

iP.2

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Fast python package to simulate *targets* observed by *surveys* 

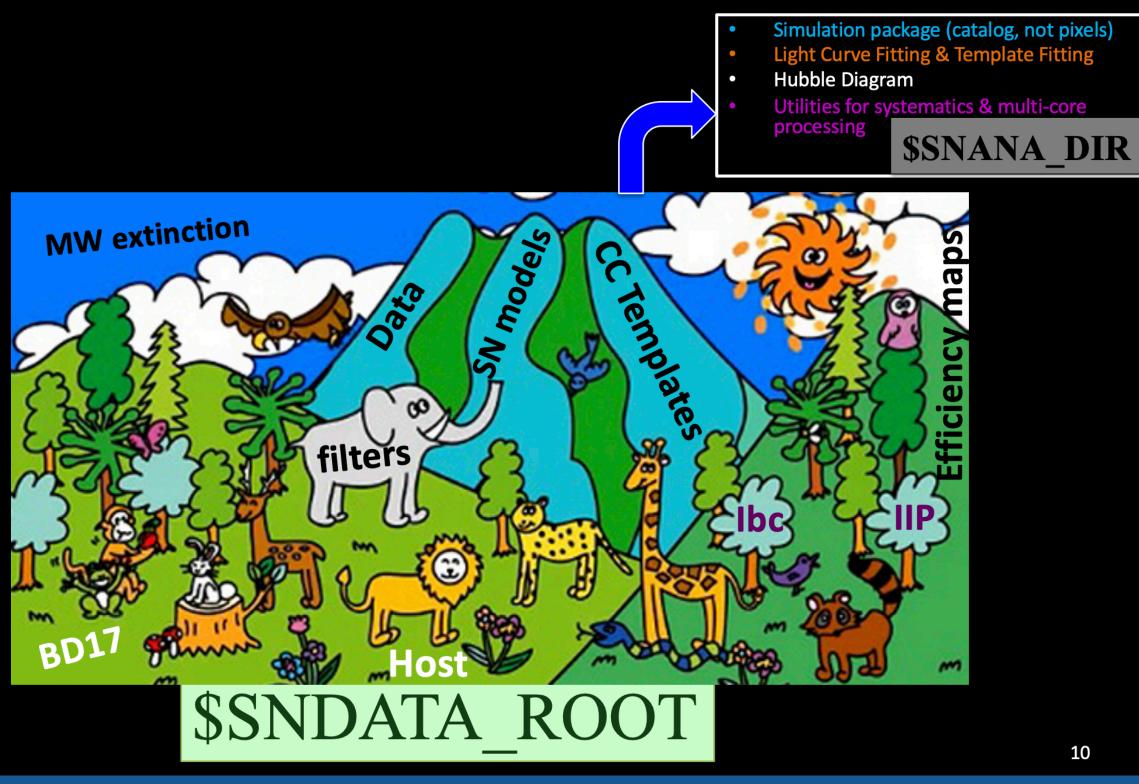
# & see also pip install modeldag

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### Source: SNANA tutorial

# Architecture: Environment



# **Simulation: SNANA**

Became the central tools SN cosmo:

**Simulate Transient** 

Simulate effect of selection function

**Correct for bias selection** 

**Fit cosmology** Scolnic+18; Brout+22; Riess+22; Popovic 2024; Vincenzi+24

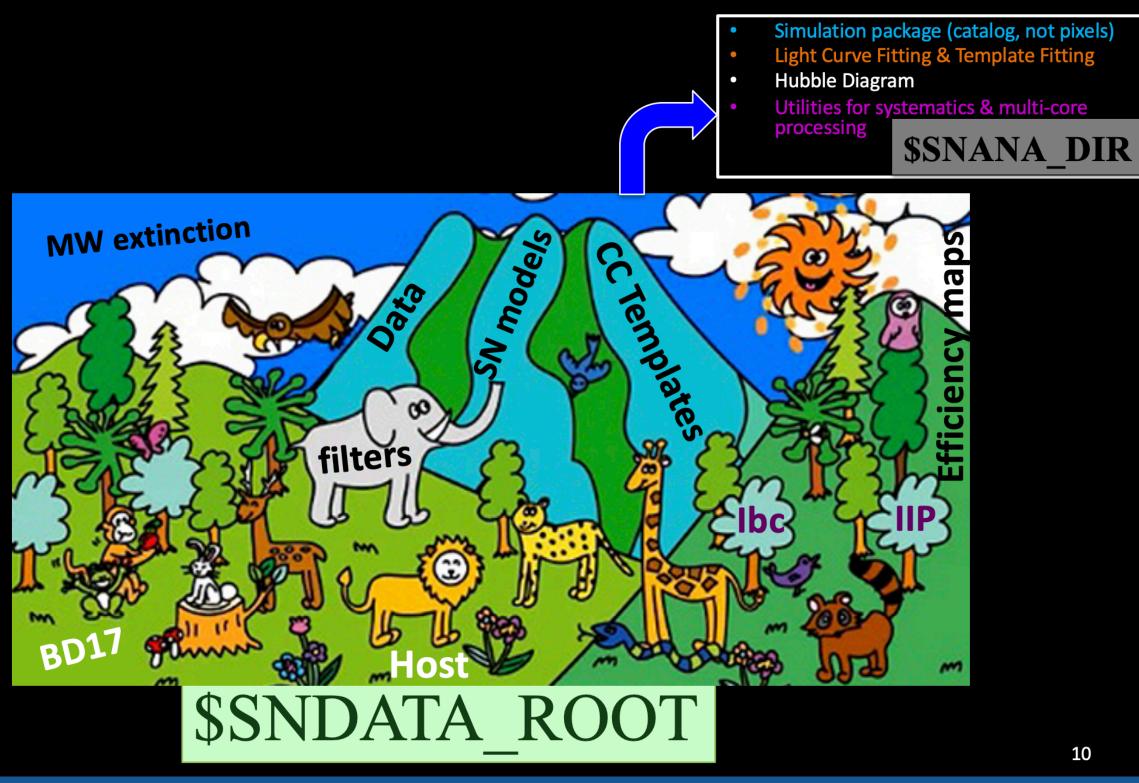
**Preparing:** 

LSST, Roman Plasstic / Elasticc



### Source: SNANA tutorial

# Architecture: Environment





# **Simulation: SNANA**

Became the central tools SN cosmo:

Simulate Transient

Simulate effect of selection function

**Correct for bias selection** 

**Fit cosmology** 

Scolnic+18; Brout+22; Riess+22; Vincenzi+24

SNANA is central for cosmological analyses... Yet, hard to install and complex to use.

pip install skysurvey





# targets

## Nature | An object in the sky



pip install **skysurvey** 

# skysurvey concept aims to replace SNANA "step 1: simulation"

## See also simsurvey | snsim



# dataset

The data you collect





# targets

## *Nature* | *An object in the sky*

import skysurvey

snia = skysurvey.SNeIa.from\_draw(10\_000, zmax=0.3, tstart = "2018-04-01",tstop = "2020 - 12 - 01")

