

Curvature Wave Equations and Ringdown Non-linearities

David Pereñiguez
Niels Bohr Institute



Based on: [[Phys.Rev.D 109 \(2024\) 4, 044048](#)], and [[arXiv:2408.13557](#)].

With: Vitor Cardoso, Jaime Redondo-Yuste and Gowtham Rishi Mukkamala.

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Introduction

—— • *Nonlinear dynamical structure of GR* • ——

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- Ringdown (RD) → intrinsic structure about GR and its sources

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- Ringdown (RD) \longrightarrow intrinsic structure about GR and its sources
- $M \gtrsim 10^5 M_{\odot}$: mostly a loud RD

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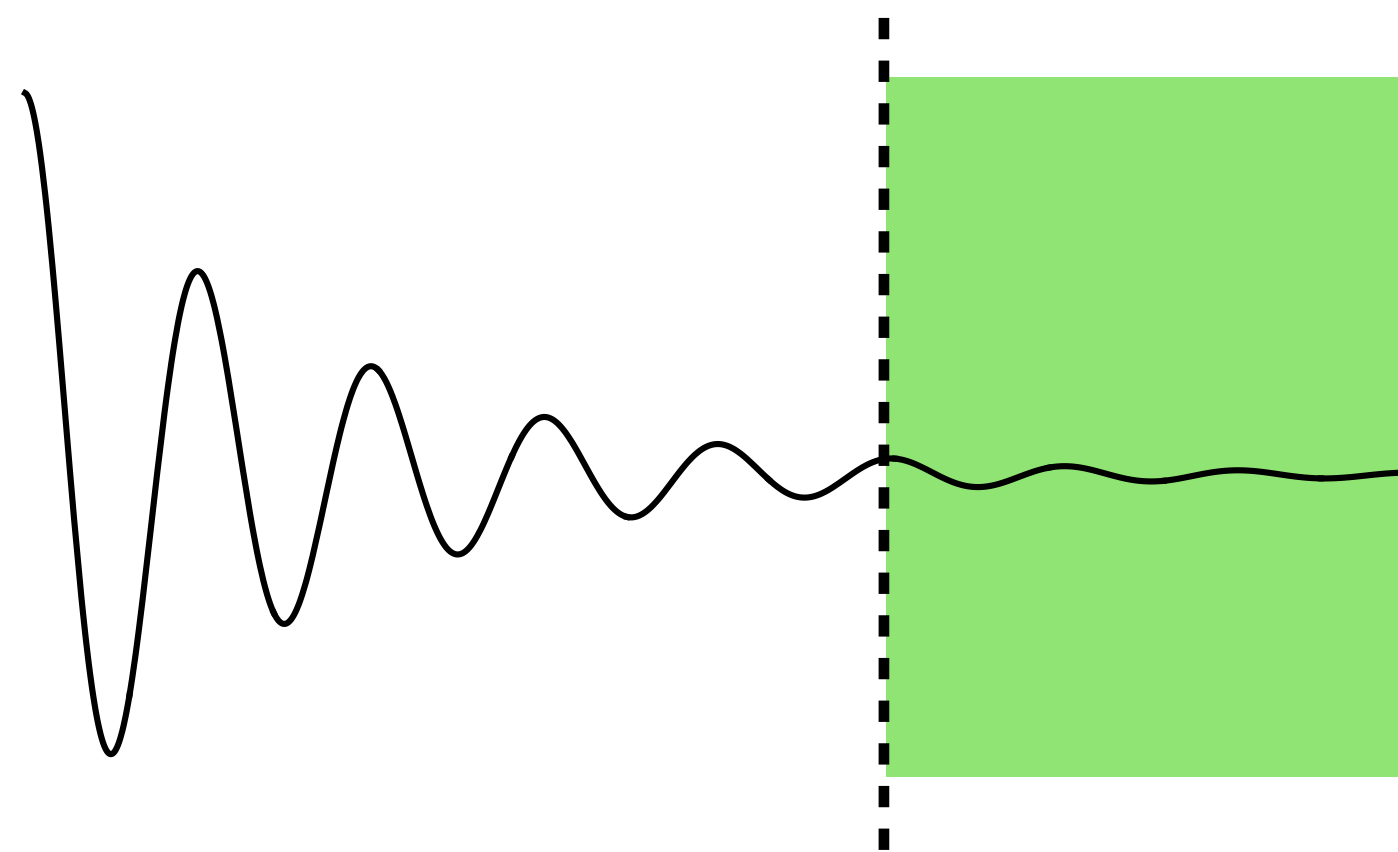
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- Limitation: linear dynamics

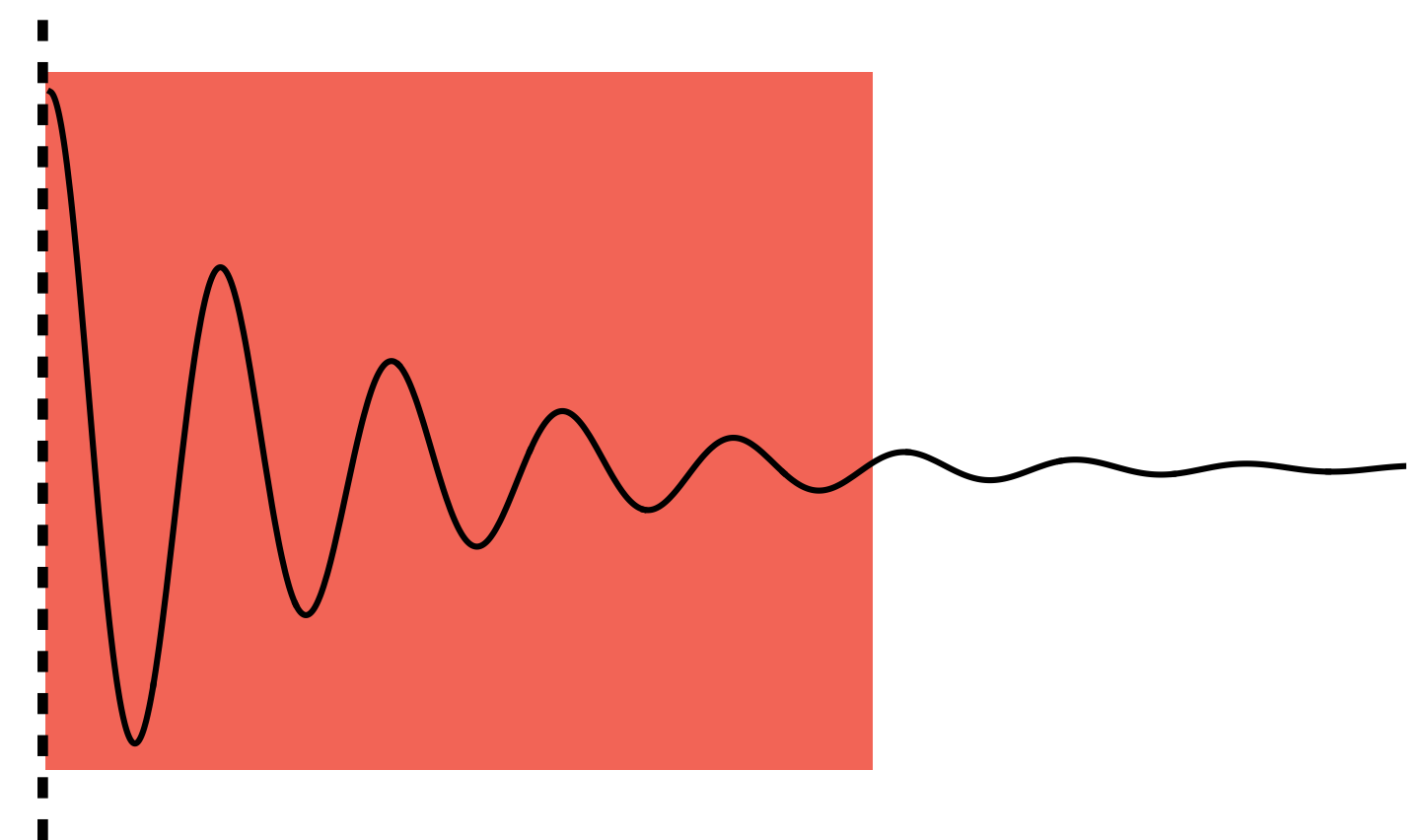
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$$t - t_{\text{merger}} \geq 15 M$$

