

The `ztf_in2p3` pipeline for DR2.5 & beyond.

a.k.a `ztf_in2p3` : what does this pipeline do.

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The *ztf**in2p3* pipeline

How does it work ?

What does it do ?

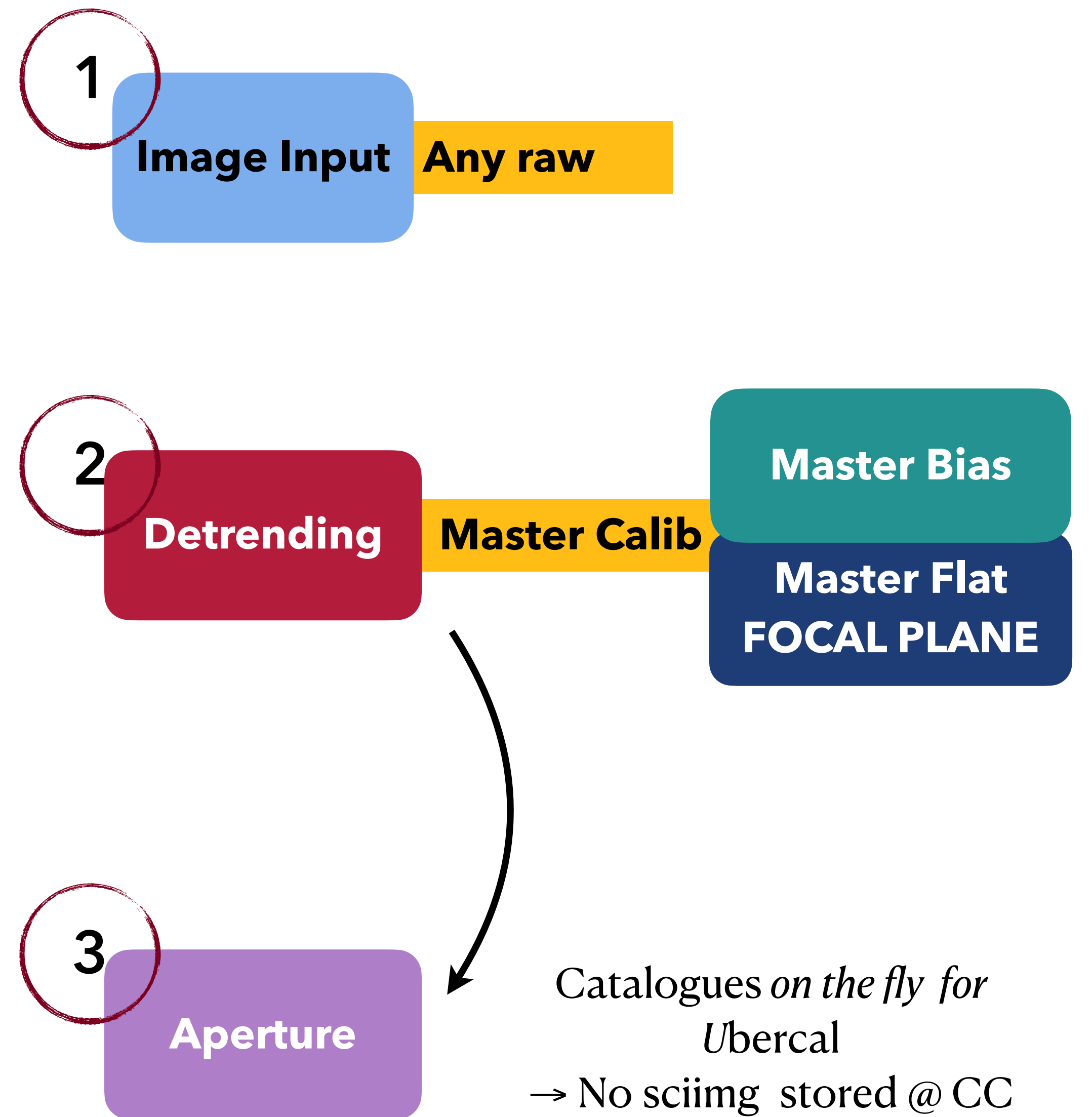
- Image I/O and handling (e.g WCS) : *ztfimg*
- Calibration and everything else : *ztf**in2p3*
- 3 Main steps from *Raw* to *Aperture catalogues* :

Pipeline is developed within two dedicated python package :

- *ztf**in2p3* (Pipes)
- *ztfimg* (I/O + pixel corrections)

*ztf**in2p3* (Pipes) : Made to work @ CC.

ztfimg (Pipes) : Should work on any laptop



The *ztf**in2p3* pipeline



1 Reading raw images (CCDs)

- Image Input and handling (e.g WCS) : `ztfimg`
- Image Input is the basis → Applies to any raw opened in the pipeline :

Image Input :