### Lots of pre-built targets, *easy* to customize

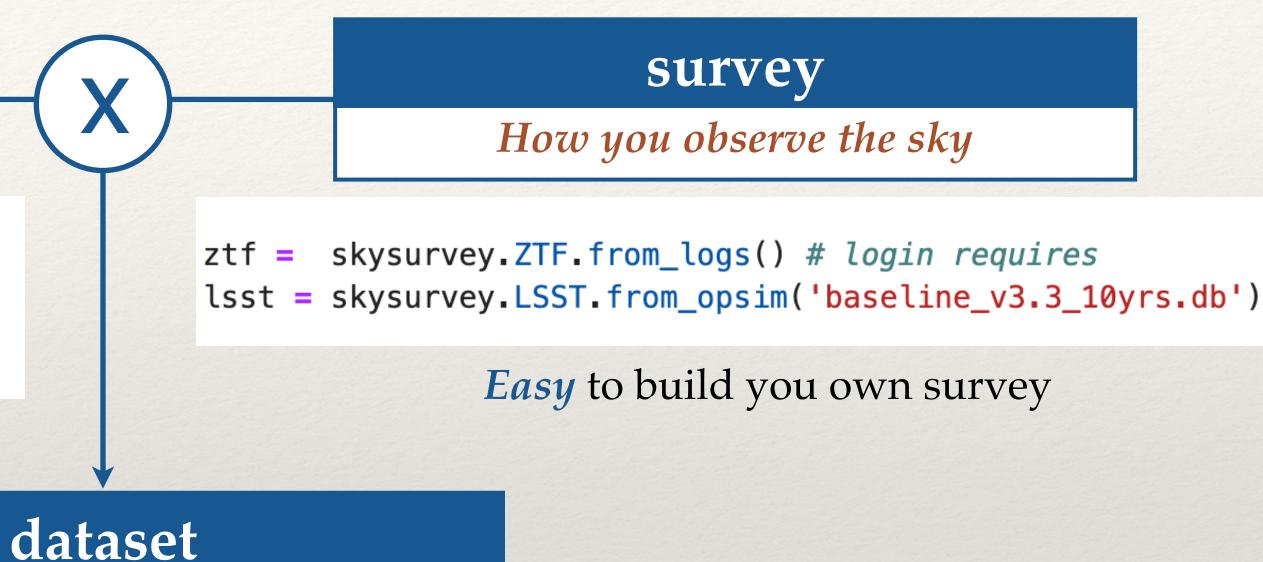


dset = skysurvey.DataSet.from\_targets\_and\_survey(snia, ztf)

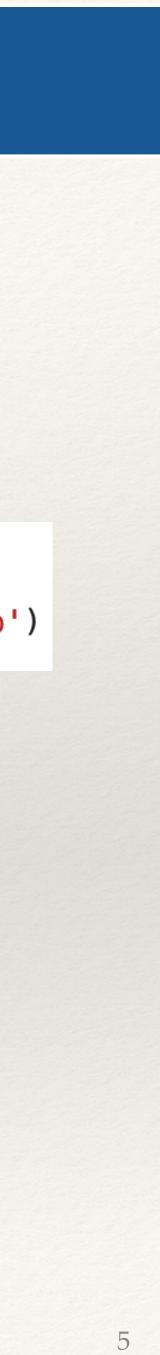
Always as *easy* as that

pip install skysurvey

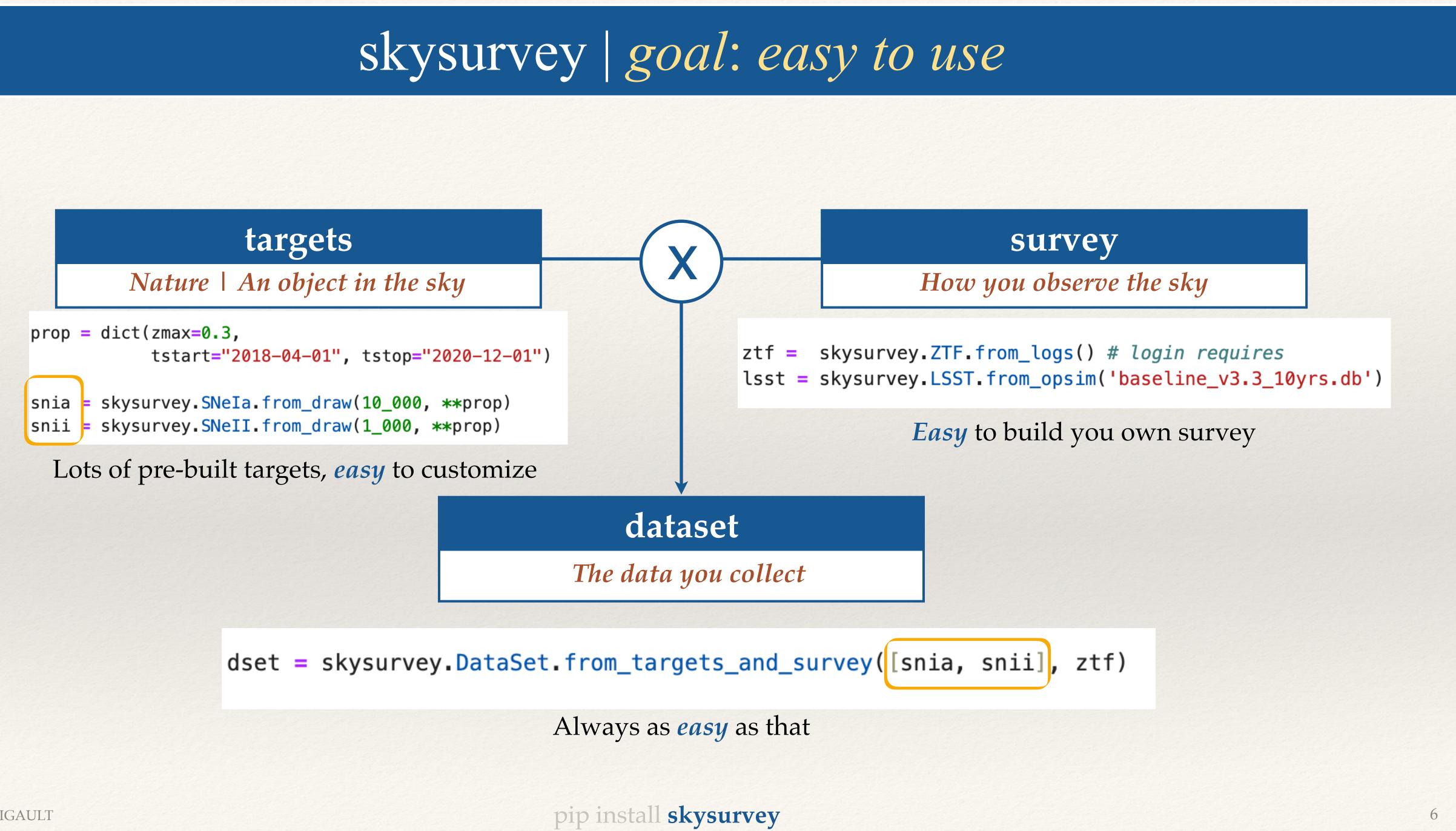
# skysurvey goal: easy to use



The data you collect







## skysurvey based within the python environment pandas SciPy <sup>S</sup> NumPy

# targets

*Nature* | *An object in the sky* 

snia.data

	z	x1	с	tO	ra	dec	magabs	mwebv	magobs	x0	template
0	0.07705	-0.030	1.221450	58487.214844	336.697937	7.605778	-15.358439	0.134286	22.429092	0.000017	salt2
1	0.06865	-0.700	-0.027583	58659.839844	60.706551	55.346775	-19.233326	0.719195	18.290983	0.000766	salt2
2	0.06905	0.320	0.351301	58782.136719	23.629887	-0.468616	-18.197527	0.036892	19.339998	0.000292	salt2

### dset.data *survey* .index fieldid rcid index index 5821 26009040 3 59104.39 3 59106.3 26054356 3 59112.3 26331019

## survey

*How you observe the sky* 

### survey.data

	gain	zp	skynoise	mjd	band	ra	dec	fieldid_sur
1867451	1	30	234.6674346923828	58800.0078125	desr	18.292482376098633	1.2328510284423828	11
1867554	1	30	234.6674346923828	58800.0078125	desr	18.292482376098633	1.2328510284423828	11
1867634	1	30	234.6674346923828	58800.0078125	desr	18.292482376098633	1.2328510284423828	11