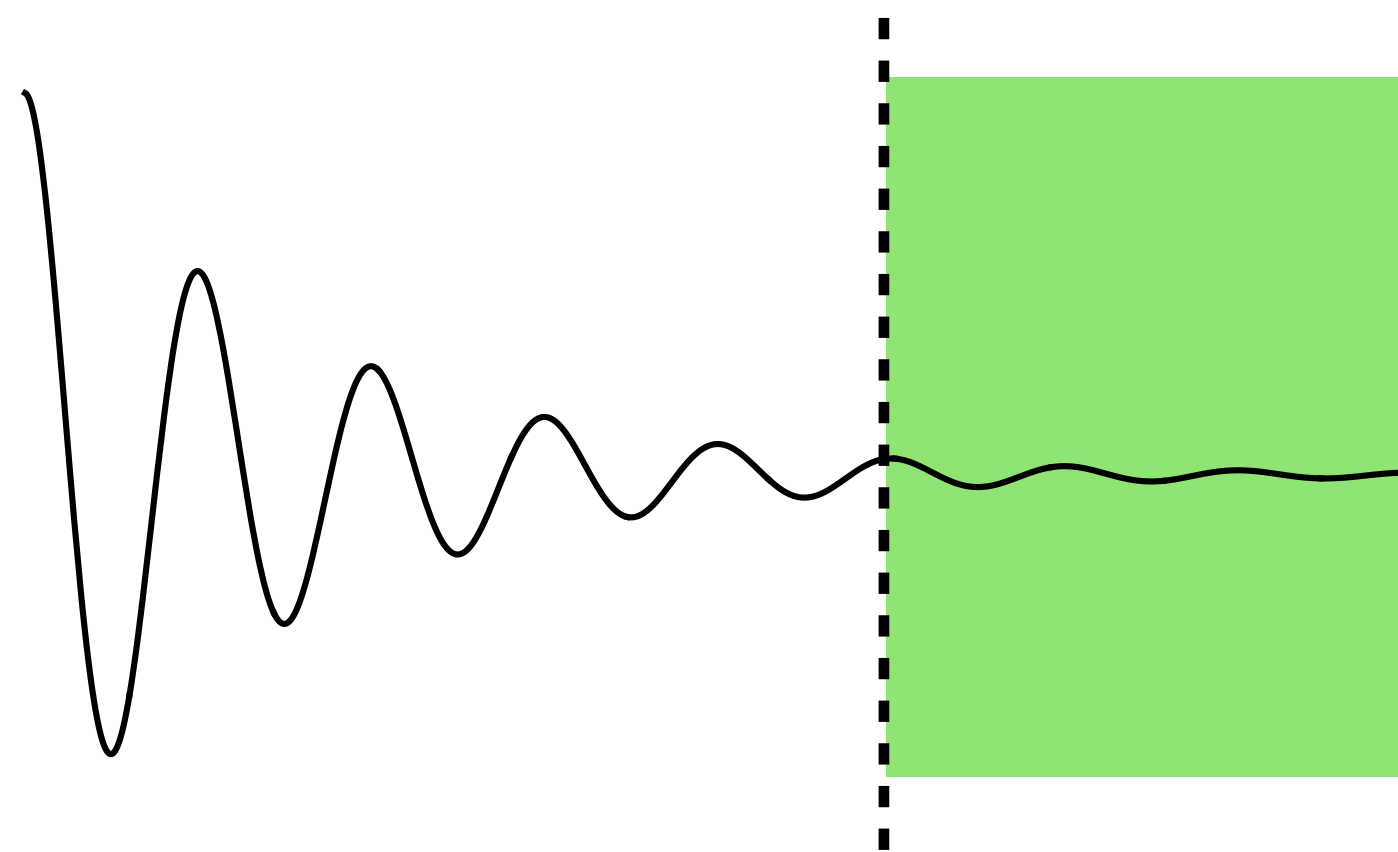


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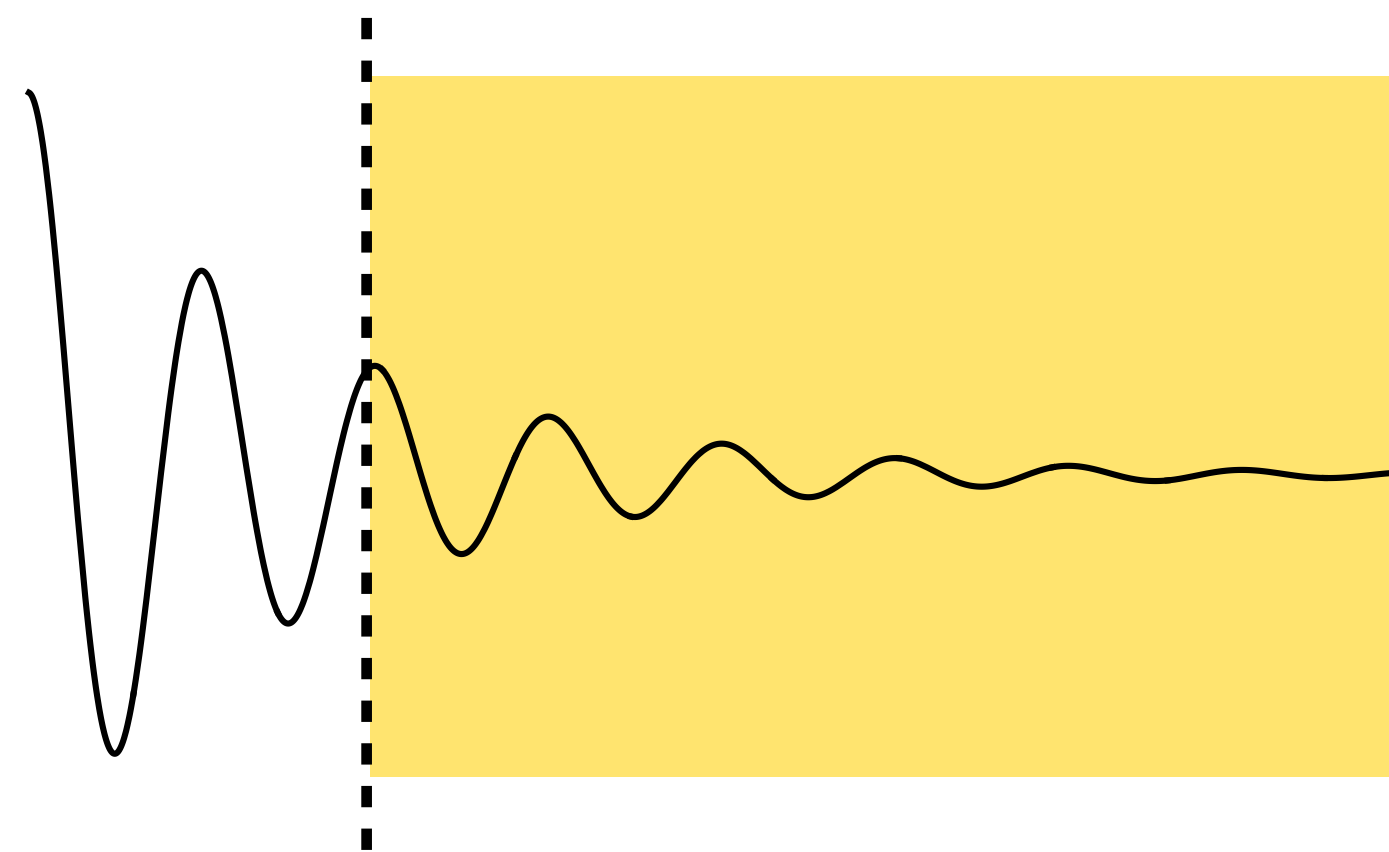
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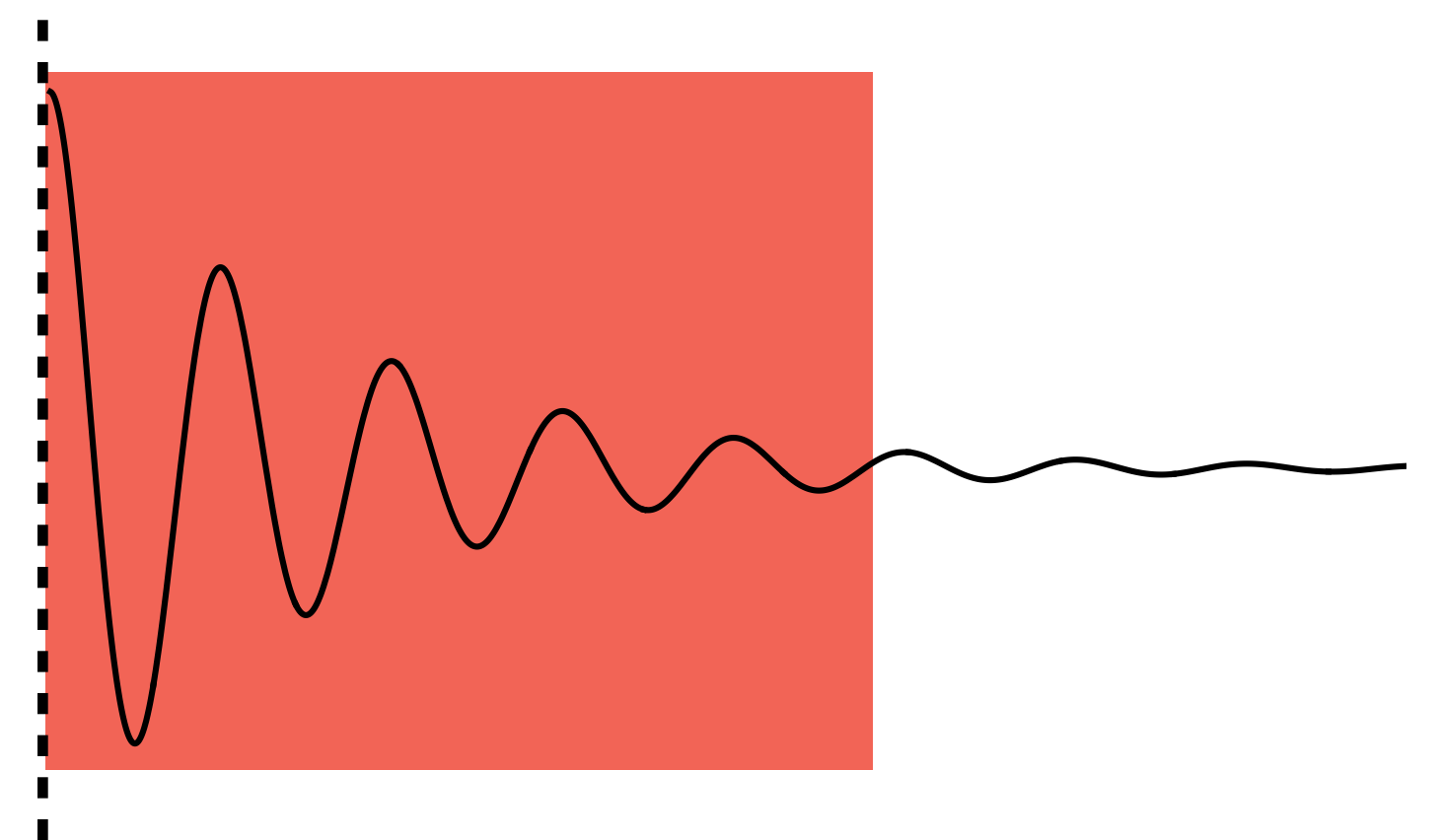
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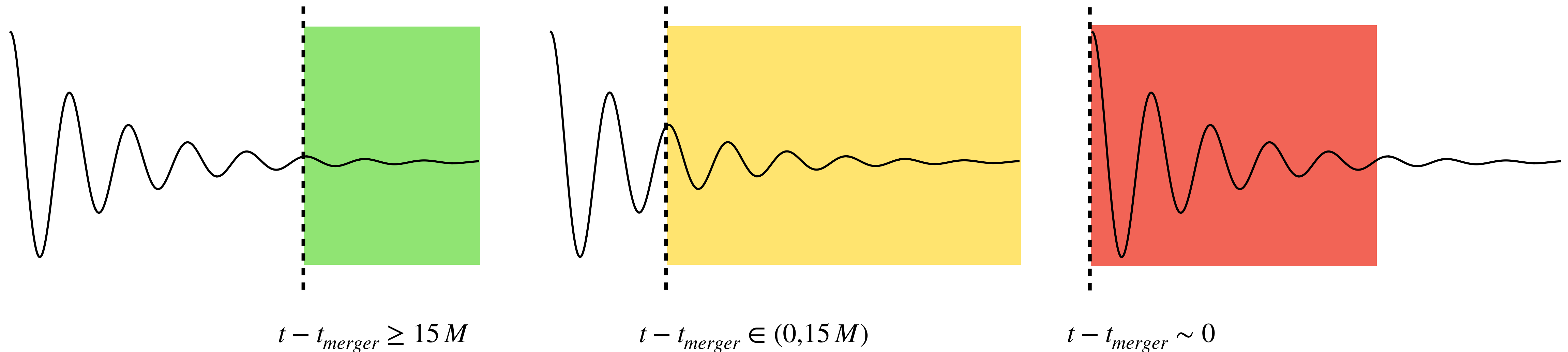


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- Goal: Nonlinear RD [LISA Waveform WG 2311.01300 '23]



State of the Art (part of)

- Quadratic fluctuations

[Gleiser+ '96, Campanelli+ '99]

