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Estimate of the magnetic contribution to acceleration noise in LISA Pathfinder

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Arxiv:



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masses". 2024, arXiv:2407.04427

Pathfinder test masses". 2024, arXiv:2407.04431.



M Armano et al. "Precision measurements of the magnetic parameters of LISA

M Armano et al. "Magnetic-induced force noise in LISA Pathfinder free-falling test



Top image credit: https://spacenews.com/lisa-pathfinders-success-boosts-likelihood-of-future-gravity-waveobservatory/ Bottom image credit: LISA Study Definition Report - Red Book (2024)

LISA Pathfinder

- 2 TMs in free-fall
- Mission from 2015 to 2017
- Beyond LISA requirements
- Δg: residual acceleration between TMs along axis joining them (x axis)



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LISA Pathfinder DDS

- Data and Diagnostics Subsystem
 - Temperature subsystem
 - Magnetic subsystem
 - Radiation monitor
- Magnetic Diagnostic Subsystem
 - 4 tri-axial fluxgate magnetometers
 - 2 induction coils



Magnetic Diagnostic Subsystem

Bulky

DDS

• 4 tri-axial fluxgate magnetometers



• 2 injection coils

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Power consuming



Small remanence More compact

SDS

2 audio-band measuring coils

6 bi-axial AMR magnetometers



2400 turns of copper wire 5.65 cm radius



Same coils as for LPF Measuring signals in the 50-500 Hz range

- Remanent magnetic moment (Mr)
- Magnetic susceptibility (χ)
- Background magnetic field and gradient at TM location
- Homogeneity and stationarity of magnetic properties



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- TM behaves like a magnetic dipole: $\vec{F} = (\vec{m} \cdot \vec{\nabla})\vec{B}$ $\vec{N} = \vec{m} \times \vec{B} + \vec{r} \times (\vec{m} \cdot \vec{\nabla})\vec{B}$ $\vec{m} = \vec{m_r} + \frac{\chi}{\mu_0}\vec{B}$
- Magnetic forces and fields dominate the background and other sources during injections

$$\vec{F} = \left((\overrightarrow{m_r} \cdot \vec{\nabla}) \vec{B} + \frac{\chi}{\mu_0} (\vec{B} \cdot \vec{\nabla}) \vec{B} \right) V \qquad \vec{N} = \left(\overrightarrow{m_r} \times \vec{B} + \vec{r} \times \left[(\overrightarrow{m_r} \cdot \vec{\nabla}) \vec{B} + \frac{\chi}{\mu_0} (\vec{B} \cdot \vec{\nabla}) \vec{B} \right] \right) V$$
$$\vec{B} = \vec{B}_0 + \vec{B} \vec{AC} \sin(\omega t) \qquad where \qquad \vec{B}_0 = \vec{B}_{back.} + \vec{B} \vec{DC}$$
$$\vec{F} = \vec{F_{DC}} + \vec{F_{1\omega}} + \vec{F_{2\omega}} \qquad \qquad \vec{N} = \vec{N_{DC}} + \vec{N_{1\omega}} + \vec{N_{2\omega}}$$

Terms of interest:

$$\overline{F_{DC}} = \left\langle \left(\overrightarrow{M_r} \cdot \overrightarrow{\nabla} \right) \overrightarrow{B_0} \right\rangle + \frac{\chi V}{\mu_0} \left[\left\langle \left(\overrightarrow{B_0} \cdot \overrightarrow{\nabla} \right) \overrightarrow{B_0} \right\rangle + \frac{1}{2} \left\langle \left(\overrightarrow{B^{AC}} \cdot \overrightarrow{\nabla} \right) \overrightarrow{B^{AC}} \right\rangle \right] \right]$$
$$\overline{F_{1\omega}} = \left\{ \left\langle \left(\overrightarrow{M_r} \cdot \overrightarrow{\nabla} \right) \overrightarrow{B^{AC}} \right\rangle + \frac{\chi V}{\mu_0} \left[\left\langle \left(\overrightarrow{B_0} \cdot \overrightarrow{\nabla} \right) \overrightarrow{B^{AC}} \right\rangle + \left\langle \left(\overrightarrow{B^{AC}} \cdot \overrightarrow{\nabla} \right) \overrightarrow{B_0} \right\rangle \right] \right\} \sin(\omega t)$$

$$\overrightarrow{F_{2\omega}} = \left\{ -\frac{\chi V}{2\mu_0} \left\langle \left(\overrightarrow{B^{AC}} \cdot \overrightarrow{\nabla} \right) \overrightarrow{B^{AC}} \right\rangle \right\} \cos(2\omega t)$$

$$\overrightarrow{N_{1\omega}} = \left\langle \overrightarrow{M_r} \times \overrightarrow{B^{AC}} \right\rangle \sin(\omega t)$$