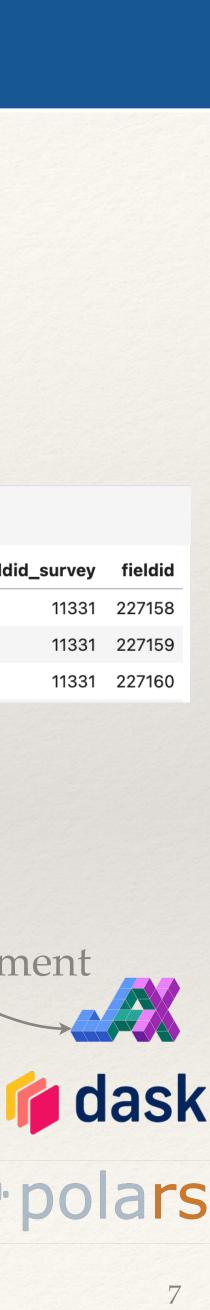
# dataset

### The data you collect

time	band	flux	fluxerr	zp	zpsys
390625	ztfr	-39.302373	36.143482	25.233383	ab
328125	ztfg	8.518026	14.717361	24.559206	ab
300781	ztfg	24.598649	20.259558	26.025703	ab

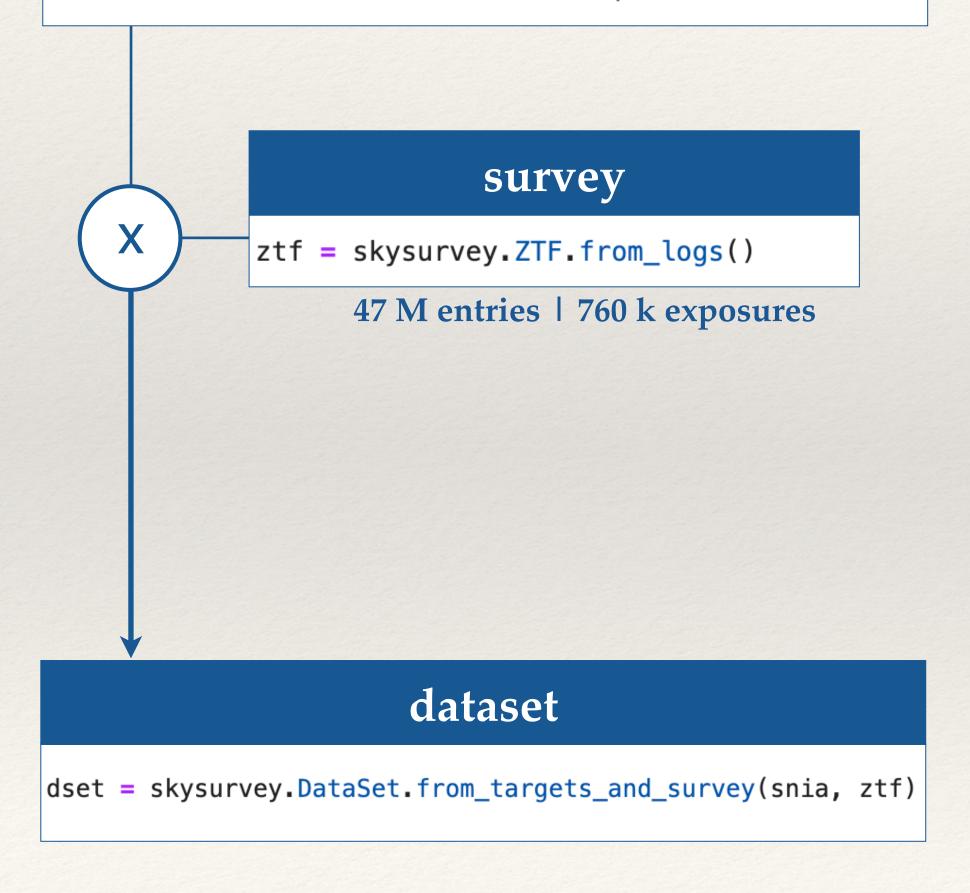
## In test development





# skysurvey | DataFrame

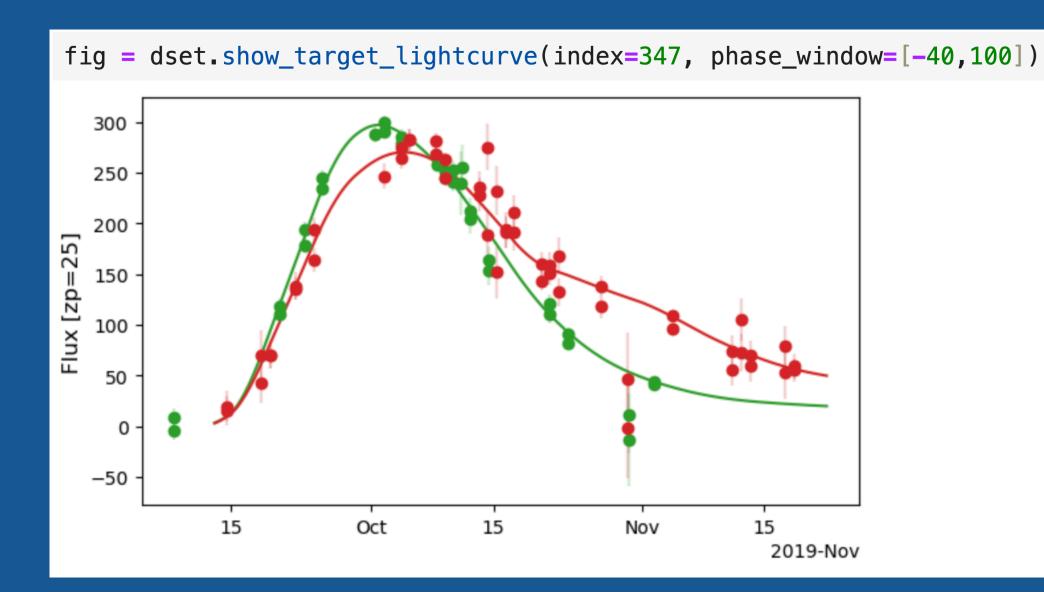


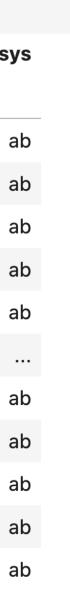


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dset	.data								
≯index	survey .index	fieldid	rcid	time	band	flux	fluxerr	zp	zpsy
5821	26009040	1	3	59104.390625	ztfr	-39.302373	36.143482	25.233383	a
	26054356	1	3	59106.328125	ztfg	8.518026	14.717361	24.559206	а
	26331019	1	3	59112.300781	ztfg	24.598649	20.259558	26.025703	а
	26338533	1	3	59112.359375	ztfr	-3.302169	36.260182	26.127666	а
	27121580	1	3	59128.250000	ztfr	-20.931607	46.604773	25.959711	a
	•••								
9727	30211656	1408	16	59203.179688	ztfr	14.887892	60.705860	26.204699	а
	30211715	1408	16	59203.179688	ztfr	90.188222	61.295684	26.208666	а
	30841192	1408	16	59221.167969	ztfr	34.804501	22.752532	26.142845	a
	36571545	1408	16	59392.414062	ztfr	-9.636968	41.250546	26.149637	а
	36571604	1408	16	59392.414062	ztfr	-18.475494	44.445759	26.151228	а
	► index 5821	index         5821       26009040         26054356       26331019         263331019       26338533         27121580       27121580         9727       30211656         30211715       30841192         36571545       36571545	<ul> <li>Survey</li> <li>index</li> <li< th=""><th>Survey         fieldid         rcid           5821         26009040         1         3           5821         26054356         1         3           26331019         11         3           26338533         1         3           27121580         1         3           9727         30211656         1408         16           30841192         1408         16           36571545         1408         16</th><th><ul> <li>Survey index inde</li></ul></th><th>Survey indexfieldidrcidtimeband5821260090401359104.390625ztfr260543561359106.328125ztfg263310191359112.300781ztfg263385331359112.359375ztfr271215801359128.250000ztfr97273021165614081659203.179688ztfr3084119214081659221.1679699ztfr3657154514081659392.414062ztfr</th><th>Survey indexfieldidrcidtimebandflux5821260090401359104.3906252tfr-39.302373260543561359106.3281252tfg8.518026263310191359112.3007812tfg24.598649263385331359128.2500002tfr-3.302169271215801359128.2500002tfr-20.9316073021165614081659203.1796882tfr14.8878923021171514081659221.1679692tfr34.8045013657154514081659392.4140622tfr-9.6369688</th><th>Survey indexfieldidrcidtimebandfluxfluxerr5821260090401359104.390625ztfr-39.30237336.143482260543561359106.328125ztfg8.51802614.717361263385331359112.300781ztfg24.59864920.259558263385331359128.250000ztfr-3.30216936.260182271215801359128.250000ztfr-20.93160746.60477397273021165614081659203.179688ztfr14.88789260.7058603021171514081659203.179688ztfr90.18822261.2956843084119214081659221.167969ztfr34.80450122.7525323657154514081659392.414062ztfr-9.63696841.250546</th><th>Mirrory indexfieldidrcidtimebandfluxfluxerrzp58212600904011359104.3906252tfr-39.30237336.14348225.2333832605435611359106.3281252tfg8.51802614.71736124.5592062633101911359112.3007812tfg24.59864920.25955826.0257032633853311359112.3593752tfr-3.30216936.26018226.1276662712158011359128.2500002tfr-20.93160746.60477325.959711097273021165614081659203.1796882tfr14.88789260.70586026.2046993084119214081659203.1796882tfr91.8822261.29568426.2086663084119214081659203.1796882tfr34.80450122.7525226.1428453657154514081659221.1679692tfr34.80450122.7525226.142845</th></li<></ul>	Survey         fieldid         rcid           5821         26009040         1         3           5821         26054356         1         3           26331019         11         3           26338533         1         3           27121580         1         3           9727         30211656         1408         16           30841192         1408         16           36571545         1408         16	<ul> <li>Survey index inde</li></ul>	Survey indexfieldidrcidtimeband5821260090401359104.390625ztfr260543561359106.328125ztfg263310191359112.300781ztfg263385331359112.359375ztfr271215801359128.250000ztfr97273021165614081659203.179688ztfr3084119214081659221.1679699ztfr3657154514081659392.414062ztfr	Survey indexfieldidrcidtimebandflux5821260090401359104.3906252tfr-39.302373260543561359106.3281252tfg8.518026263310191359112.3007812tfg24.598649263385331359128.2500002tfr-3.302169271215801359128.2500002tfr-20.9316073021165614081659203.1796882tfr14.8878923021171514081659221.1679692tfr34.8045013657154514081659392.4140622tfr-9.6369688	Survey indexfieldidrcidtimebandfluxfluxerr5821260090401359104.390625ztfr-39.30237336.143482260543561359106.328125ztfg8.51802614.717361263385331359112.300781ztfg24.59864920.259558263385331359128.250000ztfr-3.30216936.260182271215801359128.250000ztfr-20.93160746.60477397273021165614081659203.179688ztfr14.88789260.7058603021171514081659203.179688ztfr90.18822261.2956843084119214081659221.167969ztfr34.80450122.7525323657154514081659392.414062ztfr-9.63696841.250546	Mirrory indexfieldidrcidtimebandfluxfluxerrzp58212600904011359104.3906252tfr-39.30237336.14348225.2333832605435611359106.3281252tfg8.51802614.71736124.5592062633101911359112.3007812tfg24.59864920.25955826.0257032633853311359112.3593752tfr-3.30216936.26018226.1276662712158011359128.2500002tfr-20.93160746.60477325.959711097273021165614081659203.1796882tfr14.88789260.70586026.2046993084119214081659203.1796882tfr91.8822261.29568426.2086663084119214081659203.1796882tfr34.80450122.7525226.1428453657154514081659221.1679692tfr34.80450122.7525226.142845