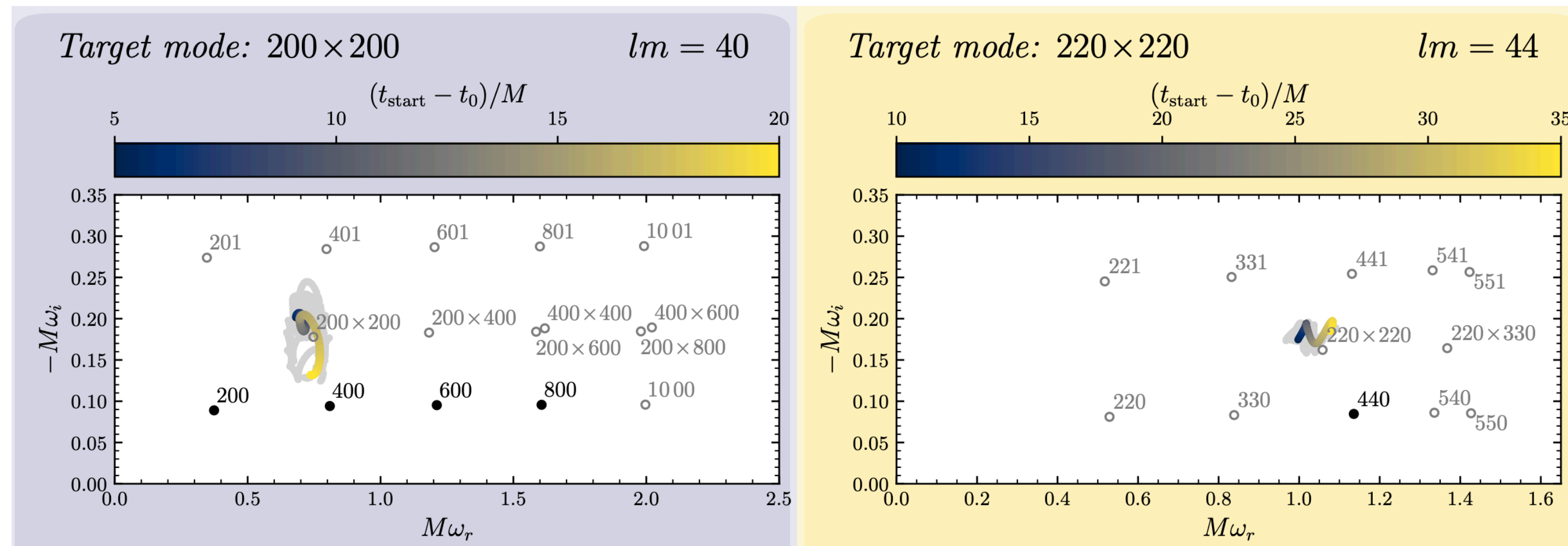
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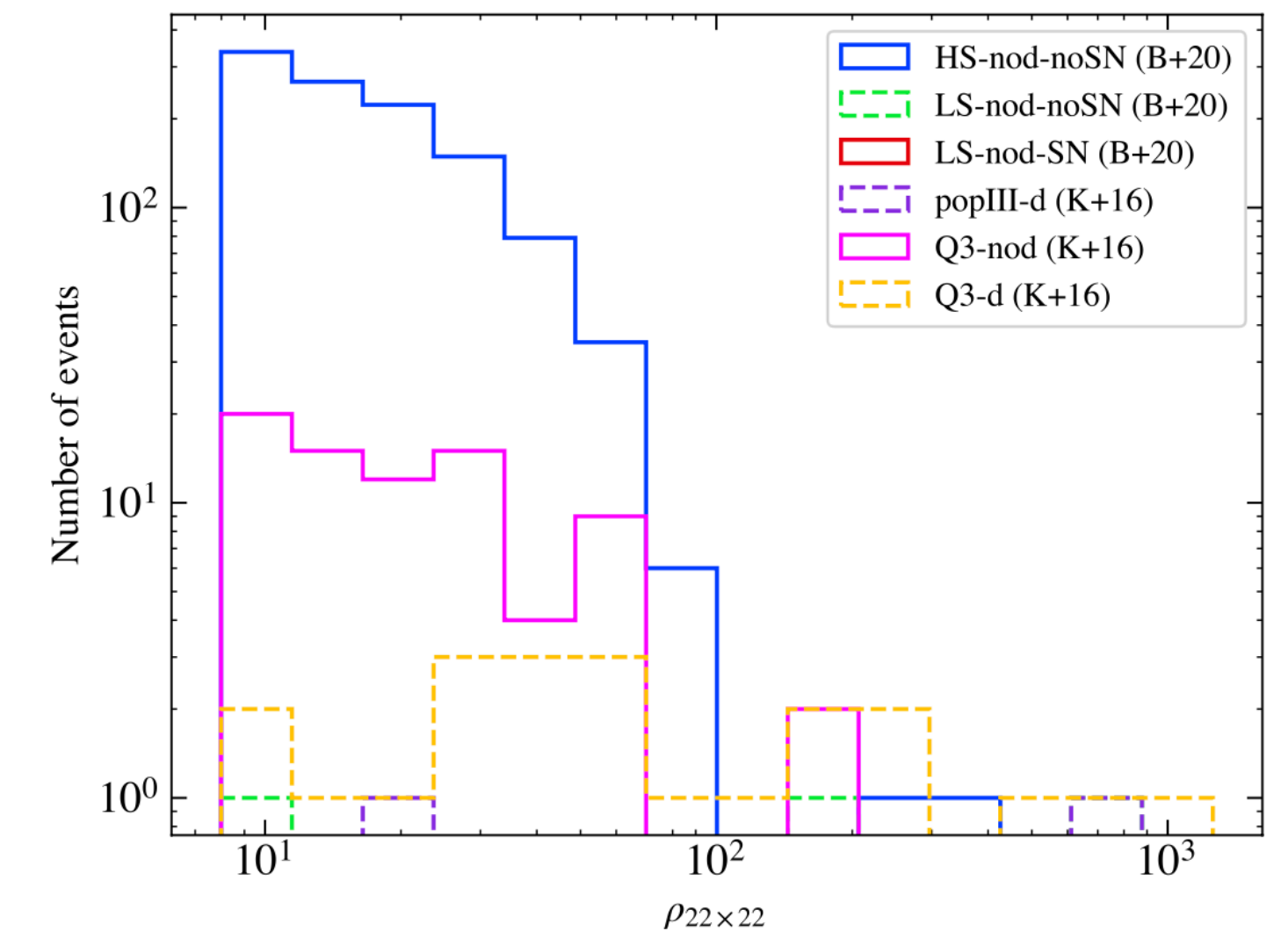
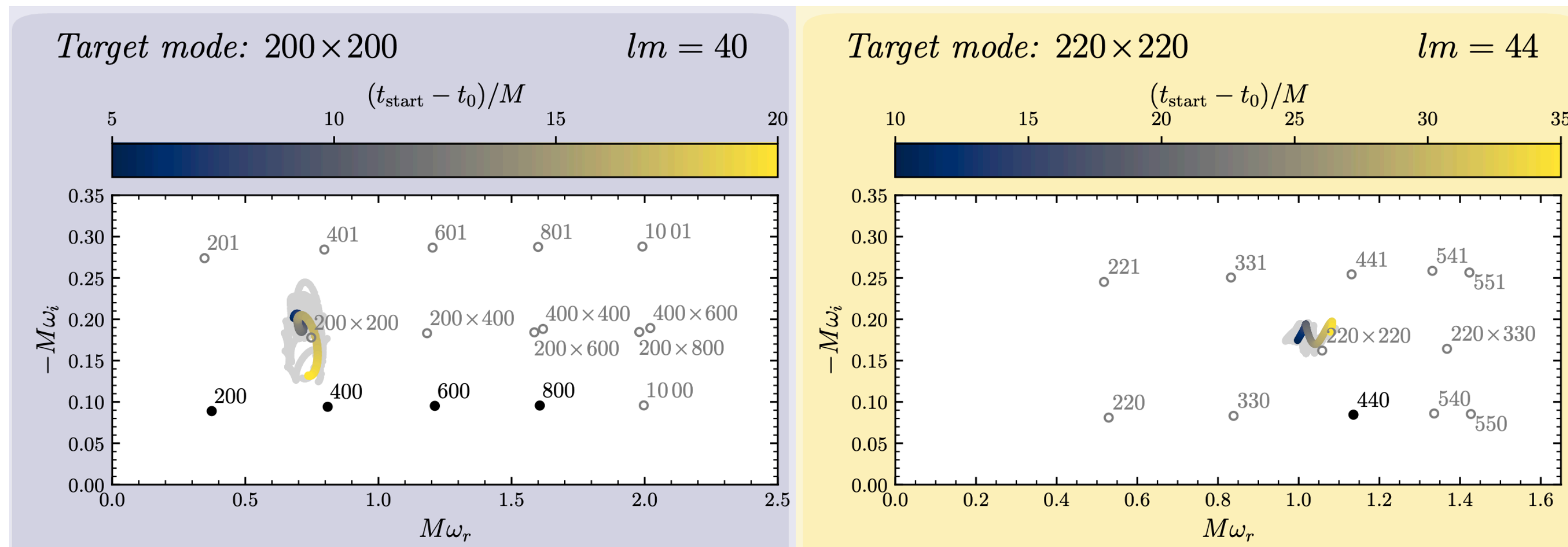
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- LISA $\sim \mathcal{O}(100)$ events per year w/ q. modes

[Yi+'24]