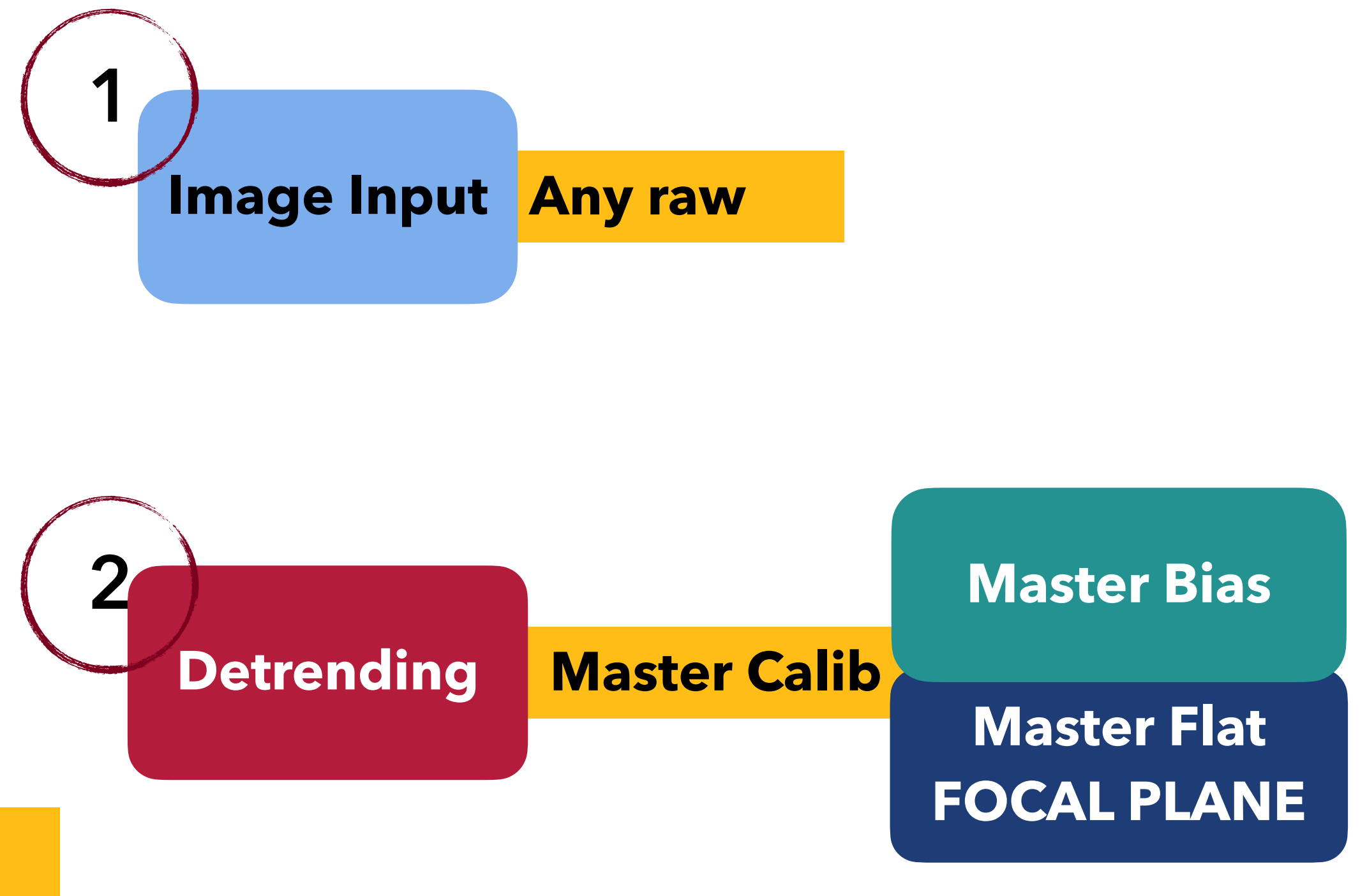
```
ztfimg.RawCCD  
ztfimg.RawQuadrant
```

1. Open
2. Pocket correction (date > 20191022)
3. Corr for NL from linearity tables / rcid (**date dependent**)
3. Corr for overscan :
 - Only last five columns used for model.
 - Model is cubic polynomial applied to sigma-clipped overscans

The *ztfin2p3* pipeline

2 Detrending raw images

- Image I/O and handling (e.g WCS) : `ztfim`
- Calibration `ztfin2p3`



Master Calib

Image I/O

Master Bias

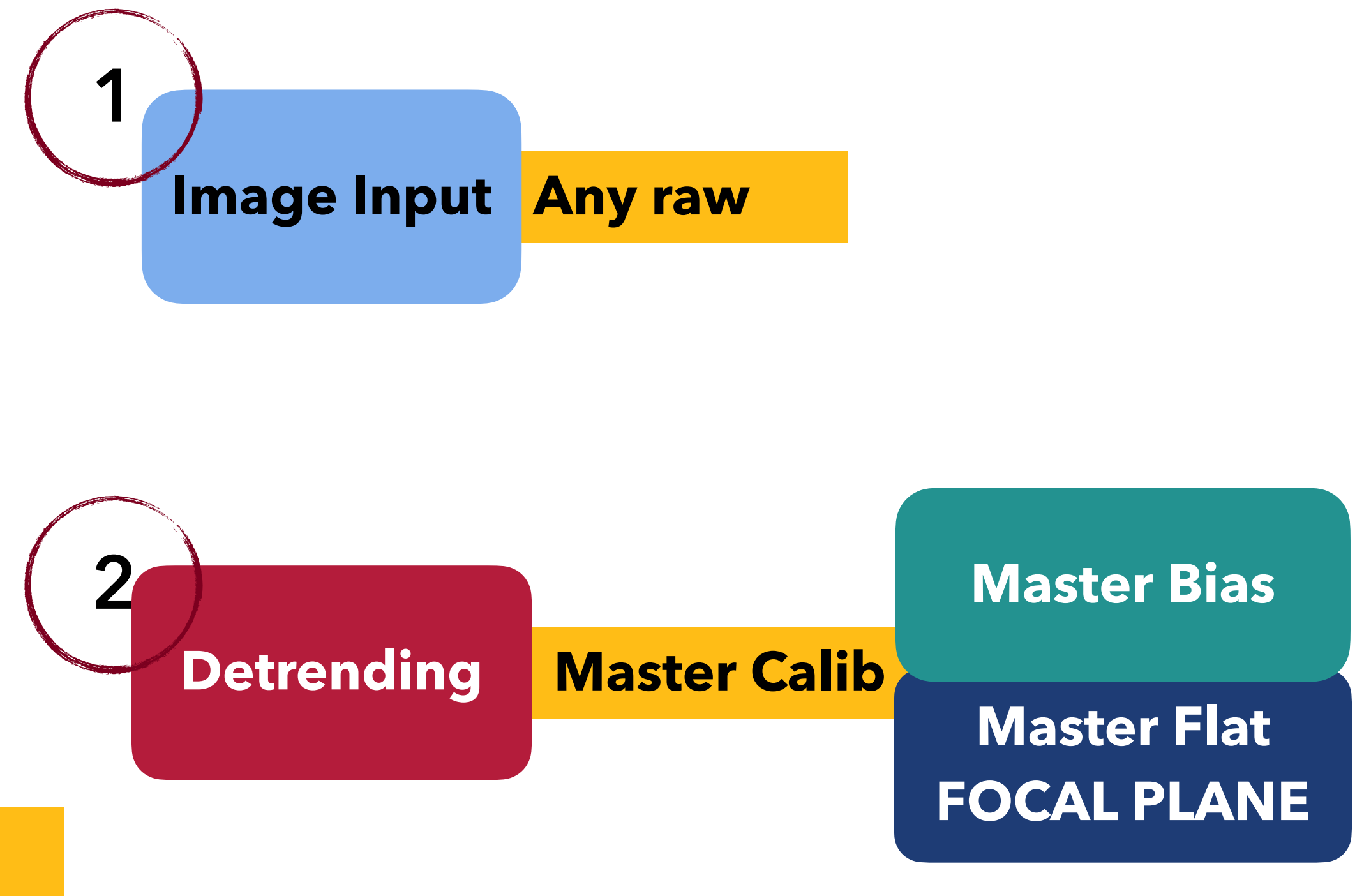
1. Use last 10 bias / day
2. Pixel-level sigma-clipped (3) median
 - **Store (with Header → package versions)**
 - **Image are store in fits files / CCD.**

```
ztfin2p3.pipe.newpipe  
class BiasPipe
```