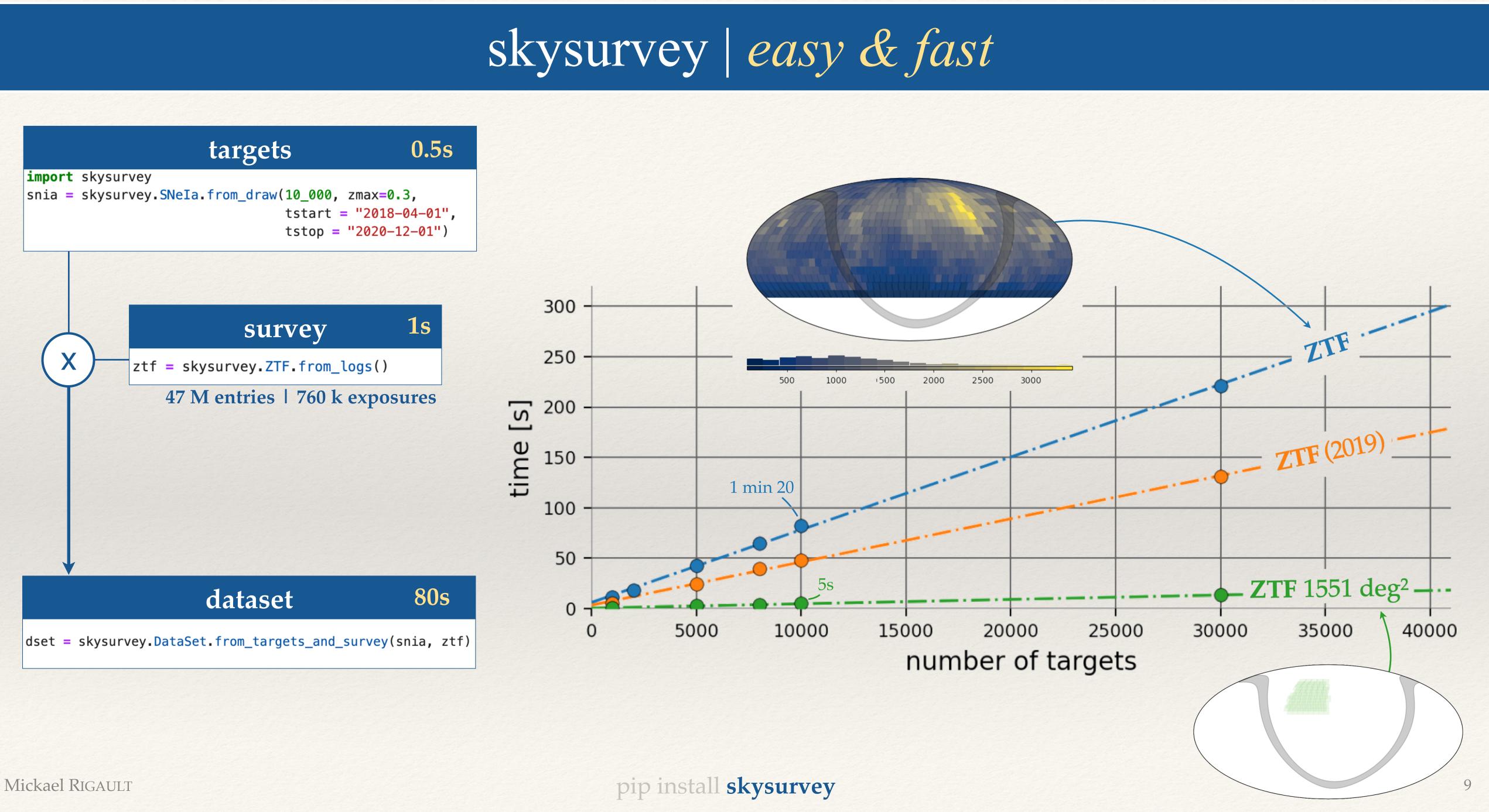
<sup>8429351</sup> rows × 8 columns



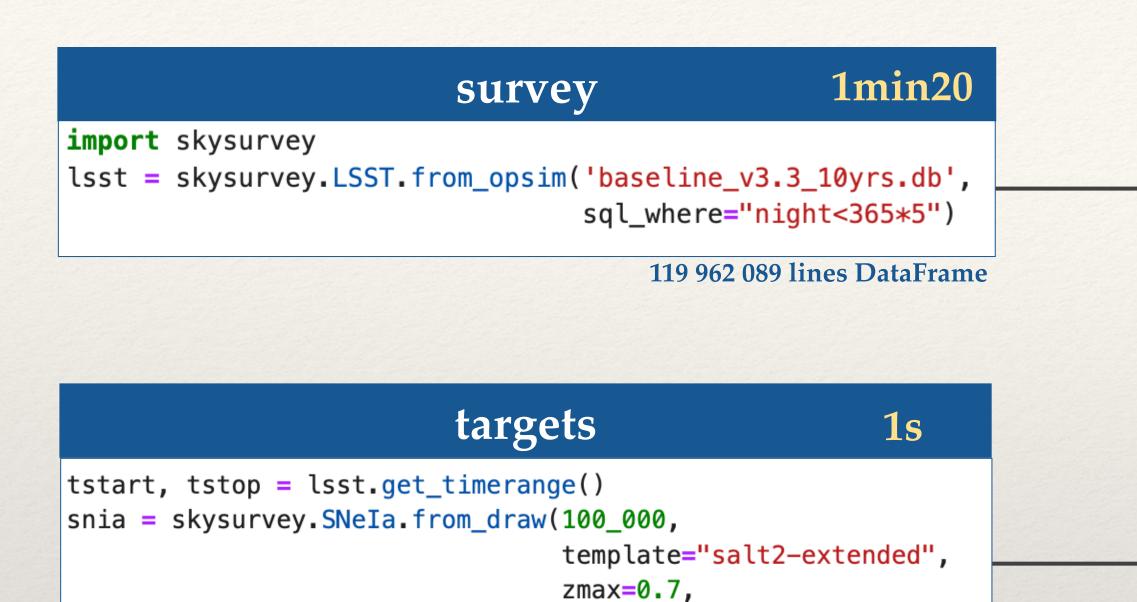












tstart=tstart, tstop=tstop)

pip install **skysurvey** 



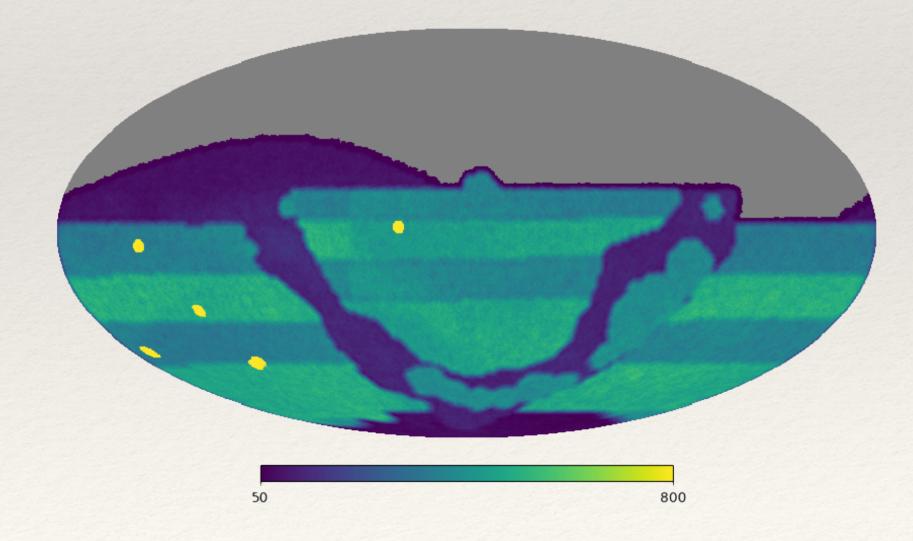






dset = skysurvey.DataSet.from\_targets\_and\_survey(snia, lsst)

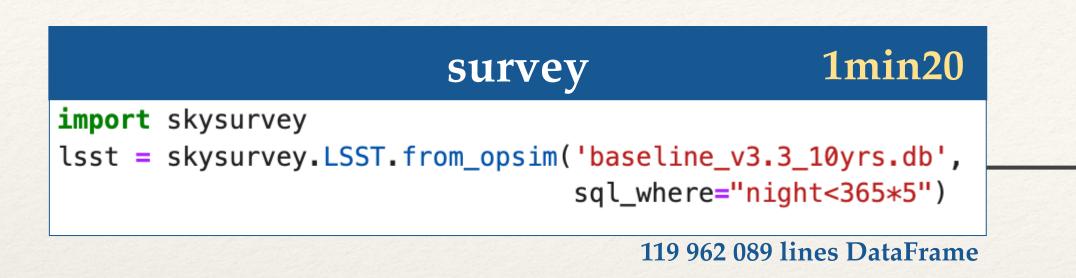
## 100 000 SNe Ia | 5 years of LSST | 3min

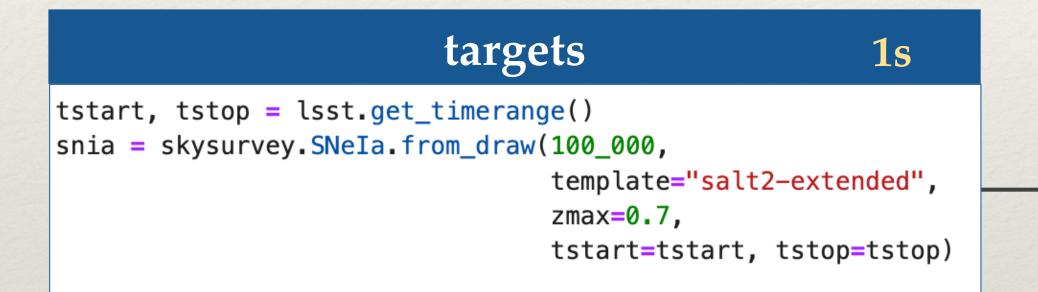