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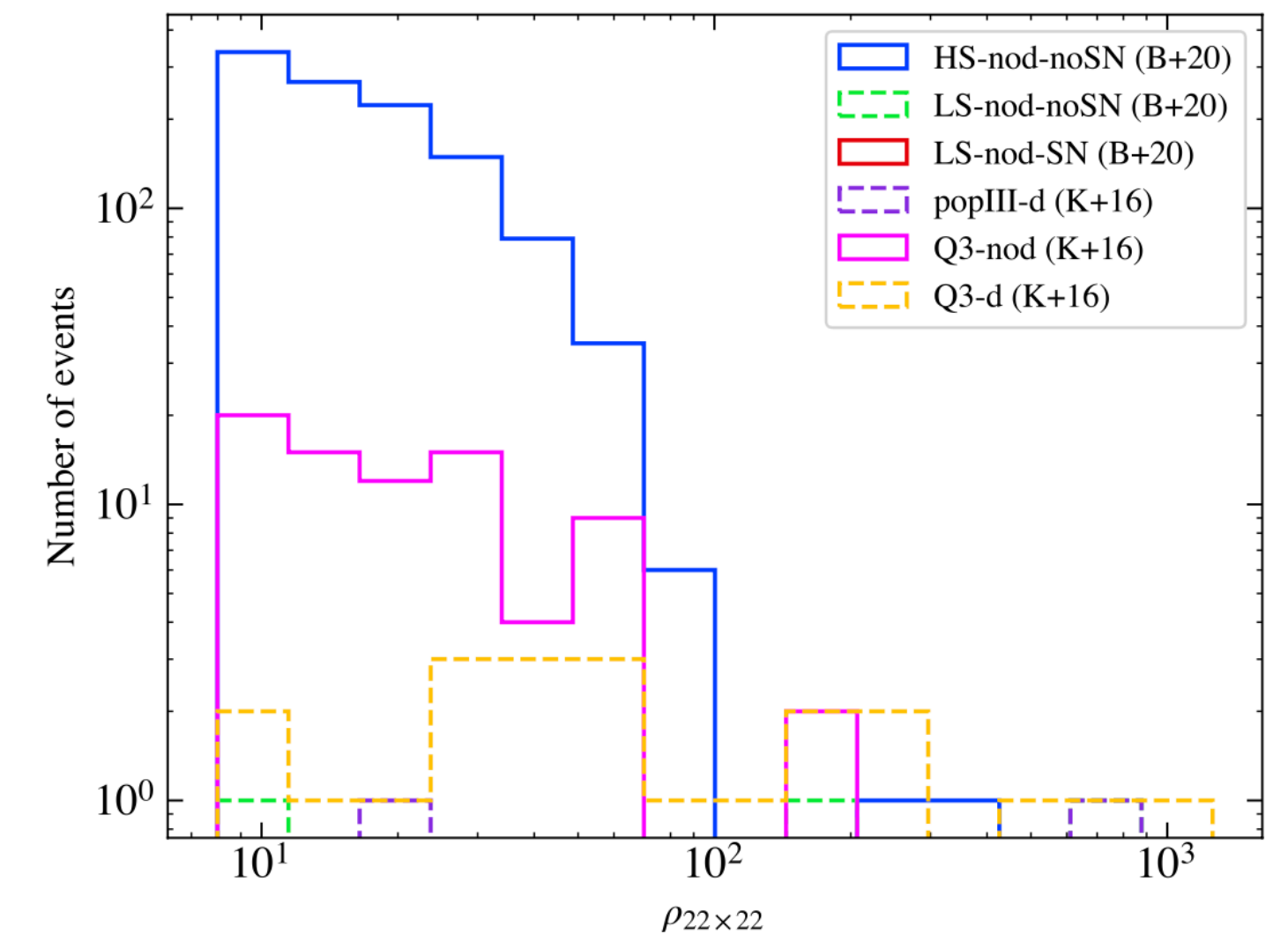
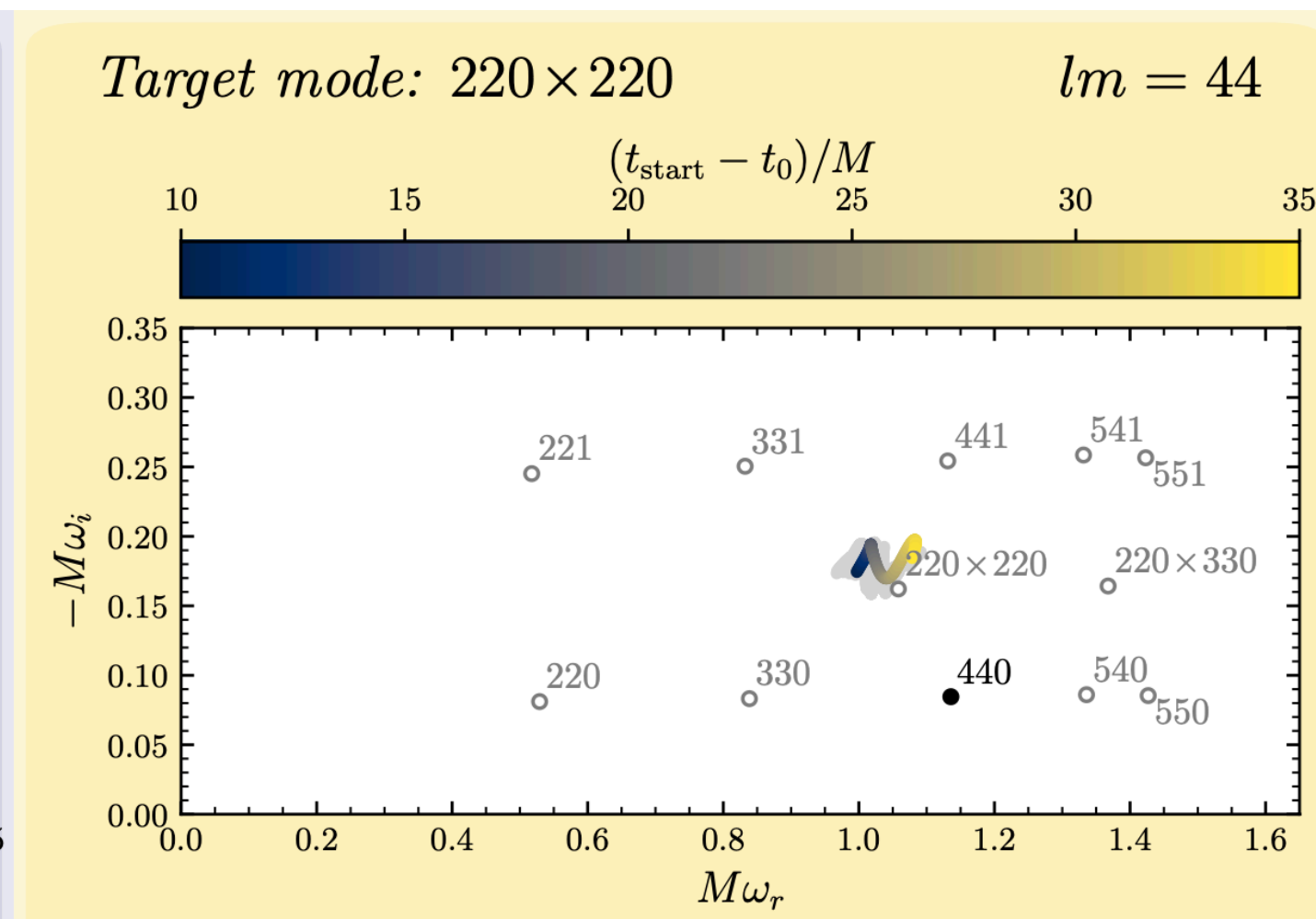
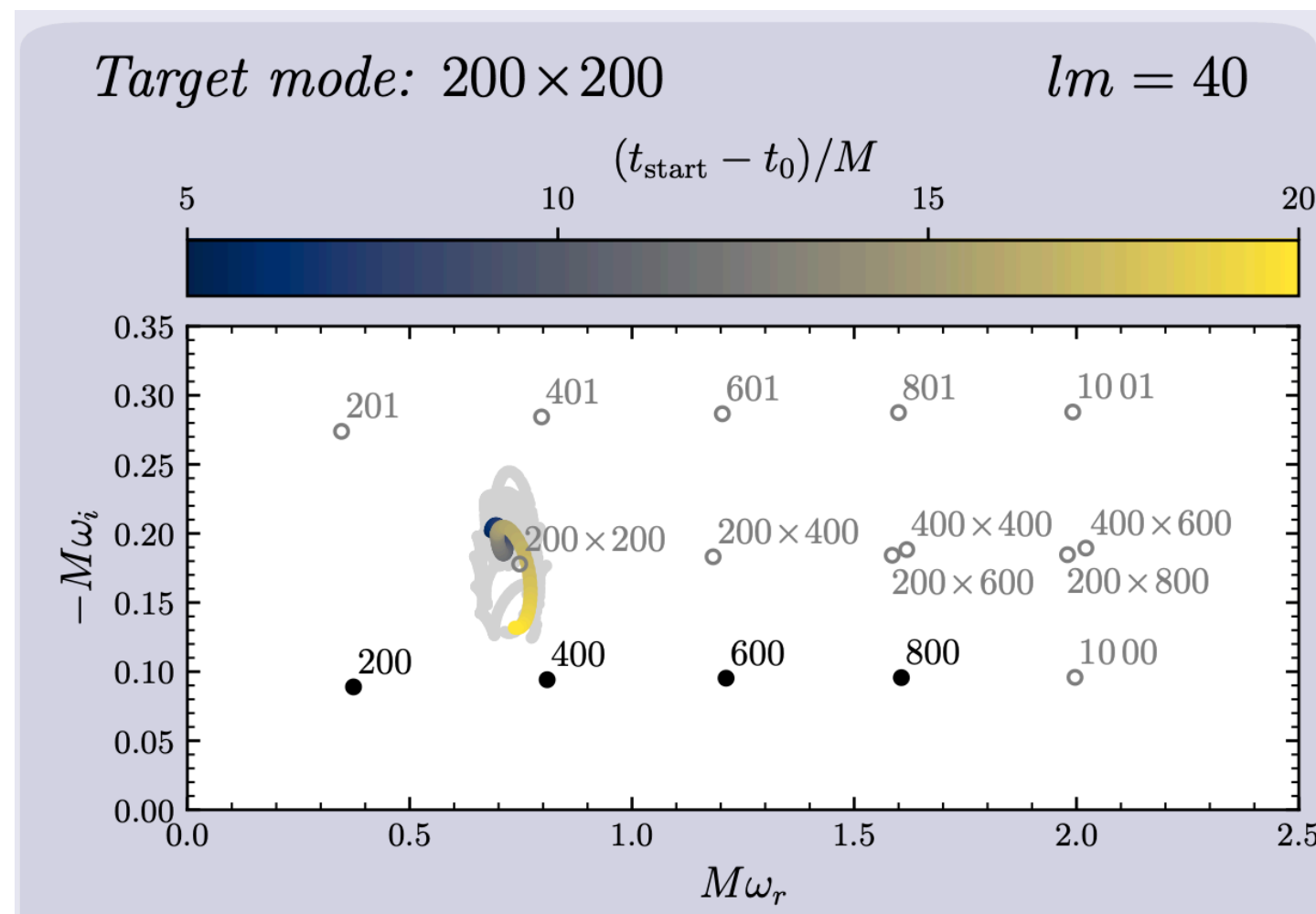
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[Yi+ '24]

→ nonlinear phenomena in RD dynamics

Absorption-Induced Excitations (AIEs)

[Bamber+ '21, Sberna+ '22, Torres+ '23, DP+ '23, May+ '24, Zhu+ '24...]



