

Institut d'Estudis Espacials de Catalunya

Estimate of the magnetic contribution to acceleration noise in LISA Pathfinder

Daniel Serrano Rubio LISA Spain 16-10-2024 dserrano@ieec.cat

Generalitat de Catalunya

Arxiv:

M Armano et al. "Precision measurements of the magnetic parameters of LISA Pathfinder test masses". 2024, arXiv:2407.04431.

M Armano et al. "Magnetic-induced force noise in LISA Pathfinder free-falling test masses". 2024, arXiv:2407.04427

UAR Universitat Autònoma de Barcelona

UNIVERSITAT POLITÈCNICA DE CATALUNYA

Top image credit: https://spacenews.com/lisa-pathfinders-success-boosts-likelihood-of-future-gravity-waveobservatory/ Bottom image credit: LISA Study Definition Report - Red Book (2024)

LISA Pathfinder

- 2 TMs in free-fall
- Mission from 2015 to 2017
- Beyond LISA requirements
- **Δg: residual acceleration** between TMs along axis joining them (x axis)

IEECR

LISA Pathfinder DDS

- Data and Diagnostics Subsystem
	- Temperature subsystem
	- **Magnetic subsystem**
	- Radiation monitor
- Magnetic Diagnostic Subsystem
	- 4 tri-axial fluxgate magnetometers
	- 2 induction coils

IEECR

Magnetic Diagnostic Subsystem

Bulky

Power consuming

• 4 tri-axial fluxgate magnetometers • **•** 6 bi-axial AMR magnetometers

Coll-1

IEECI

2400 turns of copper wire 5.65 cm radius

Same coils as for LPF Measuring signals in the 50-500 Hz range

Small remanence More compact

• **2 injection coils** • 2 audio-band measuring coils

- Remanent magnetic moment (Mr)
- Magnetic susceptibility (χ)
- Background magnetic field and gradient at TM location
- Homogeneity and stationarity of magnetic properties

5

- TM behaves like a magnetic dipole: $\vec{F} = \big(\overrightarrow{m} \cdot \overrightarrow{\nabla}\big)\overrightarrow{B}$ $\vec{N} = \overrightarrow{m} \times \vec{B} + \vec{r} \times (\overrightarrow{m} \cdot \nabla) \vec{B}$ $\vec{m} = \vec{m_r} +$ χ μ_{0} \overline{B}
- Magnetic forces and fields dominate the background and other sources during injections

$$
\vec{F} = \left\langle (\overrightarrow{m_r} \cdot \vec{\nabla}) \vec{B} + \frac{\chi}{\mu_0} (\vec{B} \cdot \vec{\nabla}) \vec{B} \right\rangle V \qquad \vec{N} = \left\langle \overrightarrow{m_r} \times \vec{B} + \vec{r} \times \left[(\overrightarrow{m_r} \cdot \vec{\nabla}) \vec{B} + \frac{\chi}{\mu_0} (\vec{B} \cdot \vec{\nabla}) \vec{B} \right] \right\rangle V
$$

$$
\vec{B} = \vec{B}_0 + \overrightarrow{B^{AC}} \sin(\omega t) \qquad \text{where} \qquad \vec{B}_0 = \vec{B}_{back.} + \overrightarrow{B^{DC}}
$$

$$
\vec{F} = \overrightarrow{F_{DC}} + \overrightarrow{F_{1\omega}} + \overrightarrow{F_{2\omega}} \qquad \qquad \vec{N} = \overrightarrow{N_{DC}} + \overrightarrow{N_{1\omega}} + \overrightarrow{N_{2\omega}}
$$

Terms of interest:

$$
\overrightarrow{F_{DC}} = \left\langle (\overrightarrow{M_r} \cdot \overrightarrow{\nabla}) \overrightarrow{B_0} \right\rangle + \frac{\chi V}{\mu_0} \left[\left\langle (\overrightarrow{B_0} \cdot \overrightarrow{\nabla}) \overrightarrow{B_0} \right\rangle + \frac{1}{2} \left\langle (\overrightarrow{B^{AC}} \cdot \overrightarrow{\nabla}) \overrightarrow{B^{AC}} \right\rangle \right]
$$
\n
$$
\overrightarrow{F_{1\omega}} = \left\{ \left\langle (\overrightarrow{M_r} \cdot \overrightarrow{\nabla}) \overrightarrow{B^{AC}} \right\rangle + \frac{\chi V}{\mu_0} \left[\left\langle (\overrightarrow{B_0} \cdot \overrightarrow{\nabla}) \overrightarrow{B^{AC}} \right\rangle + \left\langle (\overrightarrow{B^{AC}} \cdot \overrightarrow{\nabla}) \overrightarrow{B_0} \right\rangle \right] \right\} \sin(\omega t)
$$