The *ztfin2p3* pipeline

2 Detrending raw images

- Image I/O and handling (e.g WCS) : `ztfim`
- Calibration `ztfin2p3`



Master Calib

Master Flat

Image I/O

Master Bias

1. Pixel-level sigma-clipped (3) median / **LED**
2. Normalize flat / **LED** / **CDD** with **CCD Median**
3. Average CCD & norm to get **Flat / Filter**
4. Re-open everything to estimate **median level at Focal Plane level**
 - **Store in Header**
 - **Image are store as CCD normalized in fits files.**

```
ztfin2p3.pipe.newpipe  
class FlatPipe
```

The *ztf*in2p3 pipeline

2 Detrending raw images

- Image I/O and handling (e.g WCS) : `ztfimg`
- Calibration `ztf`in2p3

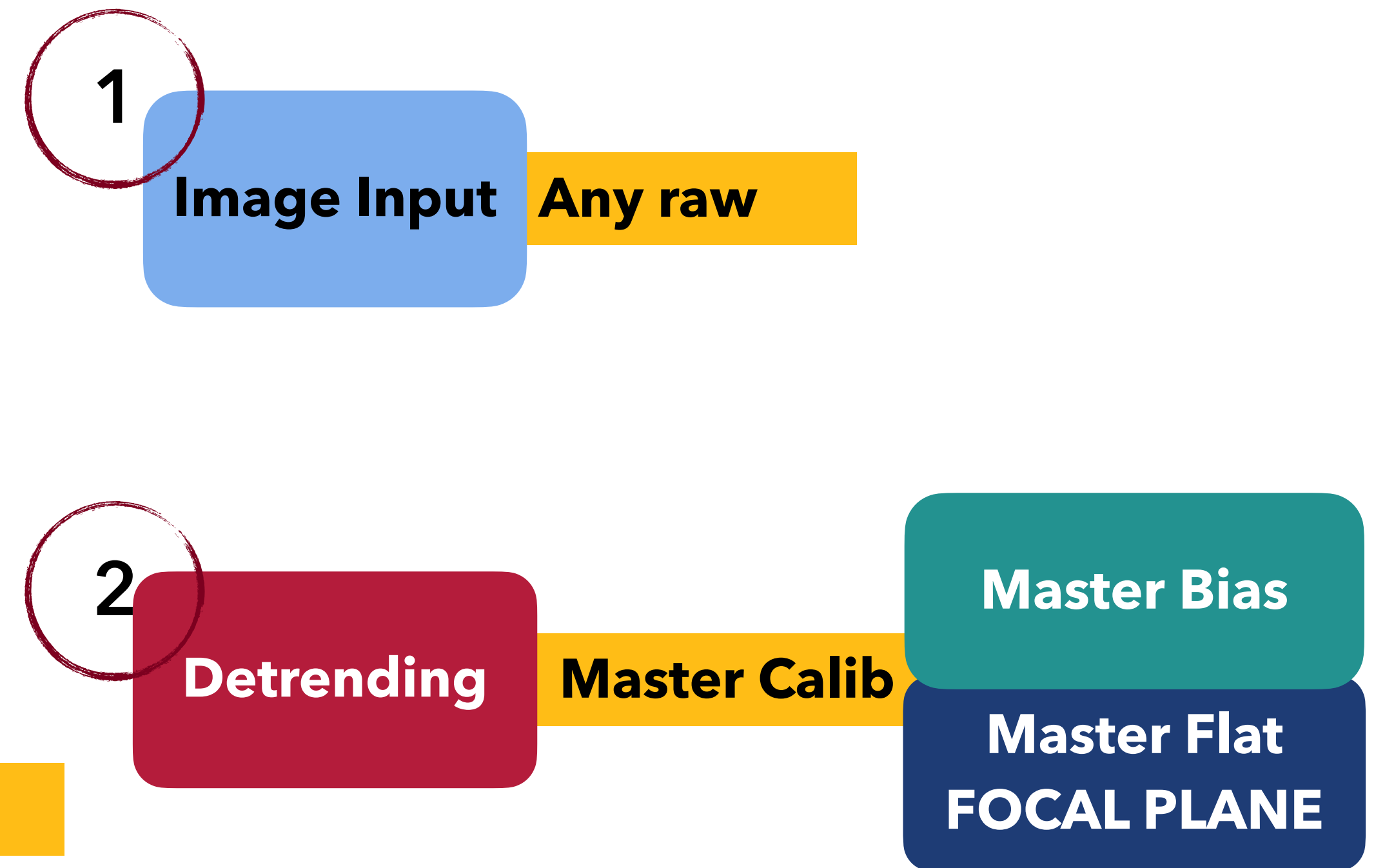


Image I/O	Detrending
Master Bias	<ol style="list-style-type: none">1. Get the nearest calib files → CCD, filter, date matching.2. Open raw file with <code>ztfimg</code> ¹3. Correct for bias ($I_{\text{raw}} - I_{\text{master bias}}$)4. Correct for flat at focal plane level $I_{\text{sci}} = \frac{I_{\text{bias corr}}}{I_{\text{Flat}}} \times \frac{N_{\text{fp}}}{N_{\text{ccd}}}$5. Separate I_{sci} into four quadrant → IPAC like sci format.6. Get associated headers & go from there
Master Flat FOCAL PLANE	

```
ztf
```