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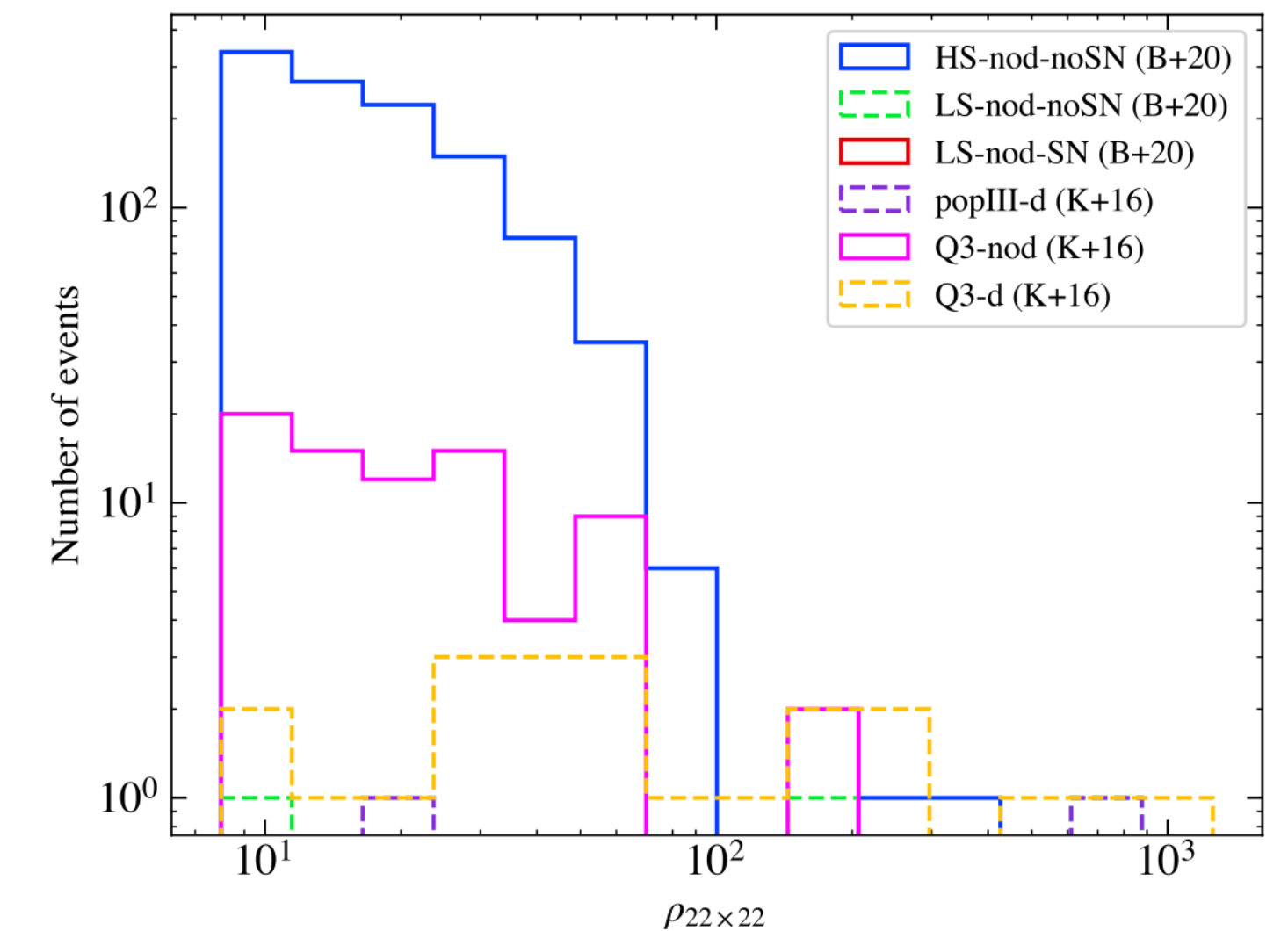
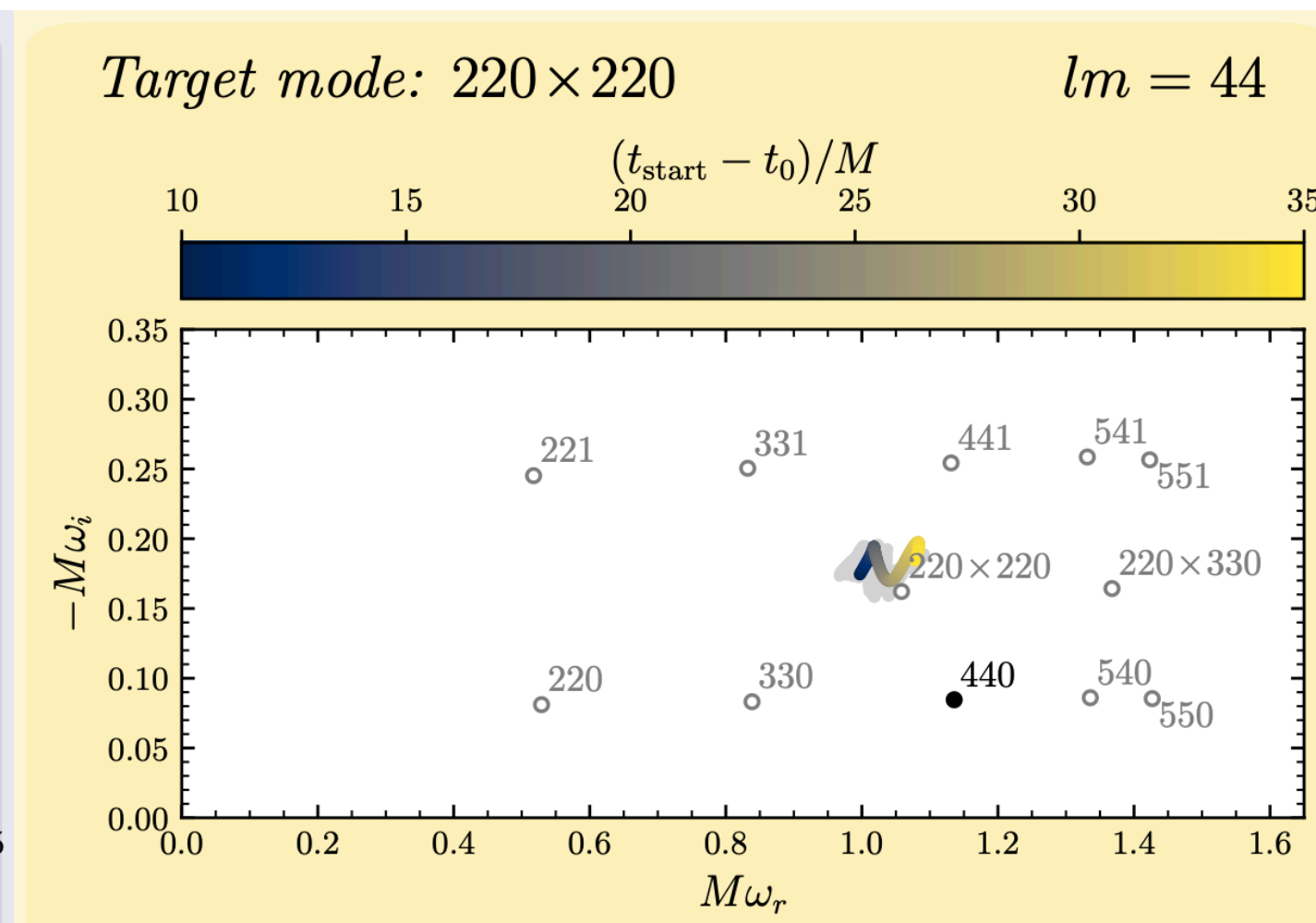
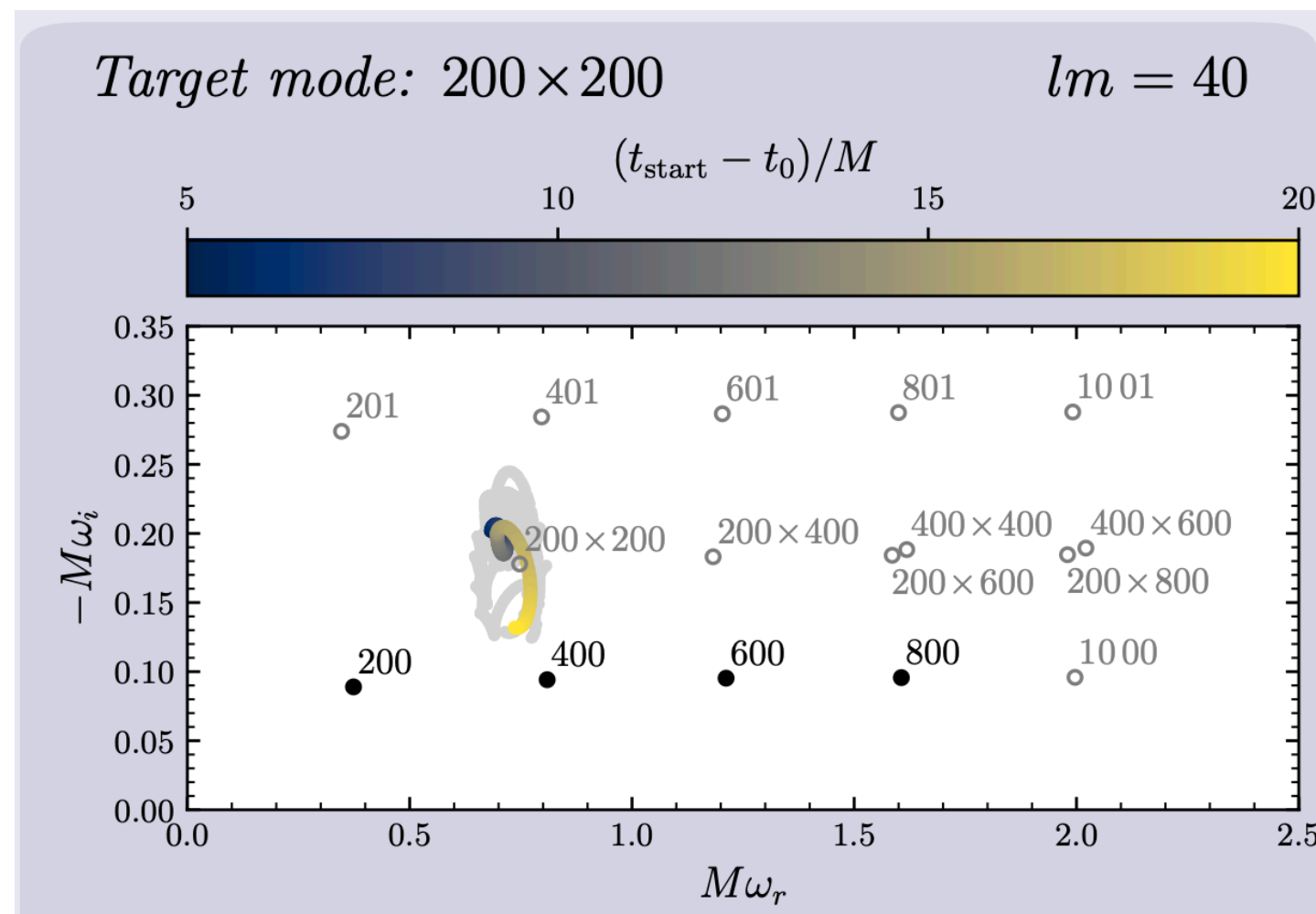
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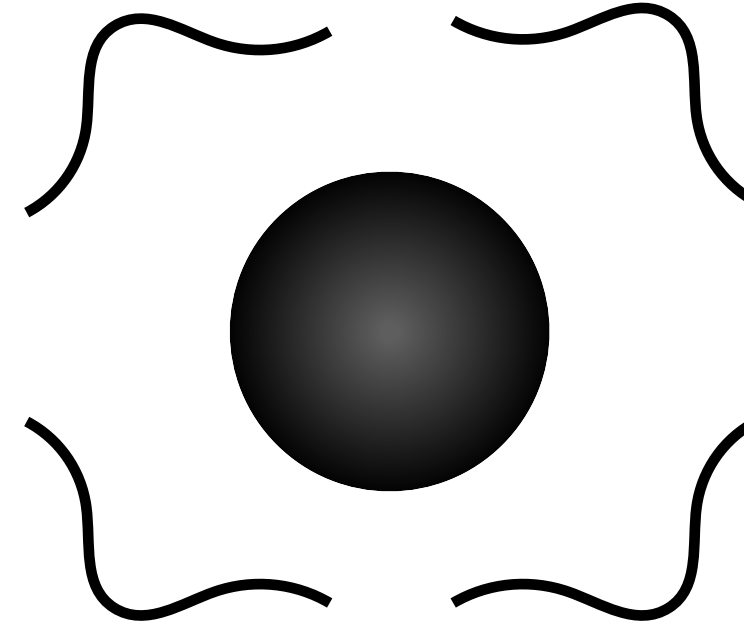
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Third-order effect in perturbation theory



Higher-order Black Hole Perturbation Theory

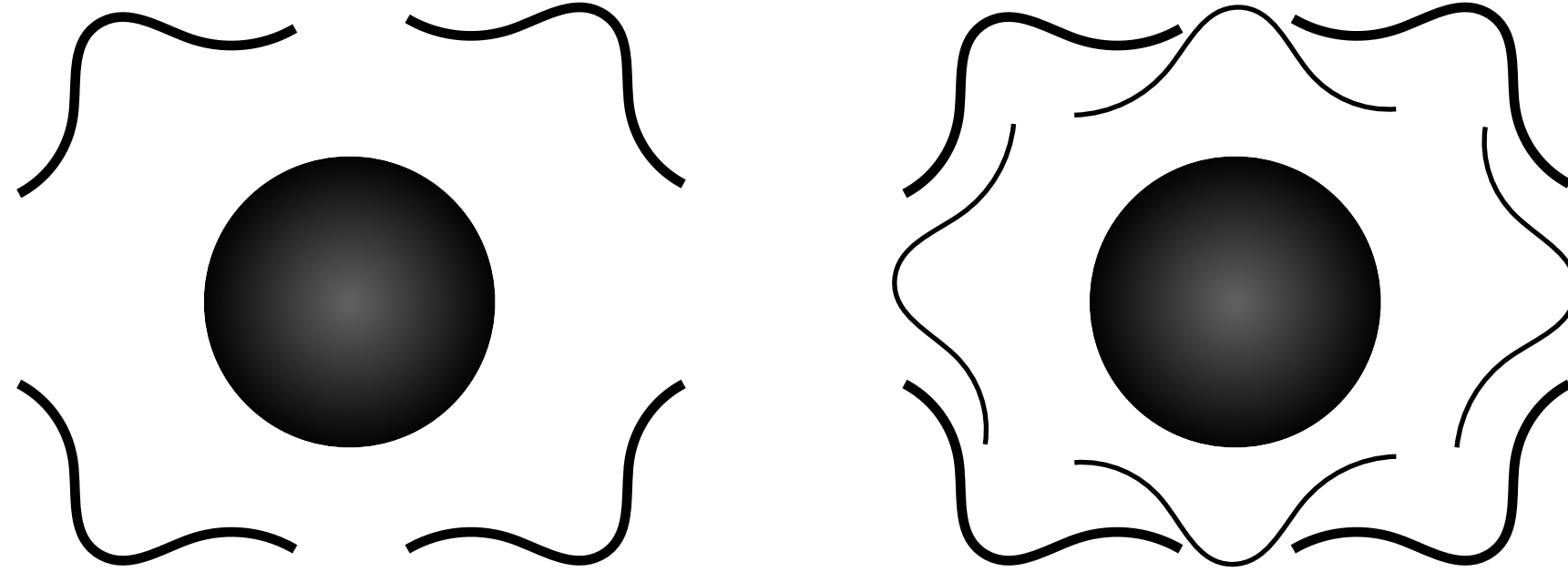
$$g_{\mu\nu} = \bar{g}_{\mu\nu} + h_{\mu\nu}^{(1)}$$