# skysurvey | easy & fast





Source: SNANA tutorial

## SIM CPU Proc-Time

U Chicago Research Computing Center: Sep 2018 for PLAsTiCC

• 117 million light curves generated in 8 hr on 40 cores  $\rightarrow$  100/sec

skysurvey: 450/sec | 1 laptop (M1)

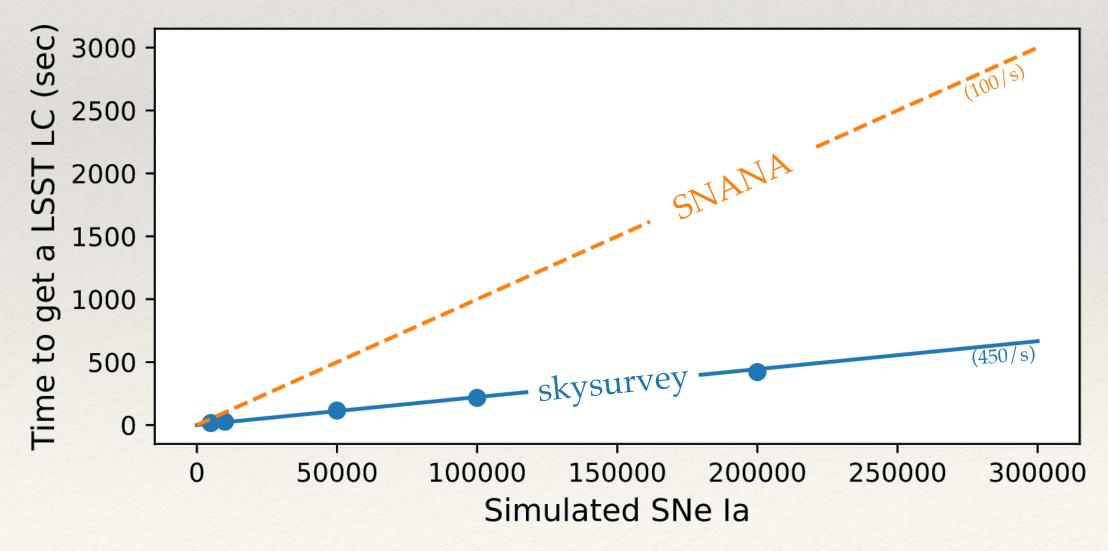
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pip install **skysurvey** 