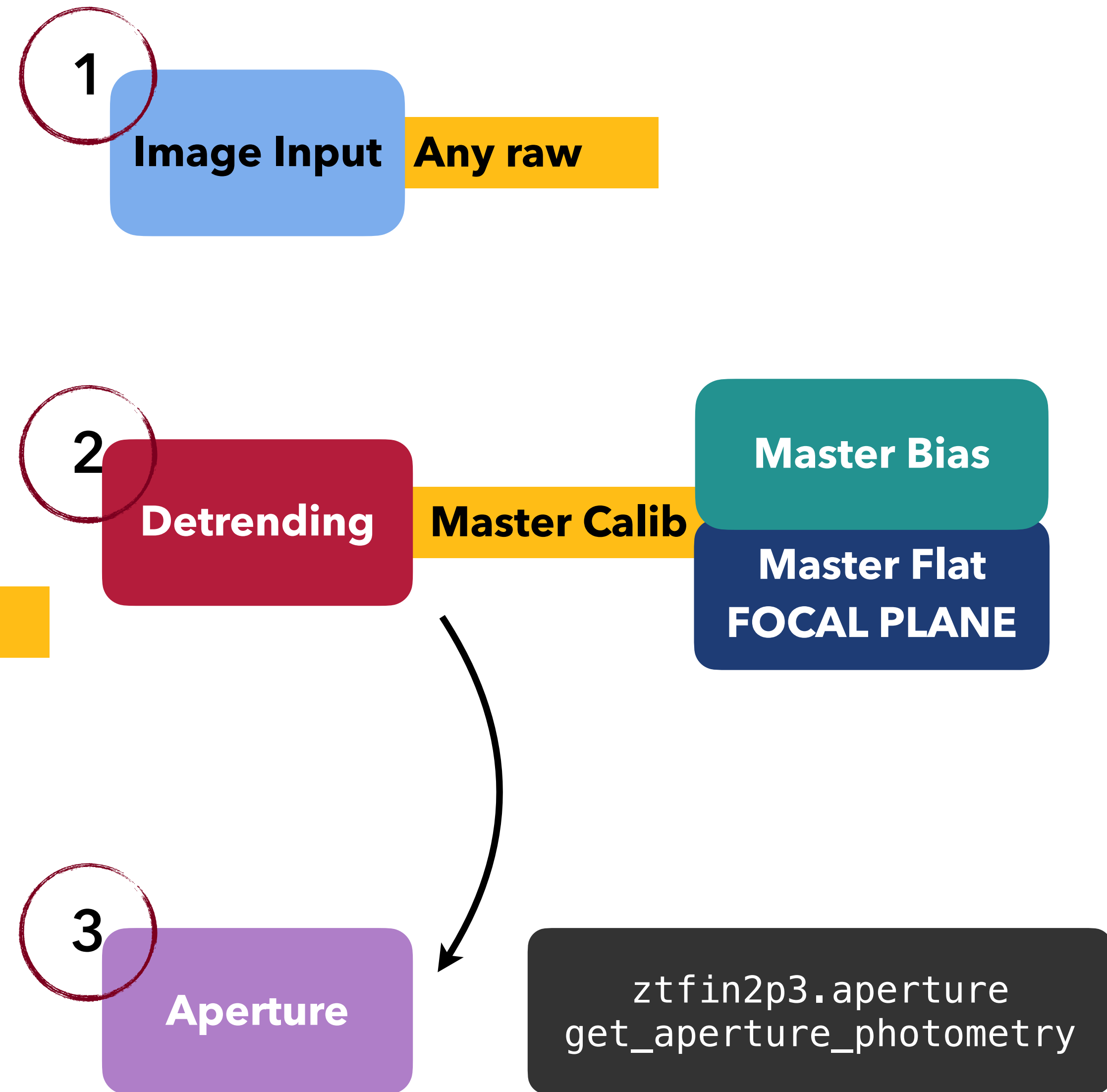
in2p3.science
build_science_image
build_science_header

```
#Output :  
ztfimg.ScienceQuadrant  
x4
```

The *ztfin2p3* pipeline

3 Aperture on detrended image

- Image I/O and handling (e.g WCS) : `ztfim`
- Calibration `ztfin2p3`



Master Calib

Master Bias

Master Flat
FOCAL PLANE

Image I/O

Aperture (Forced Photometry)

Detrending

1. Match Image to source catalogue to each quad : **GAIA DR3 - Ben ver.**
2. R.A , Dec to ra,dec with pec-vel
3. → (x, y) quadrant coord.
4. `sep` package aperture
5. Format and store to parquet