$$\mathcal{L} [h^{(1)}] = 0$$

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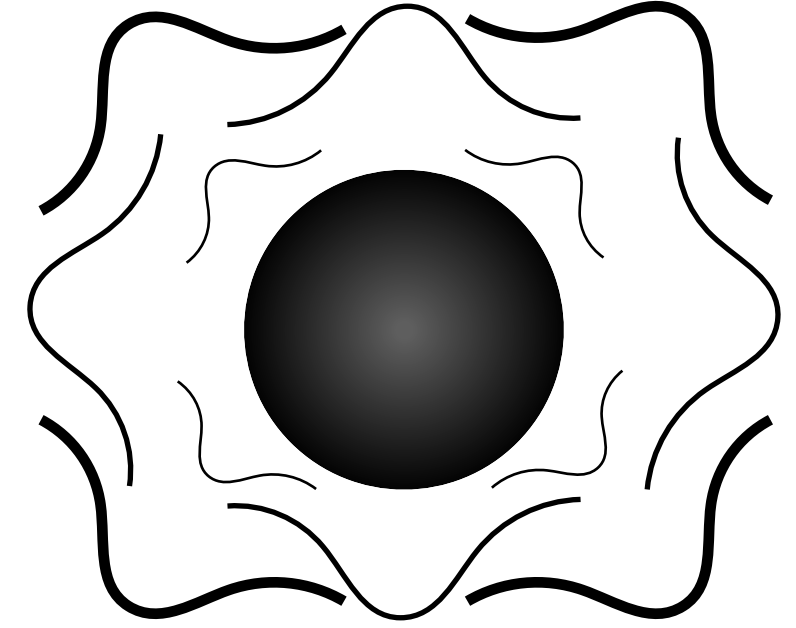
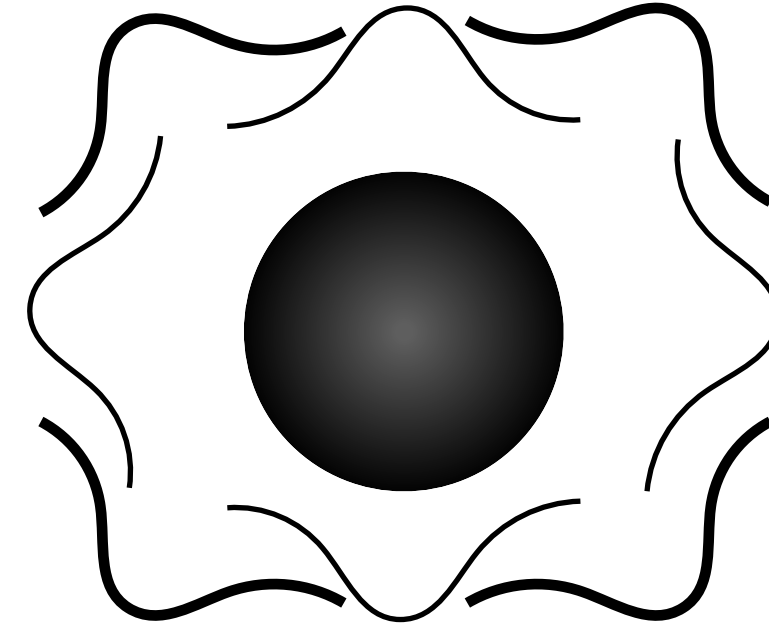
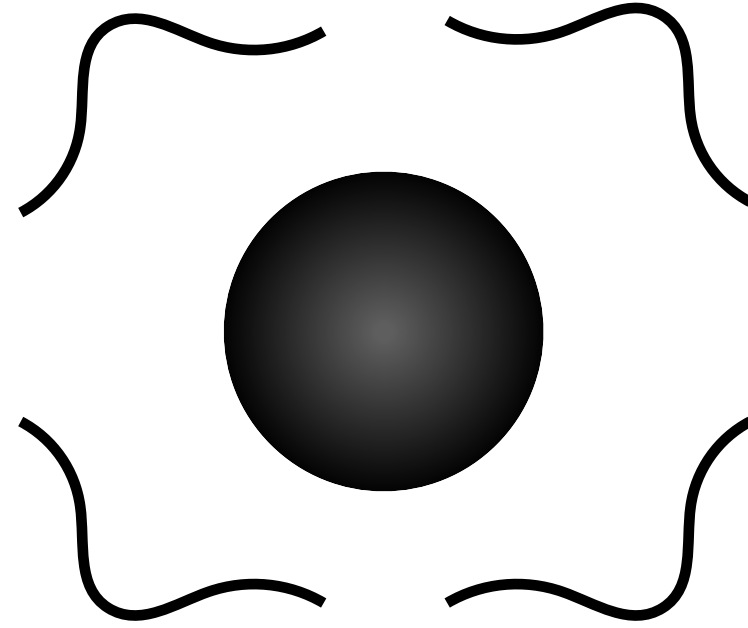
$$g_{\mu\nu} = \bar{g}_{\mu\nu} + h_{\mu\nu}^{(1)} + h_{\mu\nu}^{(2)}$$



$$\mathcal{L} [h^{(1)}] = 0, \quad \mathcal{L} [h^{(2)}] = \mathcal{S}_2 [h^{(1)}],$$

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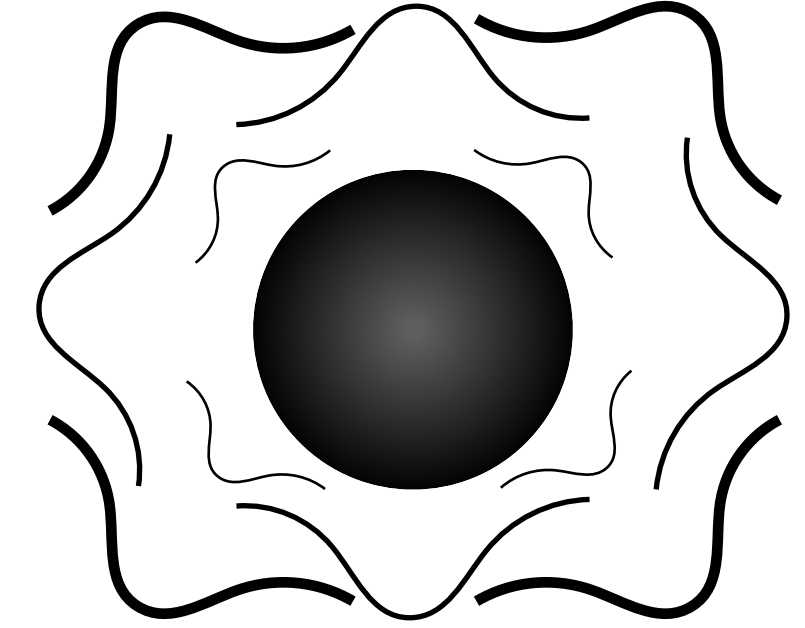
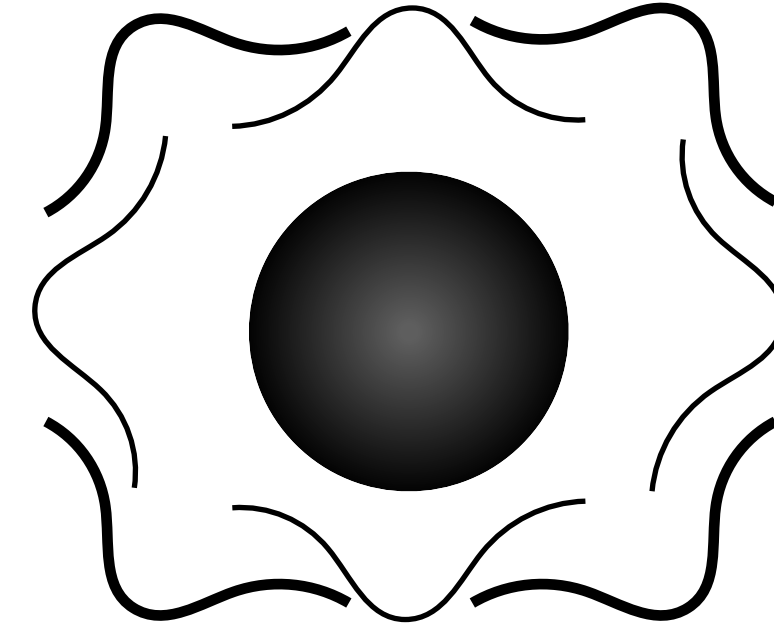
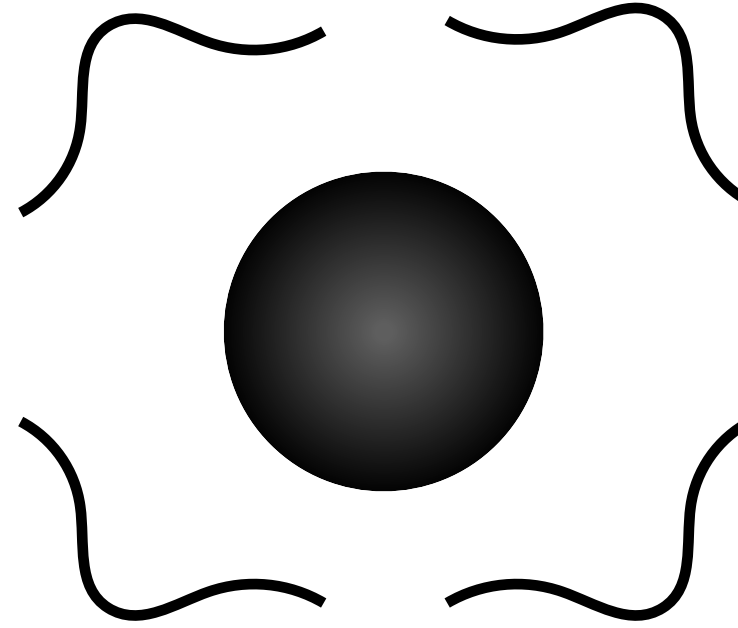
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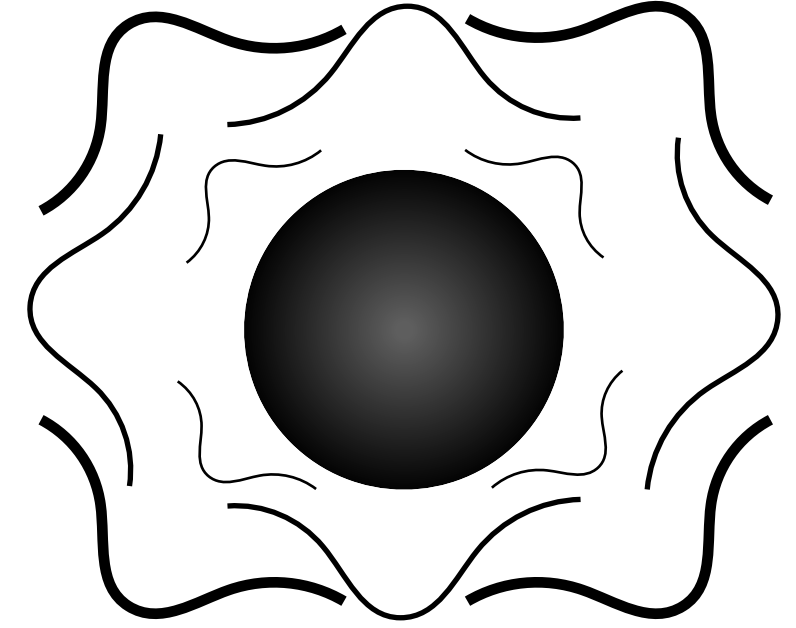
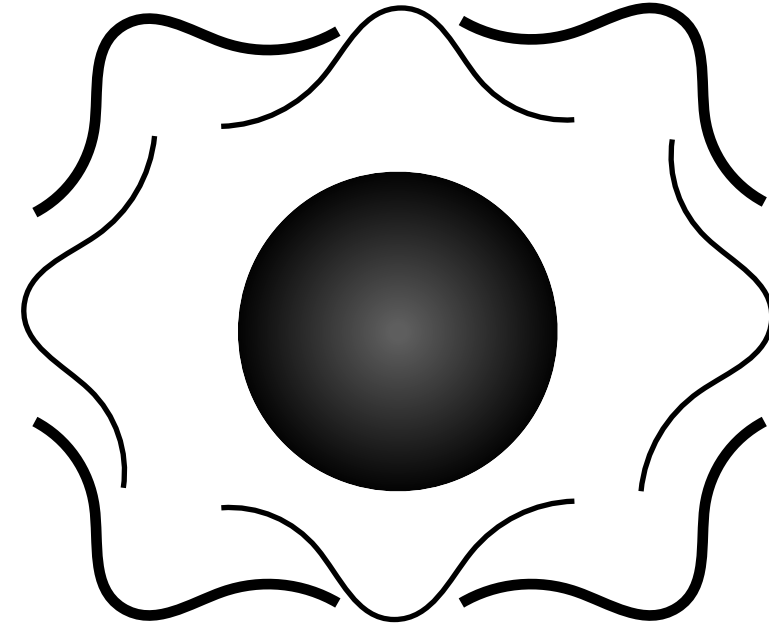
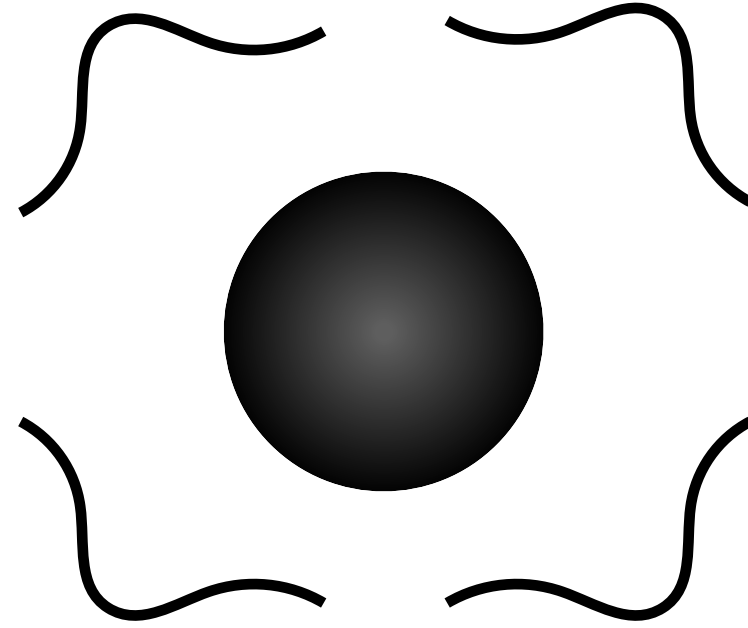
$$g_{\mu\nu} = \bar{g}_{\mu\nu} + h_{\mu\nu}^{(1)} + h_{\mu\nu}^{(2)} + h_{\mu\nu}^{(3)} + \dots$$