- Injected magnetic fields B^{DC} and B^{AC} and their gradients have to be averaged over the TM volume as defined previously by (...)
- Off-axis magnetic field of a coil involves elliptic integrals



$$B_{\rho}(x,\rho) = A_{\rho} \frac{x}{\rho^{\frac{3}{2}}} F(k) \qquad \qquad B_{x}(x,\rho) = A_{x} \rho^{-\frac{3}{2}} G(k) - \frac{\rho}{x} B_{\rho}(x,\rho)$$

$$A_{\rho} = \frac{\mu_0}{4\pi} \frac{NI}{a^{1/2}} \qquad A_x = \frac{a}{2} A_{\rho} \qquad F(k) = k \left[\frac{1 - k^2/2}{1 - k^2} E(k) - K(k) \right] \qquad G(k) = \frac{k^3}{1 - k^2} E(k)$$

When averaged over the TM volume one finds out that thanks to the symmetry of the system :

- Bx is 10 orders of magnitude larger than y and z components
- Equations for all gradients can also be calculated, such that ∂xBx is 4 orders of magnitude larger than $\partial y, zBx$
- Furthermore, a relationship between Bx and ∂xBx such that $Bx = \alpha * \partial xBx$ can be found and it is only dependent on the geometry of the system
- The torque is found not to have a 2ω component



- Equations simplified thanks to symmetry of the system
- Forces obtained by demodulating the ∆g signal at the frequencies of interest and multiplying by either the TM mass for the force or by the moment of inertia of a cube for the torque

$$N_{1\omega,\varphi} = -M_y \langle B_x^{AC} \rangle$$

$$N_{1\omega,\eta} = M_z \langle B_x^{AC} \rangle$$

$$F_{2\omega,x} = -\frac{\chi V}{2\mu_0} \langle B_x^{AC} \rangle \langle \nabla_x B_x^{AC} \rangle$$

$$F_{1\omega,x} = M_{eff,x} \langle \nabla_x B_x^{AC} \rangle \quad where \quad M_{eff,x} \equiv \left[M_x + 2 \frac{\chi V}{\mu_0} \langle B_x^{DC} \rangle \right]$$

$$\begin{split} F_{DC,x} &= \left(\frac{\chi V}{\alpha \mu_0}\right) \langle B_x^{DC} \rangle^2 + \left[\frac{M_x}{\alpha} + \frac{\chi V}{\mu_0} \left(\nabla_x B_{back,x} + \frac{B_{back,x}}{\alpha}\right)\right] \langle B_x^{DC} \rangle \\ &+ \left\{ \left(M_x + M_y + M_z\right) \nabla_x B_{back,x} + \frac{\chi V}{\mu_0} \left[3B_{back,x} \nabla_x B_{back,x} + \frac{1}{2} \langle B_x^{AC} \rangle \langle \nabla_x B_x^{AC} \rangle \right] \right\} \end{split}$$

- Experimental measurements during injections
- 11% correction to magnetic fields calculated before at the TM location due to magnetometers calibration discrepancy
- Cause: Tolerances during manufacturing and tilts/misalignments during mounting and launch



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 $N_{1\omega,\Phi} = -M_{\gamma} \langle B_{\chi}^{AC} \rangle$

 Δ g rotations demodulated at the **injected frequency** allow determination of M_y and M_z

 $M_y = (0.178 \pm 0.025) nAm^2$ $M_z = (0.095 \pm 0.010) nAm^2$



 $F_{1\omega,x} = M_{eff,x} \langle \nabla_x B_x^{AC} \rangle$ where $M_{eff,x} \equiv \left| M_x + 2 \frac{\chi V}{\mu_0} \langle B_x^{DC} \rangle \right|$