Aperture

```
ztfin2p3.aperture  
get_aperture_photometry
```

Aspects to note

→ Masks

→ Headers

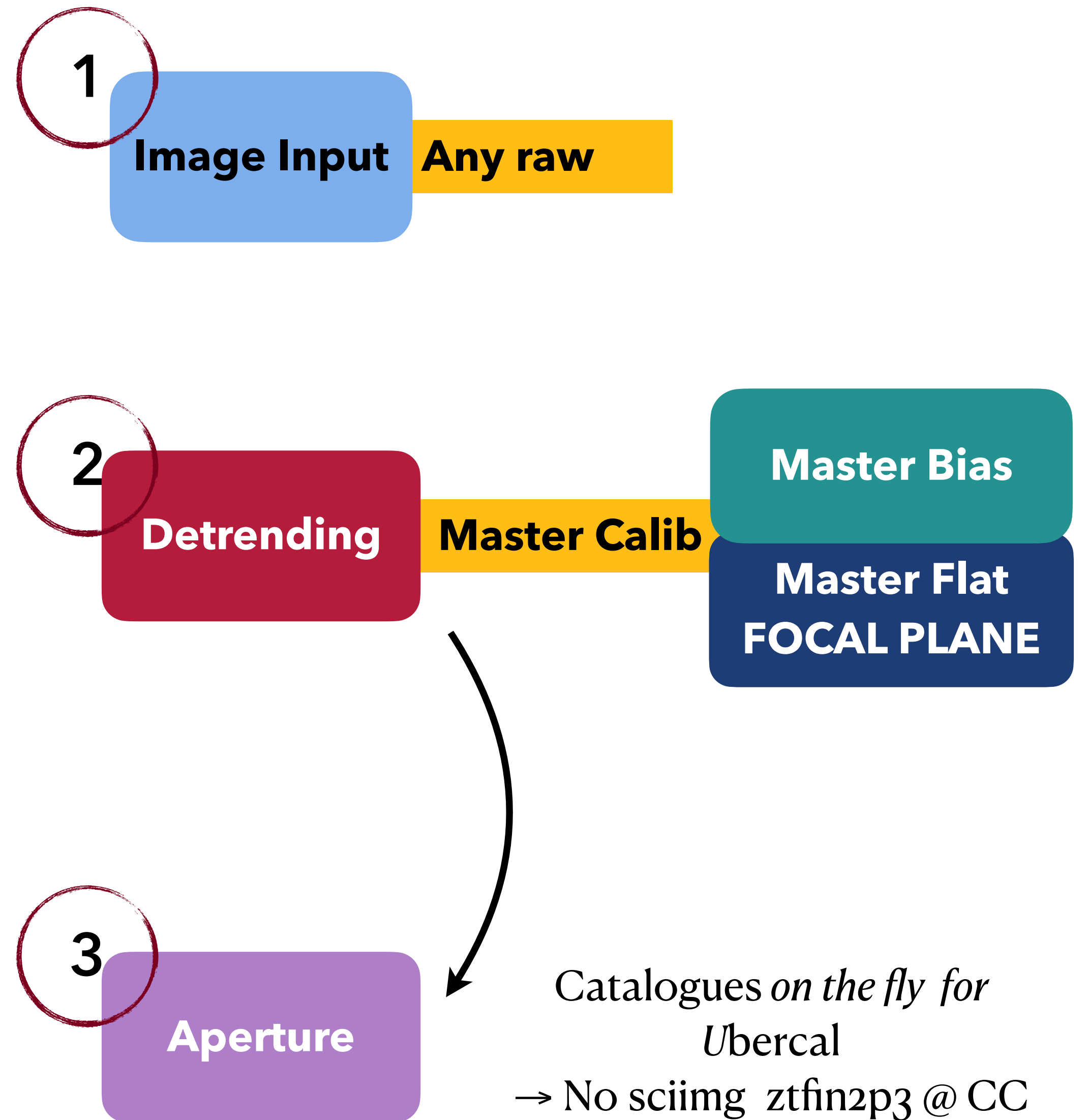
NOT REPROCESSED.
(Using ZTF IPAC masks & header)

→ Science images

NOT STORED.
Plan is : On the fly computations.

→ Cal + Raw

STORED ON DISK :
Easy to create re-create image if needed.



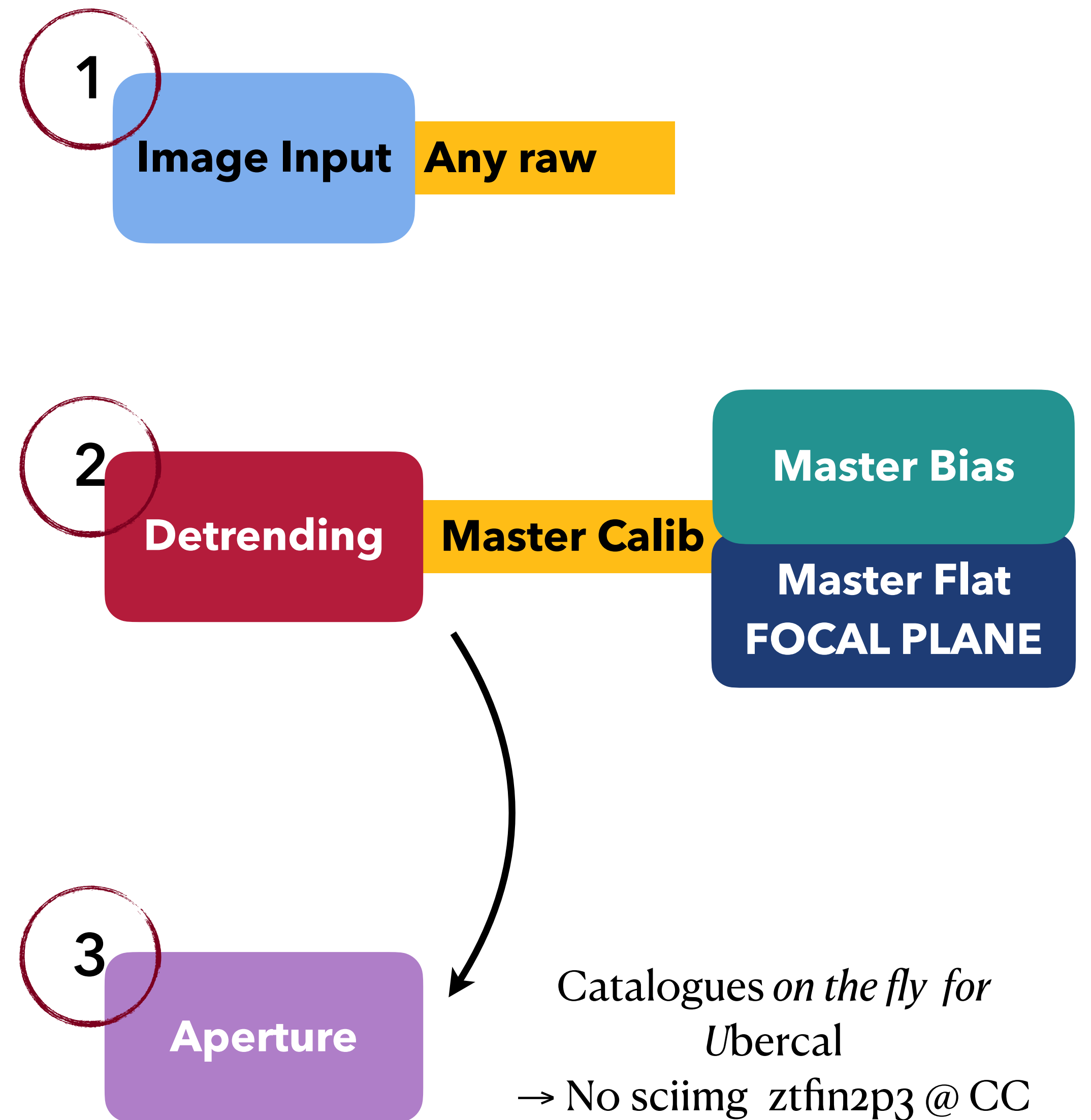
Pipeline status

« Ready to run » : Implementation aspects are coded.