$$
\overrightarrow{F_{2\omega}} = \left\{-\frac{\chi V}{2\mu_0} \left\langle \left(\overrightarrow{B^{AC}} \cdot \overrightarrow{\nabla}\right) \overrightarrow{B^{AC}} \right\rangle \right\} \cos(2\omega t)
$$

$$
\overrightarrow{N_{1\omega}} = \left\langle \overrightarrow{M_r} \times \overrightarrow{B^{AC}} \right\rangle \sin(\omega t)
$$

- Injected magnetic fields $B^{\bar D C}$ and $B^{A \bar C}$ and their gradients have to be averaged over the TM volume as defined previously by …
- Off-axis magnetic field of a coil involves elliptic integrals

$$
B_{\rho}(x,\rho) = A_{\rho} \frac{x}{\rho^{\frac{3}{2}}} F(k) \qquad B_{x}(x,\rho) = A_{x} \rho^{-\frac{3}{2}} G(k) - \frac{\rho}{x} B_{\rho}(x,\rho)
$$

$$
A_{\rho} = \frac{\mu_0}{4\pi} \frac{NI}{a^{1/2}} \qquad A_x = \frac{a}{2} A_{\rho} \qquad F(k) = k \left[\frac{1 - k^2/2}{1 - k^2} E(k) - K(k) \right] \quad G(k) = \frac{k^3}{1 - k^2} E(k)
$$

When averaged over the TM volume one finds out that thanks to the symmetry of the system :

- Bx is 10 orders of magnitude larger than y and z components
- Equations for all gradients can also be calculated, such that $\partial x B x$ is 4 orders of magnitude larger than ∂y , zBx
- Furthermore, a relationship between Bx and $\partial x Bx$ such that $Bx = \alpha * \partial x Bx$ can be found and it is only dependent on the geometry of the system
- The torque is found not to have a 2ω component

- Equations simplified thanks to symmetry of the system
- Forces obtained by demodulating the ∆g signal at the frequencies of interest and multiplying by either the TM mass for the force or by the moment of inertia of a cube for the torque

$$
N_{1\omega,\phi} = -M_{\mathcal{Y}} \langle B_x^{AC} \rangle
$$

$$
N_{1\omega,\eta}=M_{z}\langle B_{x}^{AC}\rangle
$$

$$
F_{2\omega,x} = -\frac{\chi V}{2\mu_0} \langle B_x^{AC} \rangle \langle \nabla_x B_x^{AC} \rangle
$$

$$
F_{1\omega,x} = M_{eff,x} \langle \nabla_x B_x^{AC} \rangle \quad \text{where} \quad M_{eff,x} \equiv \left[M_x + 2 \frac{\chi V}{\mu_0} \langle B_x^{DC} \rangle \right]
$$

$$
F_{DC,x}
$$
\n
$$
= \left(\frac{\chi V}{\alpha \mu_0}\right) \langle B_x^{DC} \rangle^2 + \left[\frac{M_x}{\alpha} + \frac{\chi V}{\mu_0} \left(\nabla_x B_{back,x} + \frac{B_{back,x}}{\alpha}\right)\right] \langle B_x^{DC} \rangle
$$
\n
$$
+ \left\{ (M_x + M_y + M_z) \nabla_x B_{back,x} + \frac{\chi V}{\mu_0} \left[3B_{back,x} \nabla_x B_{back,x} + \frac{1}{2} \langle B_x^{AC} \rangle \langle \nabla_x B_x^{AC} \rangle \right] \right\}
$$

- Experimental measurements during injections
- 11% correction to magnetic fields calculated before at the TM location due to magnetometers calibration discrepancy
- Cause: Tolerances during manufacturing and tilts/misalignments during mounting and launch

Δg rotations demodulated at the **injected frequency** allow determination of M_v and M_z

 $M_v = (0.178 \pm 0.025) \text{ n} A m^2$ $M_z = (0.095 \pm 0.010)$ nAm²

Δg acceleration demodulated at **twice the injected frequency** provides results of the magnetic susceptibility