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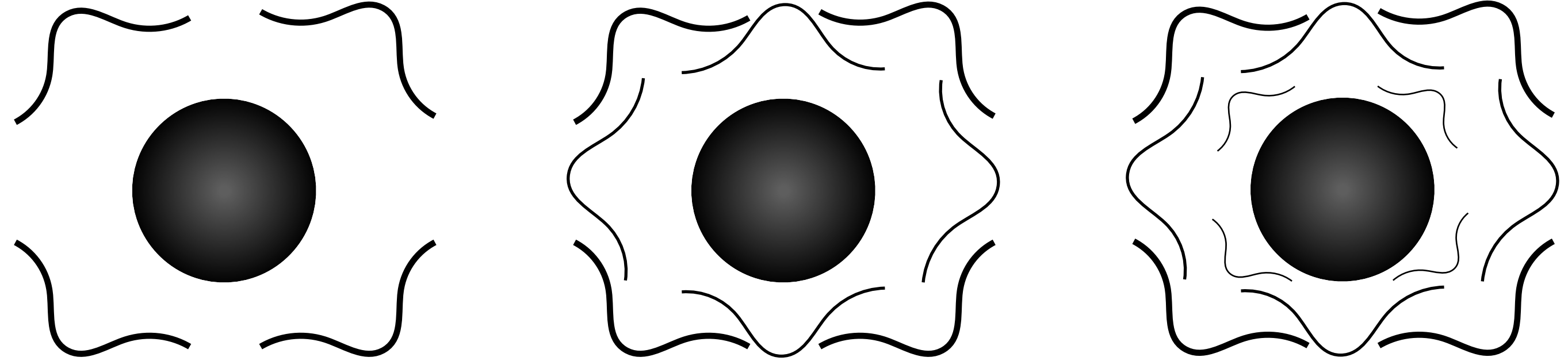


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AIEs are resonances at third-order... Long way. First steps in this talk:

Improvement at 1st Order

A geometric approach to RWZ

[Mukkamala and DP '24]

An effective approach to AIEs

Fluctuations of dynamical BHs

[Redondo-Yuste, DP and Cardoso '24]

Master Wave Equations *à la* Regge-Wheeler-Zerilli +

Spherical Symmetry: an extremely powerful assumption

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Background

$$ds^2 = g_{ab}(y)dy^a dy^b + r^2(y)\Omega_{AB}dz^A dz^B$$

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Even Sector

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$$P_a = -\tilde{\square}\tilde{h}_a + \tilde{h}^b{}_{:ab} + \frac{2}{r}\left(r^{,b}\tilde{h}_{b:a} - r_{,a}\tilde{h}^b{}_{:b}\right)$$

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$$P = \tilde{h}^a{}_{:a}$$

[Regge
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$$Q_{ab} = \tilde{p}^c{}_{(a;b)c} - \frac{1}{2}g_{ab}\tilde{p}^{cd}{}_{:cd} - \frac{1}{2}\tilde{p}^c{}_{c:ab} - \frac{1}{2}(\tilde{\square}\tilde{p}_{ab} - g_{ab}\tilde{\square}\tilde{p}^c{}_c)$$

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$$\left(\square - V_{RW,ZM}(r)\right)\Phi_{RW,Z} = 0$$

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Master Wave Equations

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- EOMs reduce to single decoupled wave equation
[Regge, Wheeler, Zerilli +]
- Even and odd sectors are *isospectral*
[Chandrasekar +]

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$$Q^b = \tilde{\square}\tilde{p}^a{}_a - \tilde{p}^{ab}{}_{:ab} - \frac{2}{r}r^{,b}\tilde{p}^a{}_{b:a} + \frac{r^{,a}}{r}\tilde{p}^b{}_{b:a} - \frac{1}{2}\frac{l(l+1)}{r^2}\tilde{p}^a{}_a + \frac{2}{r}r^{,a}\tilde{K}_{,a} + \tilde{\square}\tilde{K}$$

$$Q^\sharp = -\tilde{p}^a{}_a$$

Master Wave Equations

$$\left(\square - V_{RW,ZM}(r)\right)\Phi_{RW,Z} = 0$$

- EOMs reduce to single decoupled wave equation
[Regge, Wheeler, Zerilli +]
- Even and odd sectors are *isospectral*
[Chandrasekar +]
- Non-systematic... Geometric origin?
- Artificial difference between sectors
→ need 2 propagators

Master Wave Equations *à la* Regge-Wheeler-Zerilli +

Spherical Symmetry: an extremely powerful assumption

Background

$$ds^2 = g_{ab}(y)dy^a dy^b + r^2(y)\Omega_{AB}dz^A dz^B$$

$$h = h_{ab}(y)Y(z)dy^a dy^b + k(y)Y(z)\Omega_{AB}dz^A dz^B$$

$$+ 2j_a(y)X_A(z)dy^a dz^A \quad \text{Even Sector}$$

Odd Sector

Odd Sector

$$P_a = -\tilde{\square}\tilde{h}_a + \tilde{h}^b{}_{:ab} + \frac{2}{r}\left(r^{,b}\tilde{h}_{b:a} - r_{,a}\tilde{h}^b{}_{:b}\right)$$

$$+ \frac{1}{r^2}\left(l(l+1) - \frac{2M}{r}\right)\tilde{h}_a - \frac{2}{r^2}r_{,a}r^{,b}\tilde{h}_b$$

$$P = \tilde{h}^a{}_{:a}$$

[Regge
and Wheeler '57]

Even Sector

$$Q_{ab} = \tilde{p}^c{}_{(a;b)c} - \frac{1}{2}g_{ab}\tilde{p}^{cd}{}_{:cd} - \frac{1}{2}\tilde{p}^c{}_{c:ab} - \frac{1}{2}(\tilde{\square}\tilde{p}_{ab} - g_{ab}\tilde{\square}\tilde{p}^c{}_c)$$

$$+ \frac{2}{r}r_{,c}(\tilde{p}^c{}_{(a;b)} - g_{ab}\tilde{p}^{cd}{}_{:d}) - \frac{r^{,c}}{r}(\tilde{p}_{ab:c} - g_{ab}\tilde{p}^d{}_{d:c}) + \frac{l(l+1)}{2r^2}\tilde{p}_{ab}$$

$$- \frac{1}{r^2}g_{ab}r^{,c}r^{,d}\tilde{p}_{cd} - \frac{1}{2r^2}g_{ab}\left[l(l+1) + \frac{2M}{r}\right]\tilde{p}^c{}_c$$

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[Lenzi and Sopuerta '21]

Master Wave Equations From Curvature Wave Equations

$$U(1) : \quad \mathbb{F}_{\mu\nu} \equiv F_{\mu\nu} - i \star F_{\mu\nu}, \quad \Delta \equiv \star d \star d + d \star d \star, \quad \longrightarrow \quad \boxed{\Delta \mathbb{F}_{\mu\nu} = 0}$$

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Expand self-dual curvature & Linearise CWEs

Applied to Teukolsky's derivation

[Ryan '74]

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Use to geometrize RWZ

[Mukkamala and DP '24]

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as a corollary*

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Geometric origin of mast. eqs.



Even-Odd-symmetric



Bonus: RW eq. for even sector!

Ringdown of Dynamical Spacetimes

An effective approach to gravitational AIEs

Pure radiation field:

$$G_{\mu\nu} = \Phi K_\mu K_\nu, \quad K^\mu K_\mu = 0$$

Solution (Vaidya $m(v)$):

$$ds^2 = -f(v, u)du dv + r^2(u, v)d\Omega^2$$

Wave equation:

$$\left[\partial_{uv}^2 - \frac{f}{r} \left(\frac{3m(v)}{r^2} - \frac{l(l+1)}{2r} \right) \right] \Psi = \frac{2f}{r^2} F(v)$$

Scalar and electromagnetic relaxation: [\[Abdalla+ '06, Chirenti+ '11\]](#)