Right now the main work is **the data** :

→ *Downloading from IPAC which is in progress.*

Data	Status
Headers (2018 - 2023)	✓ All good
Masks (2018 - 2023)	↻ Downloading
Raws sci (2018 - 2023)	↻ Downloading
Raws flats & bias (2018 - 2023)	✓ ↻ Some flats & biases are hiding on disk @ IRSA



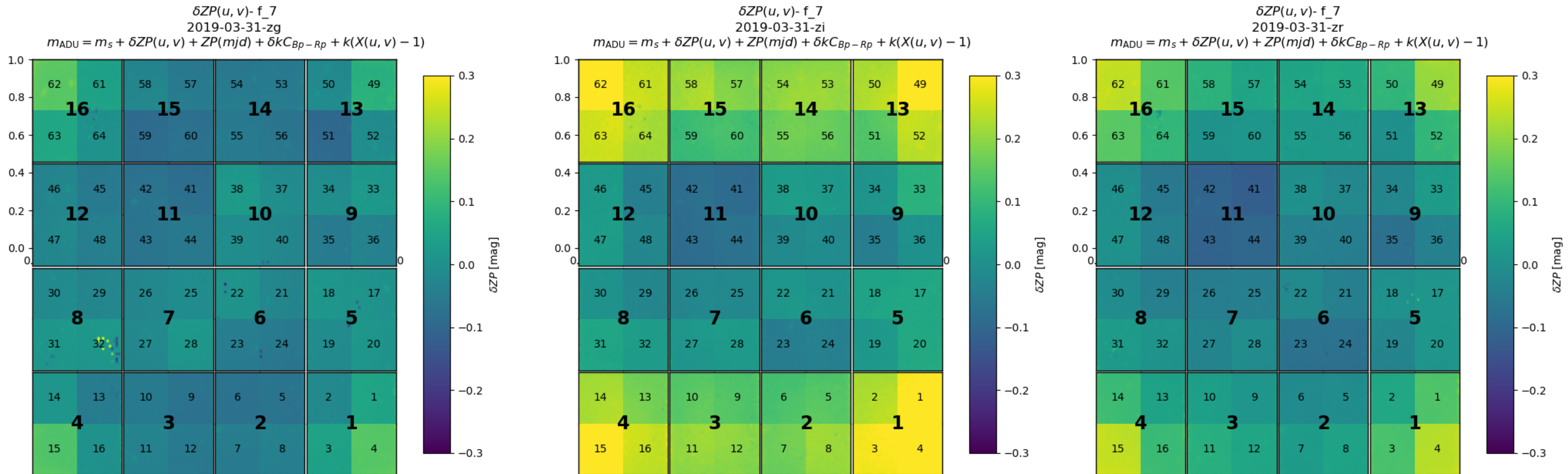
So we have a pipeline : What's next

✨Validation✨

- Pipeline will be tested both ways (in conjunction with Mathieu's work - just after) :
 - Starflats fits to aperture catalogues.
 - Ubercal from Ben.
- A **toy sample** has been devised by Ben on a small region of the sky between 2018 & 2023. → Dec : $[-5, 35]$ deg R.A : $[345, 65]$ deg in all ZTF bands.
(~1 million quadrants to process / band)
- Starflats : 3 starflats sequence were processed as of now (2018, 2019, 2021)