## LSST opsim (dev)

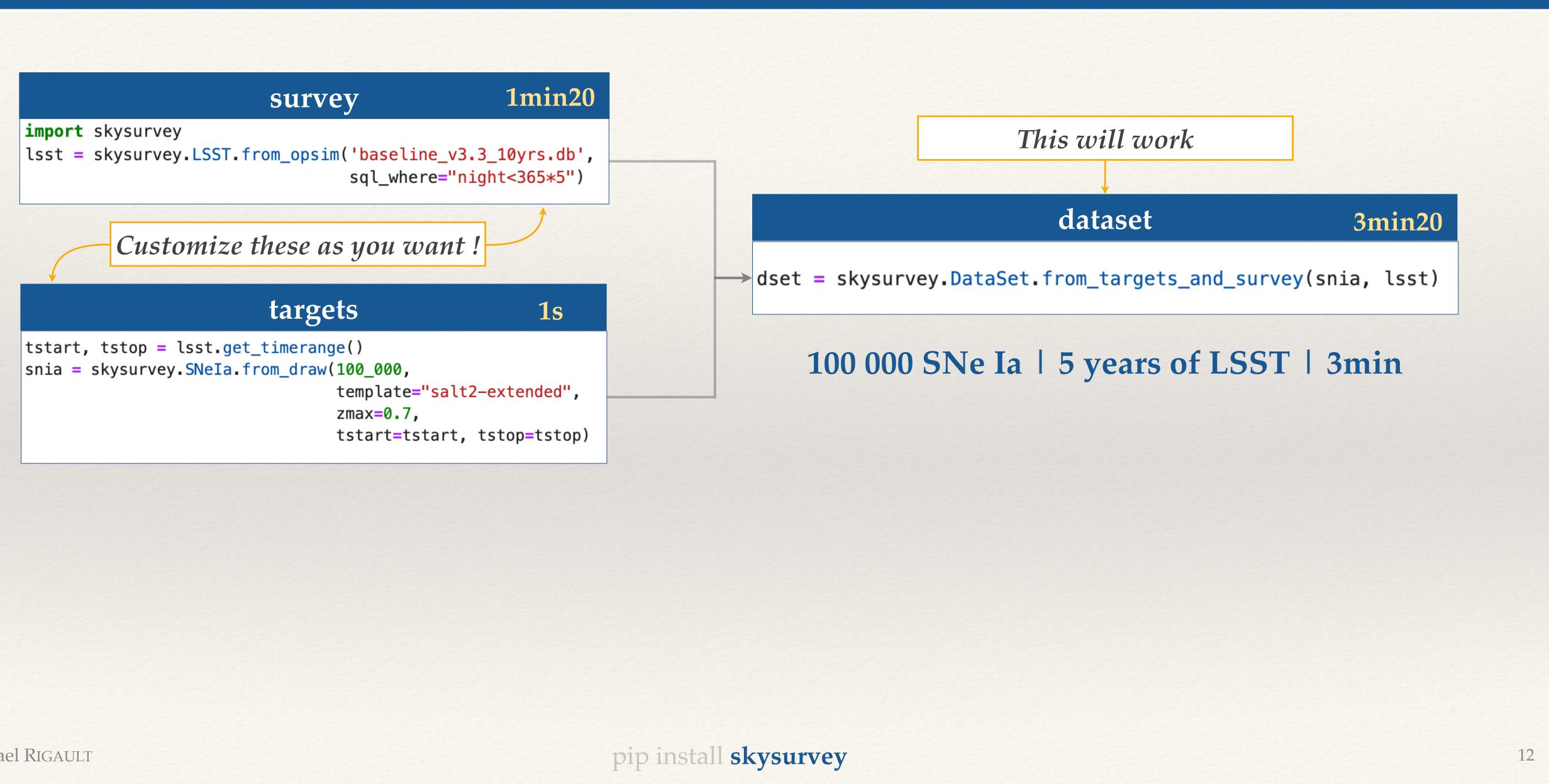


# 100 000 SNe Ia | 5 years of LSST | 3min







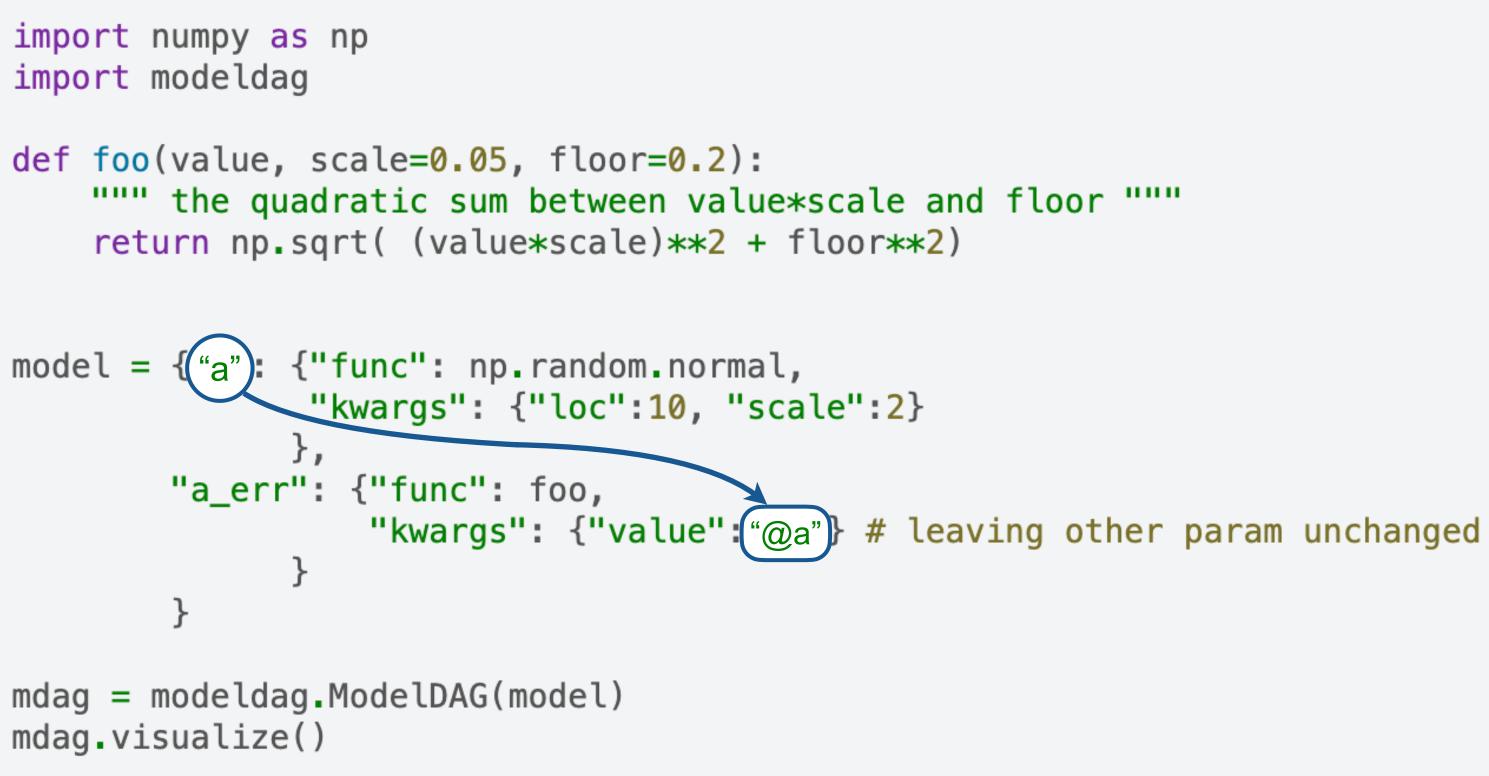




LSST opsim (dev)