 $\chi_{2mHz} = (-3.43 \pm 0.58) * 10^{-5}$ $\chi_{6mHz}=$ (-2.65 \pm 0.62) * 10⁻⁵ $\chi_{10 mHz} = (-3.35 \pm 0.12)*10^{-5}$ $\chi_{30mHz}=(-4.73\pm0.34)*10^{-5}$ Δg acceleration demodulated at **the injected frequency** against the gradient of the AC field have $M_{eff,x}$ as the slope of the fit. $M_{eff,x}$ linear dependence with the DC field determines M_r as the offset and χ as the slope

$$
M_x = (0.140 \pm 0.138) \, nAm^2 \qquad \qquad \chi_{5mHz} = (-3.3723 \pm 0.0069) \cdot 10^{-5}
$$

www.ieec.cat dserrano@ieec.cat

 -1.236 -1.237

 -1.238

 $\sqrt{2}$ -1.239 $\frac{1}{6}$ -1.240

> -1.241 -1.242

 -1.243

A quadratic fit of the dependence of F_{DC} with the injected DC field allows the determination of χ_{DC} , $B_{back,x}$ and $\nabla_x B_{back,x}$.

 $(225 + 0.15) \cdot 10^{-5}$ The latter two are of great importance as this is the **only way** to determine the precise values of both within the TM location (magnetometers are located too far away)

IEECR

$$
\chi_{DC} = (-3.35 \pm 0.15) * 10^{-5}
$$

$$
B_{back,x} = (414 \pm 74) nT
$$

$$
\nabla_x B_{back,x} = (-7400 \pm 2100) nT/m
$$

$$
F_{DC,x}
$$
\n
$$
= \left(\frac{\chi V}{\alpha \mu_0}\right) (B_x^{DC})^2 + \left[\frac{M_x}{\alpha} + \frac{\chi V}{\mu_0} \left(\nabla_x B_{back,x} + \frac{B_{back,x}}{\alpha}\right)\right] (B_x^{DC})
$$
\n
$$
+ \left\{ (M_x + M_y + M_z) \nabla_x B_{back,x} + \frac{\chi V}{\mu_0} \left[3B_{back,x} \nabla_x B_{back,x} + \frac{1}{2} \langle B_x^{AC} \rangle \langle \nabla_x B_x^{AC} \rangle \right] \right\}
$$

TMs magnetic parameters extraction residual

12

dserrano@ieec.cat

Magnetic contribution to acceleration noise

• From the general formula of the force of a dipole we can derive the amplitude spectrum in acceleration

 $S_{\Delta g}^{1/2} = \frac{\chi V}{M_{TM}\mu_0}$ $\overrightarrow{\nabla}B_{\chi}\big\}S_{\overrightarrow{B}}^{1/2}$

- We have only obtained $\nabla_x B_x =$ (-7400 ± 2100) nT/m. This value can be attributed to NTCs thermistors at the EH
- Rest of gradients were found by a Monte -Carlo simulation of the NTCs surrounding the TM

dserrano@ieec.cat

Magnetic contribution to acceleration noise

- Fluctuations of magnetic field $(S$ \overline{B} $\frac{1}{2}$) originated by interplanetary magnetic field
- Amplitude Spectrum Density (ASD) during February noise run, 2017

Contribution at 0.1 mHz: 1.46 $^{+3.73}_{-0.77}$ % **(in noise power)**

Magnetic contribution to acceleration noise

• Magnetic fluctuations show non-stationarities related to solar wind speed variations

M. Armano et al. Spacecraft and interplanetary contributions to the magnetic environment on-board LISA Pathfinder. Monthly Notices of the Royal Astronomical Society, 494(2):3014–3027, 04 2020.

IEECR

Conclusions

TMs magnetic parameters

- $|\vec{M}| = (0.245 \pm 0.081)$ nAm² < 10 nAm²
- $B_{back.x} = (414 \pm 74) \text{ nT}$
- $\nabla_x B_{back.x} = (-7400 \pm 2100) \text{ nT/m}$
- $\chi = (-3.3723 \pm 0.0069) * 10^{-5}$ at 5 mHz

Magnetic induced acceleration noise contribution to Δg

- At 1 mHz: $0.25^{+0.15}_{-0.08}$ $fms^{-2}Hz^{-1/2} < 12fms^{-2}Hz^{-1/2}$
- **Non-stationarities** increase contribution by a factor of 4.6

Thanks for you attention!