Example starflats 2019-03-31

Focal plane residuals with Quadrants normalization



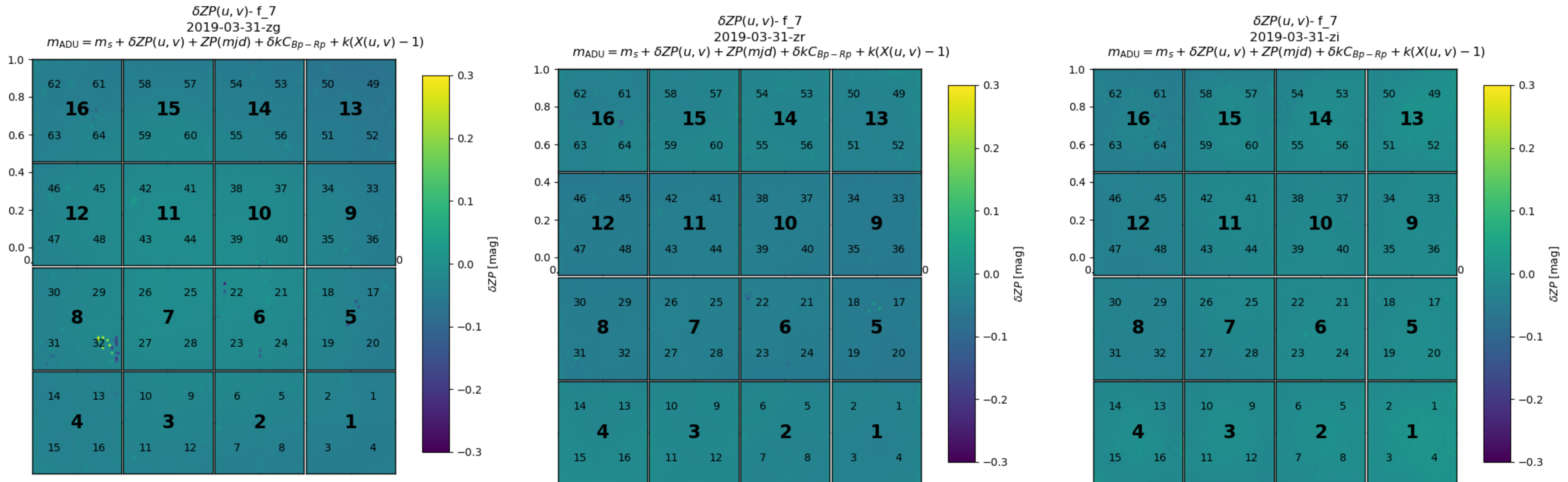
zg

zr

zi

Pipeline validation — as it is

Focal plane residuals with Focal Plane normalization



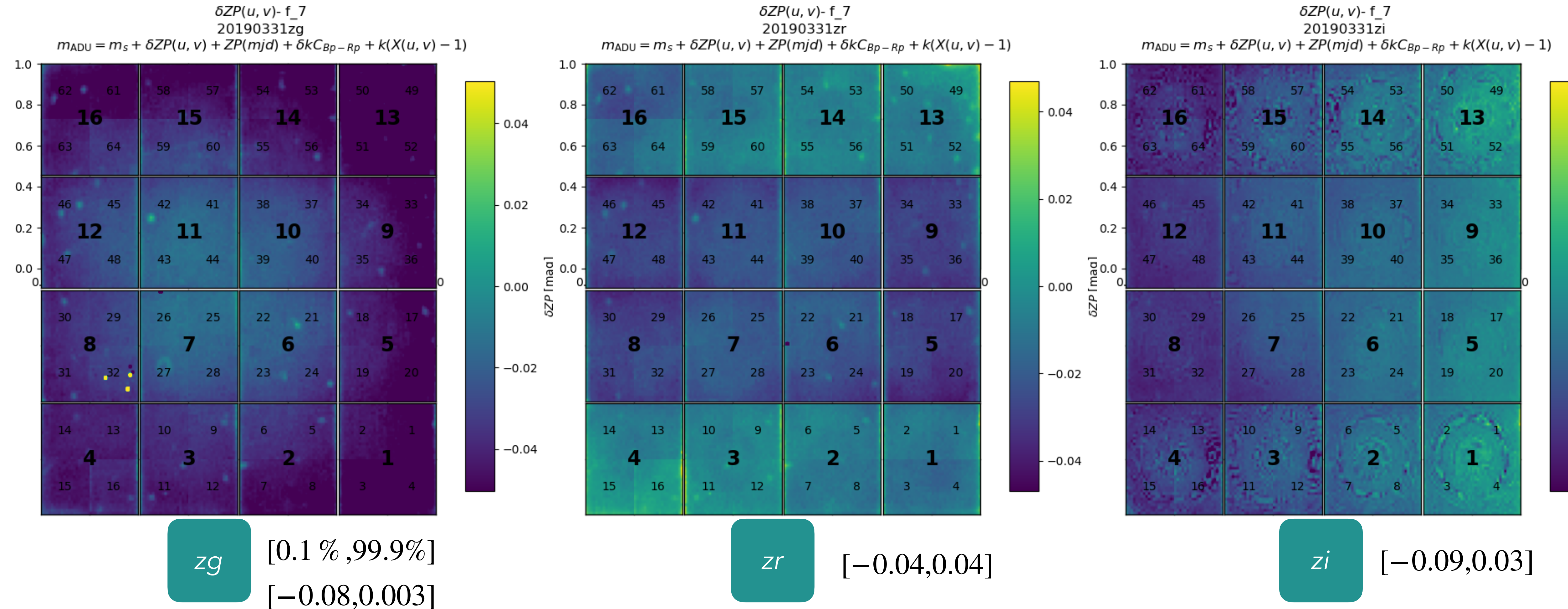
zg

zr

zi

Pipeline validation — as it is

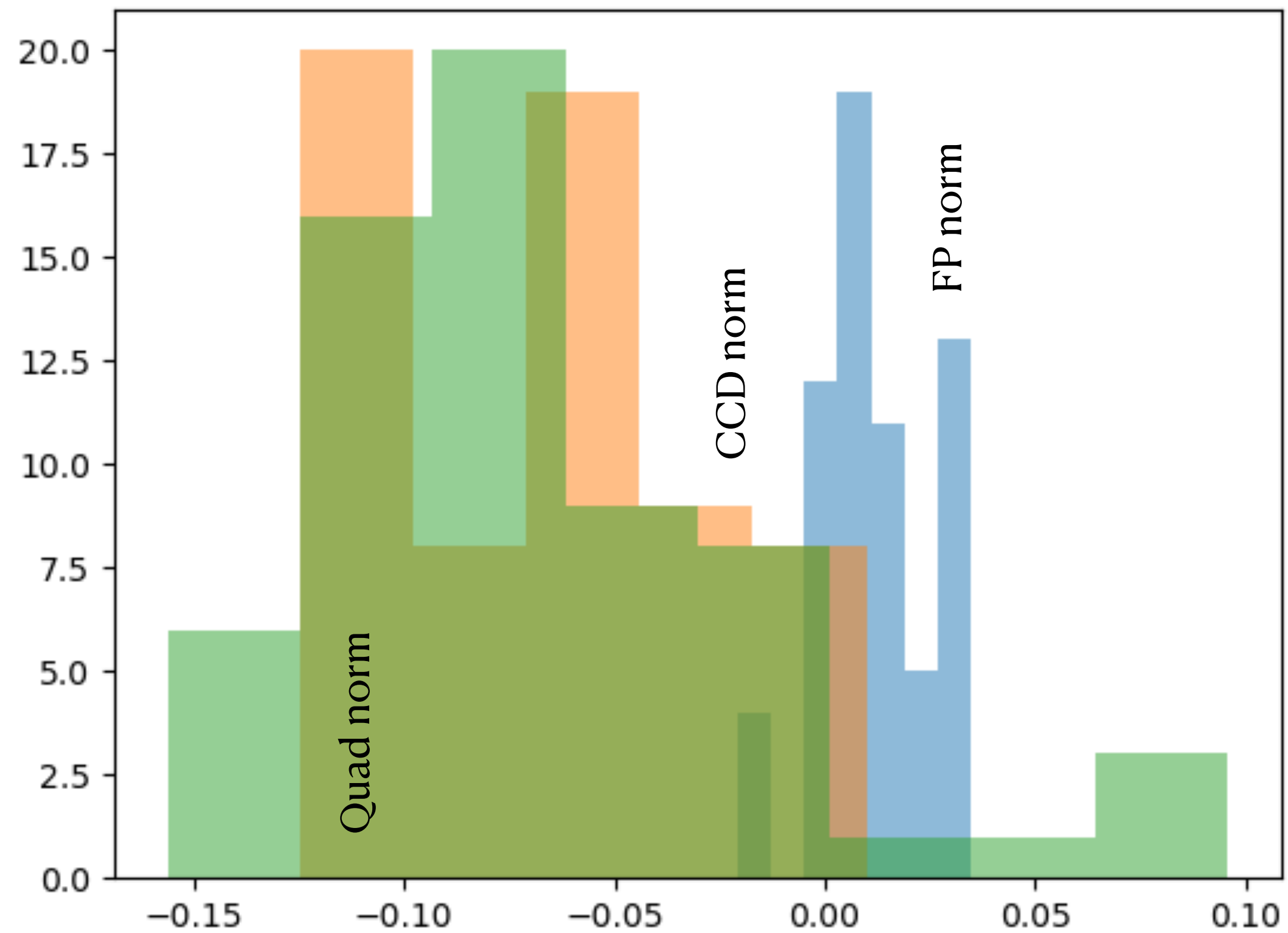
Focal plane residuals with Focal Plane normalization