# *Easy* to build *complex* transient



# DAG

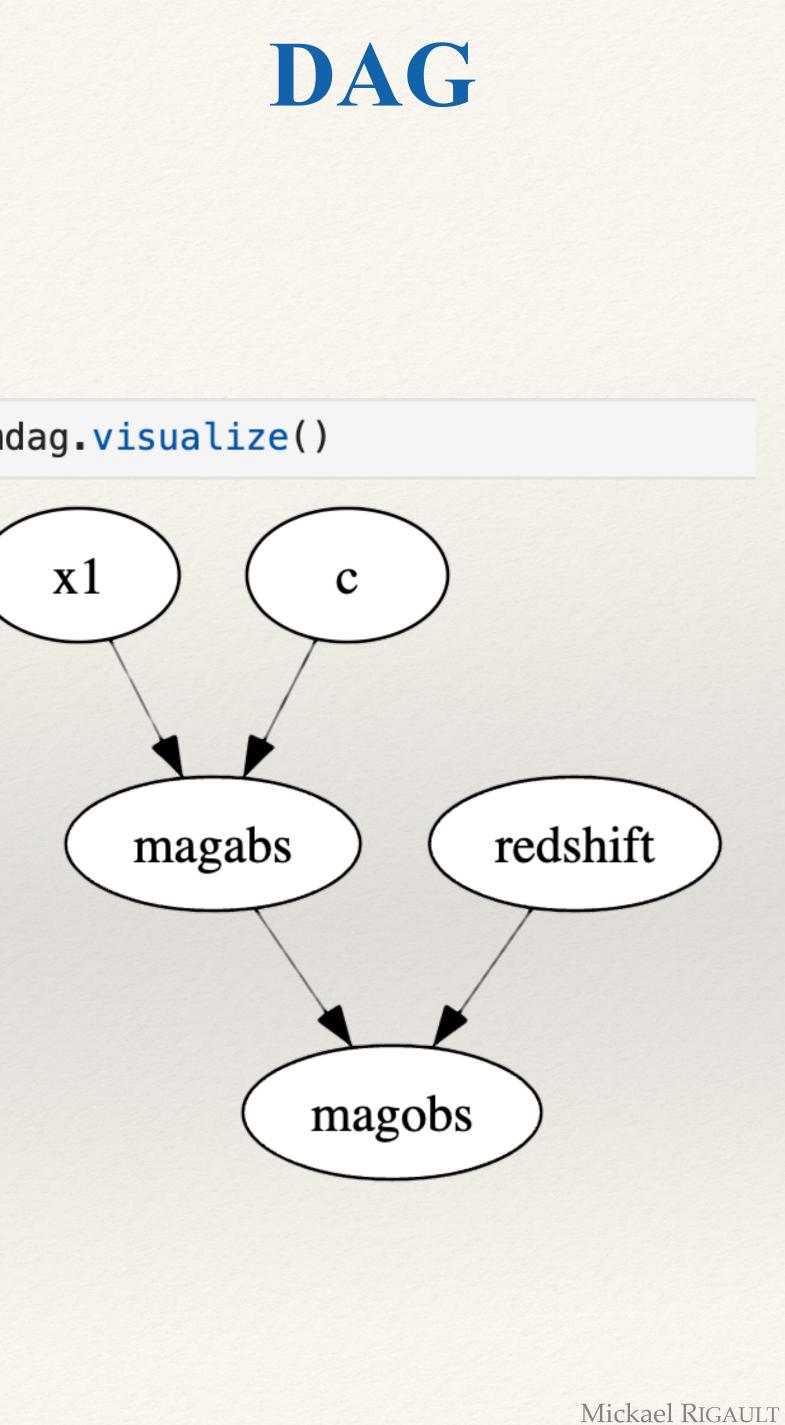
a a\_err

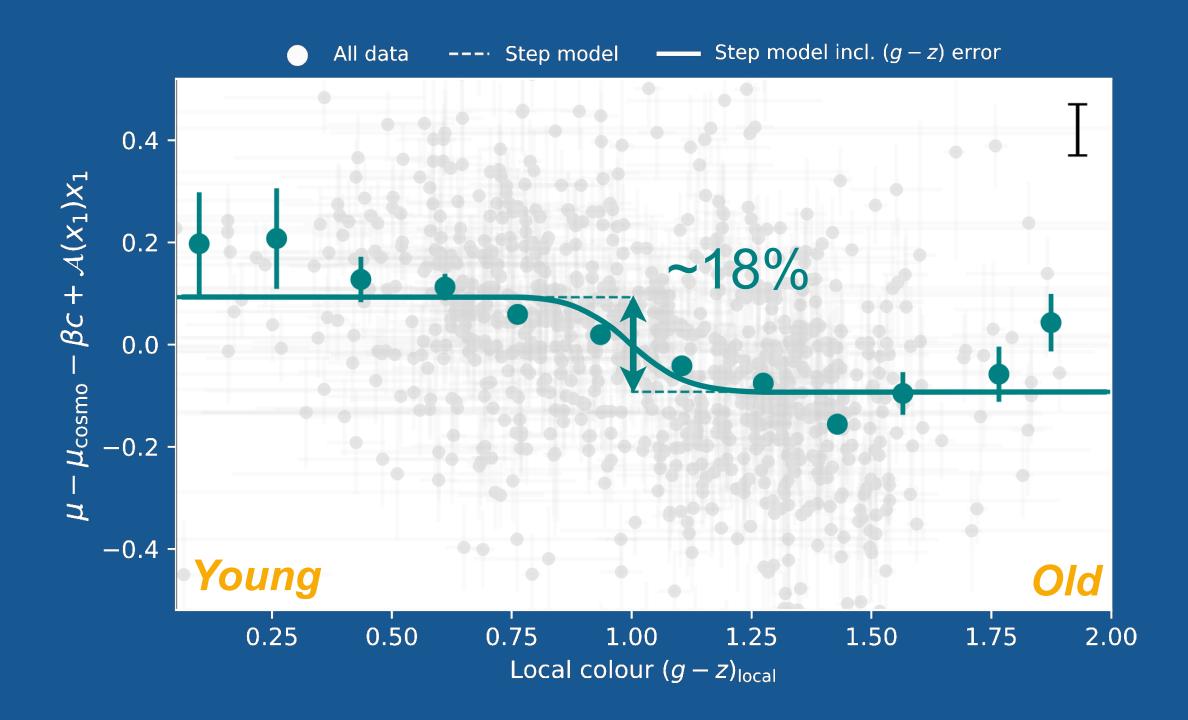


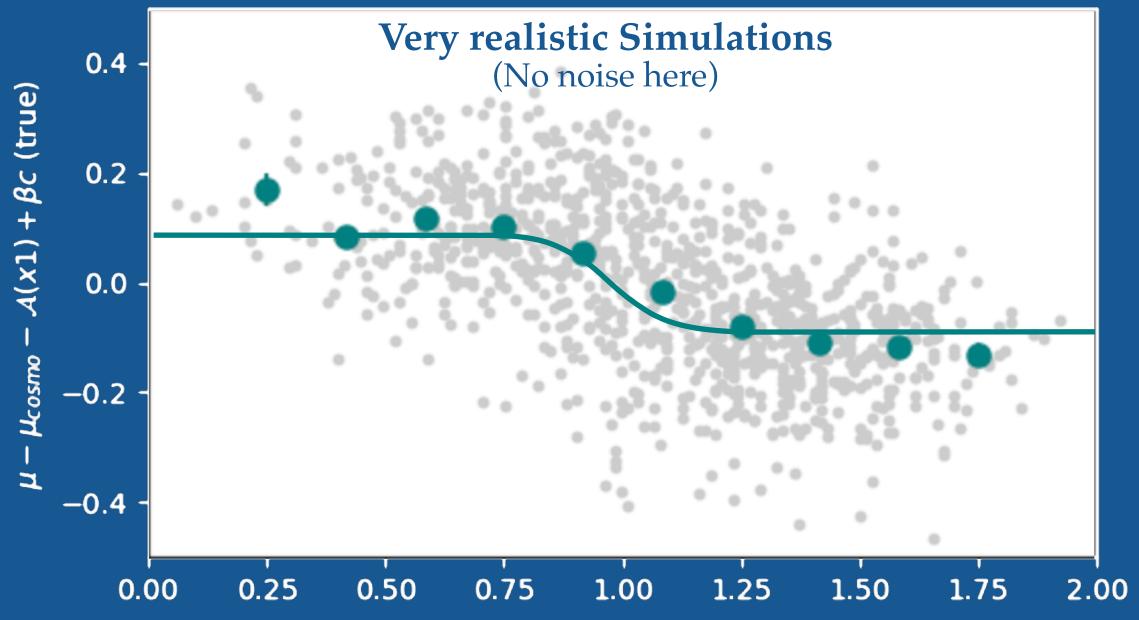
# *Easy* to build *complex* transient

```
_____
   Your functions #
#
     _____
def tripp98(x1, c, alpha=-0.14, beta=3.3, sigmaint=0.1, mabs=-19.3):
    """ This is the Tripp 98 relation """
    magabs = x1*alpha + c*beta + mabs # natural SN Ia magnitude (non-standardized)
    magabs_scattered = np.random.normal(loc=magabs, scale=sigmaint) # add intrinsic scatter
    return magabs_scattered
def magabs_to_magobs(magabs, redshift, cosmology=Planck18):
    """ distance_modulus(cosmo) = m - M """
    mu = Planck18.distmod(redshift).value
    return mu + magabs
    _____ #
# Create the dict #
modeldict = {"x1" {"func": np.random.normal, "kwargs": {"loc": 0, "scale":1}},
             "c": {"func": stats.lognorm.rvs, "kwargs": {"s": 0.7, "loc":-0.2, "scale":0.2}},
             "magabs": {"func": tripp98,
                       "kwargs": {"x1": "@x1"] "c": "@c"}},
            "redshift": {"func": draw_from_rate},
             "magobs": {"func": magabs_to_magobs,
                       "kwargs": {"magabs":"@magabs", "redshift": "@redshift"}},