∆g acceleration demodulated at **twice the injected frequency** provides results of the magnetic susceptibility

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$$\begin{split} \chi_{2mHz} &= (-3.43 \pm 0.58) * 10^{-5} \\ \chi_{6mHz} &= (-2.65 \pm 0.62) * 10^{-5} \\ \chi_{10mHz} &= (-3.35 \pm 0.12) * 10^{-5} \\ \chi_{30mHz} &= (-4.73 \pm 0.34) * 10^{-5} \end{split}$$

 Δ g acceleration demodulated at **the injected frequency** against the gradient of the AC field have $M_{eff,x}$ as the slope of the fit. $M_{eff,x}$ linear dependence with the DC field determines M_x as the offset and χ as the slope

$$M_x = (0.140 \pm 0.138) nAm^2$$
 $\chi_{5mHz} = (-3.3723 \pm 0.0069) * 10^{-5}$



A quadratic fit of the dependence of $F_{DC,x}$ with the injected DC field allows the determination of χ_{DC} , $B_{back,x}$ and $\nabla_x B_{back,x}$.

The latter two are of great importance as this is the **only way** to determine the precise values of both within the TM location (magnetometers are located too far away) $y = (-2.25 \pm 0.15) \pm 10^{-5}$

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$$\chi_{DC} = (-3.35 \pm 0.15) * 10^{-5}$$
$$B_{back.,x} = (414 \pm 74) nT$$
$$\nabla_x B_{back.,x} = (-7400 \pm 2100) nT/m$$

$$F_{DC,x} = \left(\frac{\chi V}{\alpha \mu_0}\right) \langle B_x^{DC} \rangle^2 + \left[\frac{M_x}{\alpha} + \frac{\chi V}{\mu_0} \left(\nabla_x B_{back,x} + \frac{B_{back,x}}{\alpha}\right)\right] \langle B_x^{DC} \rangle \\ + \left\{ \left(M_x + M_y + M_z\right) \nabla_x B_{back,x} + \frac{\chi V}{\mu_0} \left[3B_{back,x} \nabla_x B_{back,x} + \frac{1}{2} \langle B_x^{AC} \rangle \langle \nabla_x B_x^{AC} \rangle\right] \right\}$$



TMs magnetic parameters extraction residual



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Magnetic contribution to acceleration noise

 From the general formula of the force of a dipole we can derive the amplitude spectrum in acceleration

$$S_{\Delta g}^{1/2} = \frac{\chi V}{M_{TM}\mu_0} \langle \vec{\nabla} B_{\chi} \rangle S_{\vec{B}}^{1/2}$$

- We have only obtained $\nabla_x B_x = (-7400 \pm 2100)$ nT/m. This value can be attributed to NTCs thermistors at the EH
- Rest of gradients were found by a Monte-Carlo simulation of the NTCs surrounding the TM



Magnetic contribution to acceleration noise

- Fluctuations of magnetic field $(S_{\vec{B}}^{1/2})$ originated by interplanetary magnetic field
- Amplitude Spectrum Density (ASD) during February noise run, 2017



Contribution at 0.1 mHz: $1.46^{+3.73}_{-0.77}$ % (in noise power)



Magnetic contribution to acceleration noise

 Magnetic fluctuations show non-stationarities related to solar wind speed variations



M. Armano et al. Spacecraft and interplanetary contributions to the magnetic environment on-board LISA Pathfinder. Monthly Notices of the Royal Astronomical Society, 494(2):3014–3027, 04 2020.



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Conclusions

TMs magnetic parameters

- $|\vec{M}| = (0.245 \pm 0.081) \text{ nAm}^2 < 10 \text{ nAm}^2$
- $B_{back,x} = (414 \pm 74) \text{ nT}$
- $\nabla_x B_{back,x} = (-7400 \pm 2100) \text{ nT/m}$
- $\chi = (-3.3723 \pm 0.0069) * 10^{-5}$ at 5 mHz

Magnetic induced acceleration noise contribution to Δg

- At 1 mHz: $0.25^{+0.15}_{-0.08} fms^{-2}Hz^{-1/2} < 12 fms^{-2}Hz^{-1/2}$
- Non-stationarities increase contribution by a factor of 4.6



Thanks for you attention!