Pipeline validation — as it is

Focal plane residuals with Focal Plane normalization

Fitted gain / quadrant on the starflat

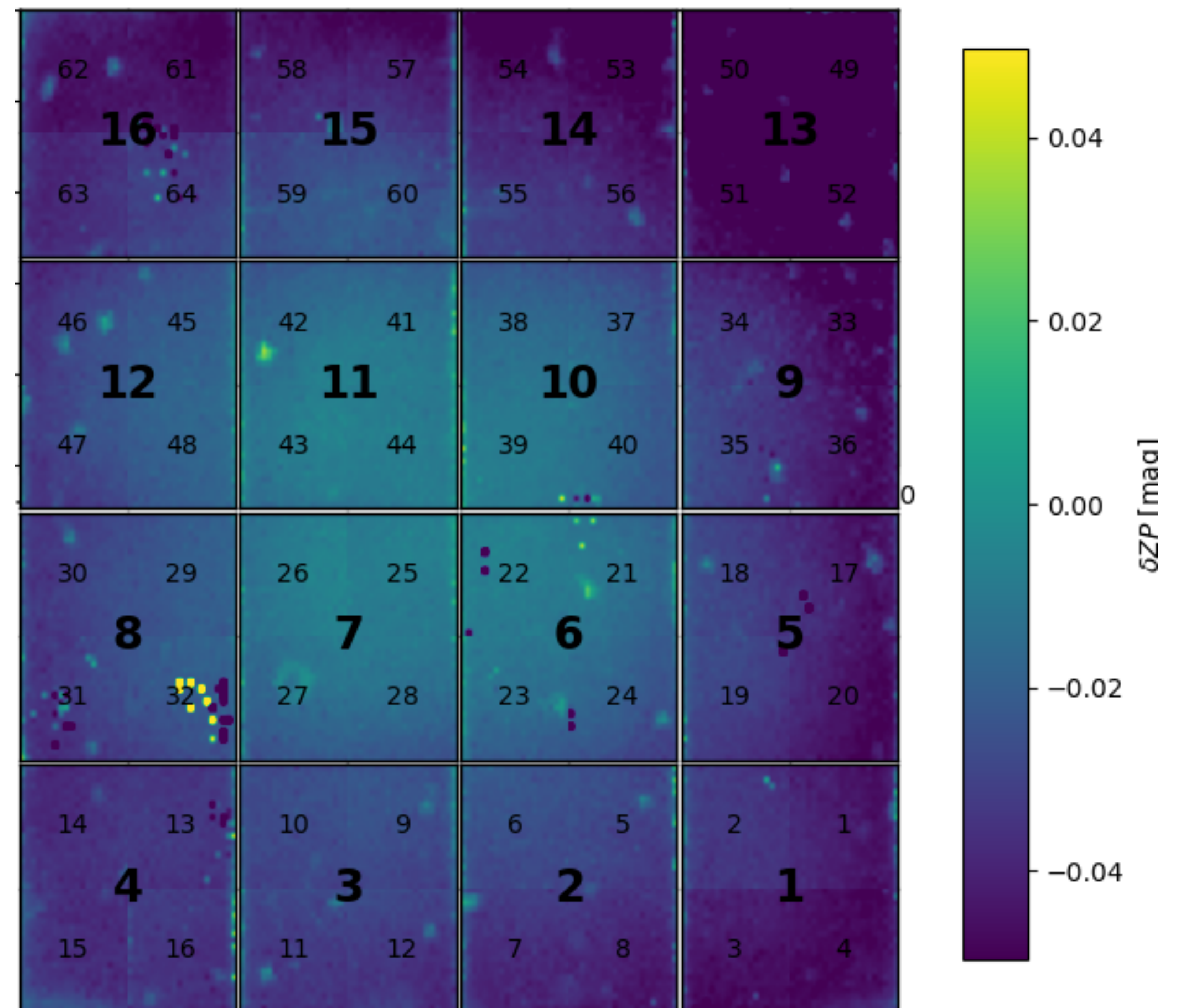


Fitted gain / Quadrant

$$\delta ZP(u, v) - f_7$$

2019-03-31-zg

$$m_{\text{ADU}} = m_s + \delta ZP(u, v) + ZP(\text{mjd}) + \delta k C_{Bp-Rp} + k(X(u, v) - 1)$$



ztf_{in}2p3 - Near future additions

ztf_{in}2p3 : what will this pipeline do

- 🔄 Images from 2020 & on were corrected for fringe effects in the i band.
Plan is to include this correction in the pipeline & correct the images prior with that.
- 🔄 Pocket effect correction (appeared in images > 2019-10-22)
→ 1st correction is implemented but needs further validations for some special cases.

ztf_{in}2p₃ - Pipeline

ztf_{in}2p₃ : what does this pipeline do.

✓ The ztf_{in}2p₃ pipeline is ready to run → Modified calibration

↻ → Some minor additions to be done.

✓ First results on the starflats sequence are promising :

→ Uniform focal plane

→ Quadrants variations are mitigated

↻ Pipeline validation on a larger scale (toy sample)

→ Dec : $[-5, 35]$ deg R.A : $[345, 65]$ deg

→ 2018 - 2023 to cover g, r and i bands.

**To be plugged with the Scene
Modelling pipeline**

**→ PSF & Aperture photometry
(not forced photometry)**

Mathieu's talk : Progress on the processing