13	ccdaqcFlags	CCD-level information on the on-board acquisition	View	byte	[10]					
14	gates	Gate information for all CCDs	View	byte	[10]					
15	aWinCoord	Along Scan Coordinate of transmitted SM and AF windows wrt ref	View	short	[10]	Time(TDI Period)				
16	aWinCoord	AC coordinate of the transmitted SM and AF windows	View	byte	[10]	Length & Distance(pixel)				
17	ismSamples	Samples from the sky mapper	View	int	[var:62]	Flux(ADU)	0	65535		IOAC(Daily/CU5)
18	afSamples	Samples from the Astrometric Field (AFs)	View	int	[106]	Flux(ADU)	0	65535		IOAC(Daily/CU5)
19	distToLastCI	Distance to last Charge injection	View	short	[10]	Length & Distance(pixel)				

The 'Table Description' dialog for 'Along Scan Coordinate of transmitted SM and AF windows wrt ref Detailed Description' contains the following text:

The astro observation records contain the raw data from the SM and AF transits. They are used by IDT itself, IDU, AVU, for double star work in CU4, and for 2D imaging and PSF calibration in CU5. This information is retrieved from SPI, ASD1, ASD5, ASD6 and the CDB (mainly AL Phasing table).

Please see CF-010 for more information

Warning: These are the absolute AL coordinates of the windows w.r.t. the reference time, that is, they already include the AL Phasing Table correction.

5 However the following correction must be done to avoid exceeding the range of a 'short':

```

7 aWinCoord_coded[i] = aWinCoord[i] - (1 <<< 14) \
8 ...that is, 16384 must be subtracted.

```

10 Thus, with the example of AL phasing table provided in GAIA.ASF.TCN.PLM.00137 (Table 6-6, p.84), it should lead to an approximate range of -27260 TDIs to +23137 TDIs.

Technical considerations (IV)

- Evaluate properly the **long-term support** (and license approach) of your language and tools
 - E.g.: JVM/JDK version (and vendor: Oracle, ehem...), Python version
 - Apache Commons, Numpy, Astropy, Pandas, ...
- Also for **project/software management tools**, e.g.:
 - Code repository: SVN, Git
 - Issue tracking: We started with Mantis, migrated to JIRA later
 - Avoid, as much as possible, having to migrate during (or close to) operations
- Clearly define, *beforehand* or ASAP, a **software licensing approach** for all partners involved (ESA, institutes, universities, contractors...)
 - Much, MUCH difficult if done later
 - We're still struggling with it
- Define **project services**, shared calendars, mailing lists...
 - Create document **templates** for the various types (did I mention LaTeX already?)
 - Clearly define a “**document codes**” approach, and create (and update) a list of authors (+initials) and institutions
 - Create a list of **acronyms** ASAP and regularly maintain it
 - Create a **Parameters Database**. Now.

The Human dimension

- Prepare “**welcome packs**” and **training resources**
 - General project info including project services, repositories, essential code tools (e.g. basic guidelines to setup the code environment), document guidelines...
 - Coordinator/manager lists
 - What are you supposed to do?
Who should you report to?
Are you responsible (or will you be) for somebody else?
Who will you work with (locally or remotely)?
Does your work overlap with someone else’s?
Who can you ask when you get blocked with your subject?
- Don’t postpone too much your tasks
 - *One JIRA a day keeps QA away*
- Accept the truth: **Some key people will leave the project at the worst time**
 - Be ready to replace him/her
 - Add redundancy before it’s too late: **distribute the knowledge**
 - **Avoid single-person projects**: train and delegate tasks
- (Astro)physicists are not (software)engineers, and vice versa
 - You will probably need “translators” in between. **Please have patience!**
 - Accept that some excellent scientists may not program well → support them

In a nutshell

You're already late

In a nutshell

You're already late

Don't desperate, prioritize

In a nutshell

You're already late

Don't desperate, prioritize

Have as much fun as possible

In a nutshell

You're already late

Don't desperate, prioritize

Have as much fun as possible
(so you'll start from the lowest-priority task)

Thank you

Finançat per



GOBIERNO
DE ESPAÑA



Next Generation
Catalunya



Generalitat
de Catalunya