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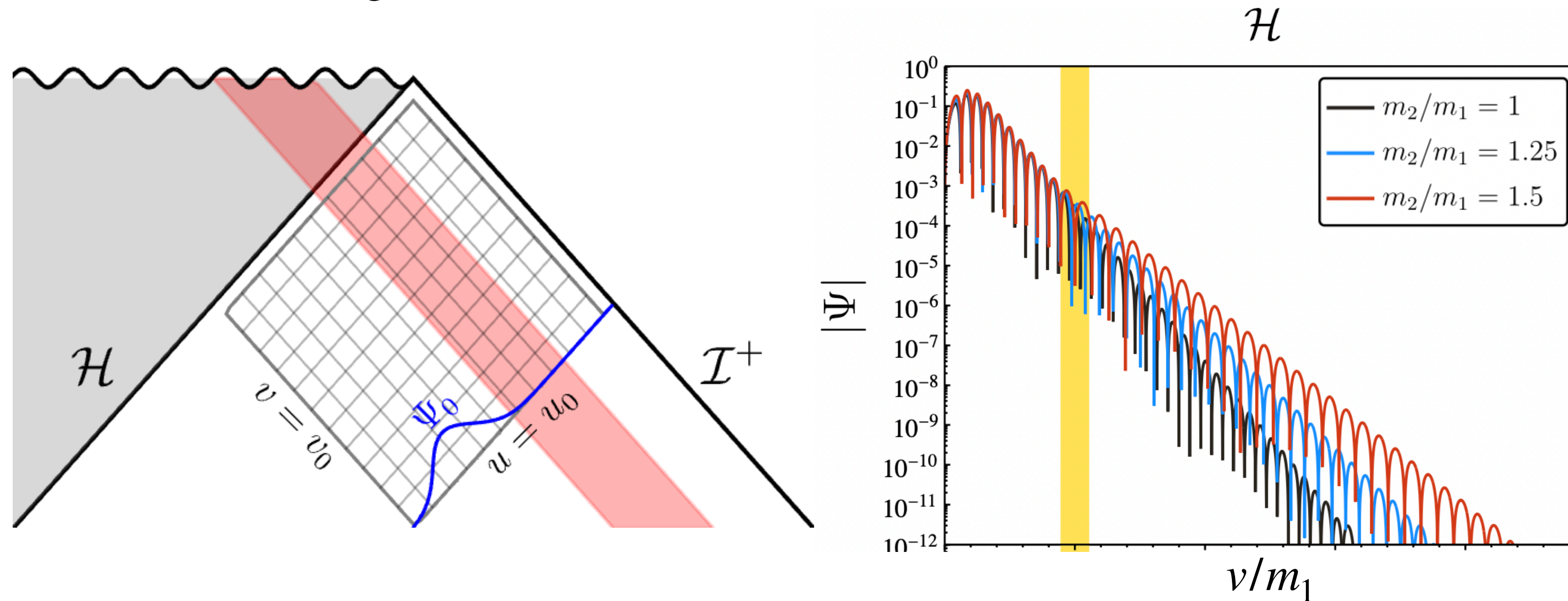
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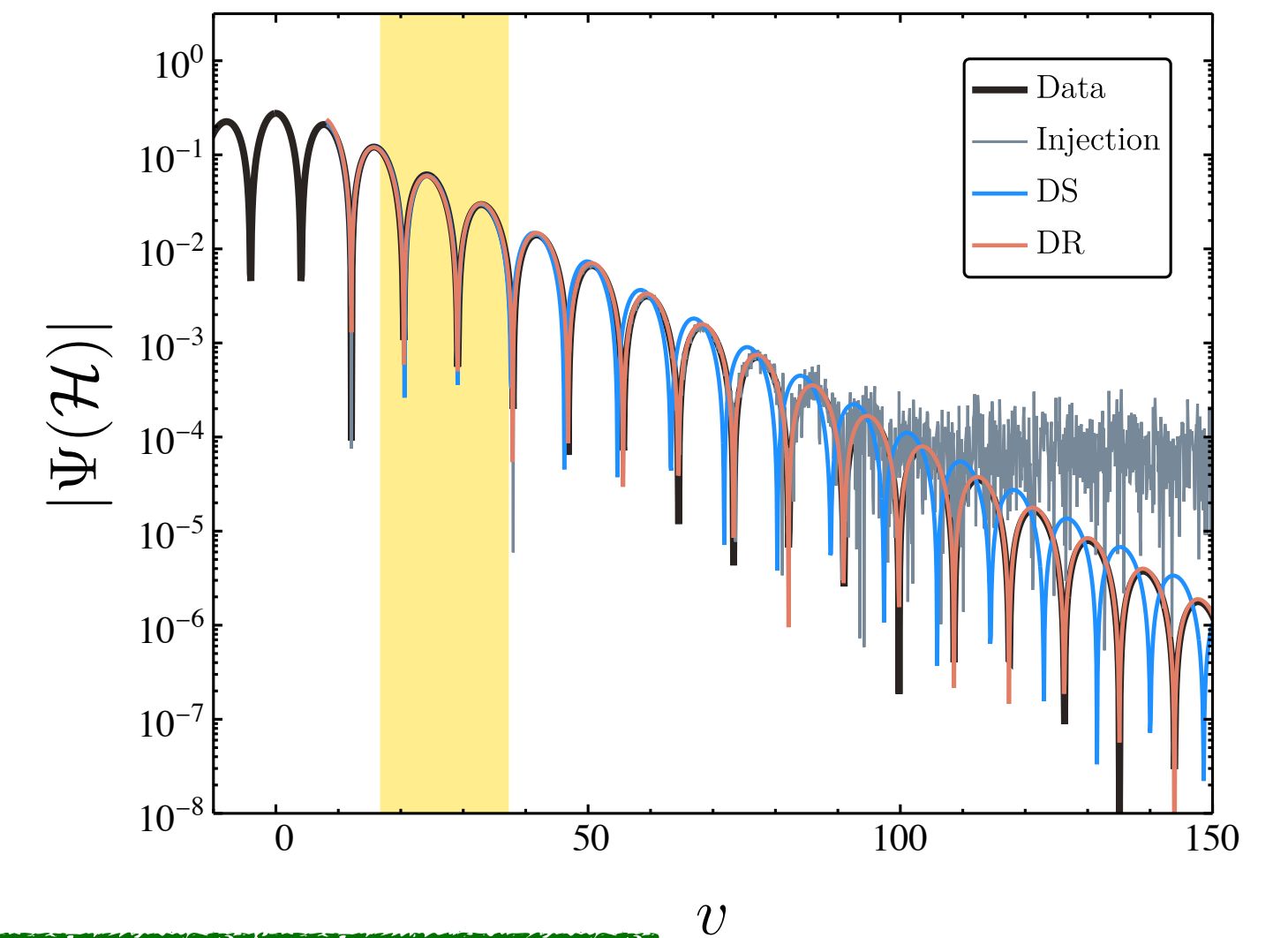
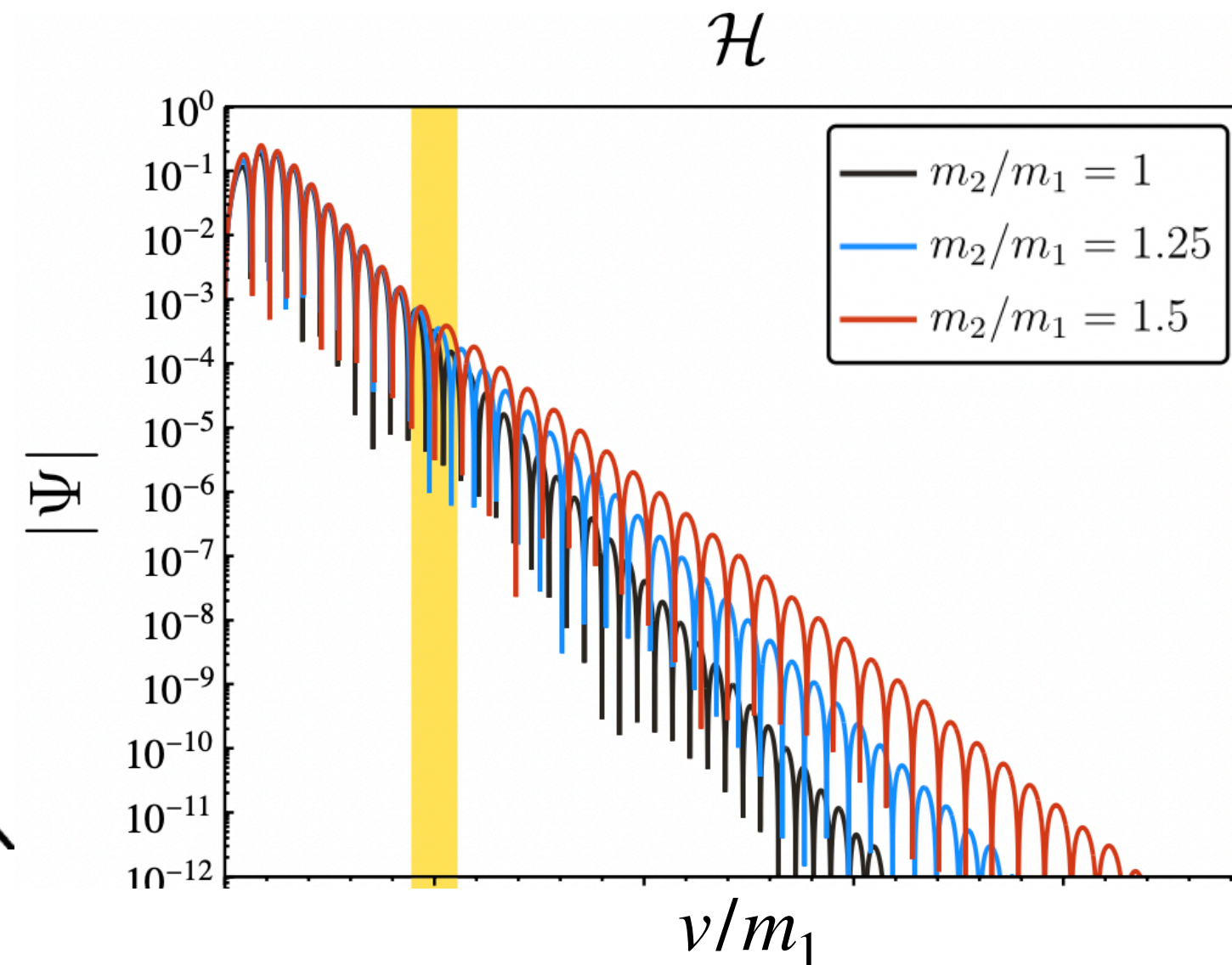
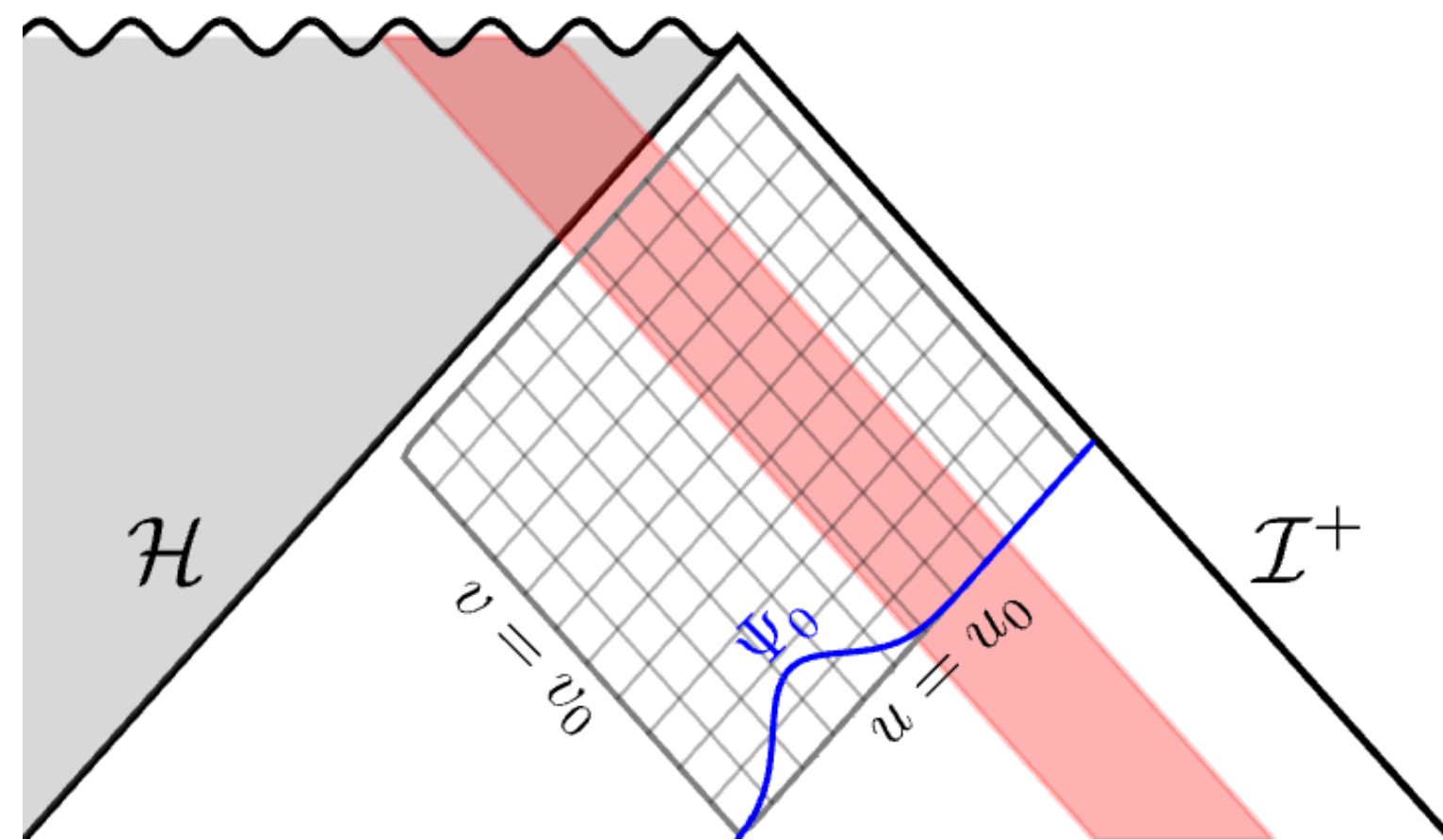
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$$m(v)\omega(v) \sim \text{constant} = m_{1,2}\omega_{1,2} \rightarrow$$

$$\Psi = A(v)e^{i\omega(v)v} + c.c., \quad A(v) = \tilde{A} \left[1 + \tilde{Q} \frac{\delta m(v)}{m_2 - m_1} \right], \quad \omega(v) = \frac{m_1 \omega_{220}}{m(v)}$$

Dynamical Ringdown

Summary and future directions

- New approach to perturbation theory in spherical symmetry based on CWEs:
 - Geometric origin of master wave equations (simplifies derivation drastically)
 - Even-odd symmetric \rightarrow isospectrality
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