     Simulate
#
mdag = modeldag.ModelDAG(modeldict)
data_sim = mdag.draw(1_000)
```

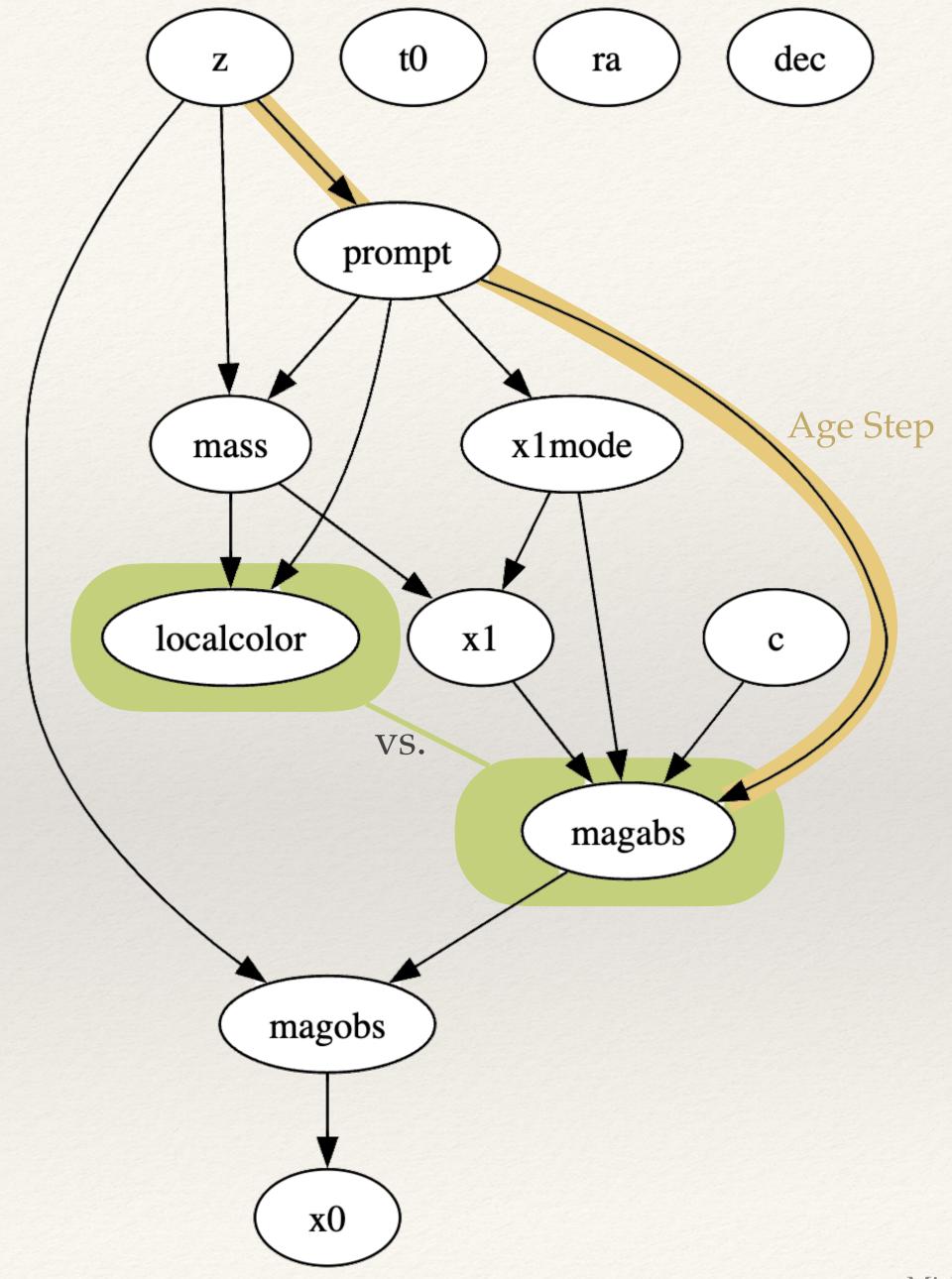
### mdag\_visualize()







## skysurvey | modeldag





snia.data												
		z	x1	С	tO	ra	dec	magabs	mwebv	magobs	x0	template
(	<b>)</b> 0.077	05	-0.030	1.221450	58487.214844	336.697937	7.605778	-15.358439	0.134286	22.429092	0.000017	salt2
	<b>1</b> 0.068	65	-0.700	-0.027583	58659.839844	60.706551	55.346775	-19.233326	0.719195	18.290983	0.000766	salt2
:	<b>2</b> 0.069	05	0.320	0.351301	58782.136719	23.629887	-0.468616	-18.197527	0.036892	19.339998	0.000292	salt2

### sn\_07pk.data

	z	tO	magabs	ra	dec	magobs	amplitude	template
0	0.04905	56041.375000	-18.313118	213.854660	-78.848862	18.451269	1.207582e-15	v19-2007pk-corr
1	0.03315	56130.003906	-16.217342	310.827576	17.159157	19.671362	3.925357e-16	v19-2007pk-corr
2	0.04845	56179.964844	-14.626013	337.447815	37.145824	22.110720	4.150845e-17	v19-2007pk-corr

### snii = skysurvey.SNeII.from\_draw(1000) snii.data

	template	z	tO	magabs	ra	dec	magobs	amplitude
0	v19-2016x-corr	0.02965	56173.551302	-17.000019	313.871430	-32.681907	18.640832	3.438910e-15
1	v19-2016bkv-corr	0.04335	56078.601439	-16.390584	294.182082	23.938362	20.096692	8.948369e-15
2	v19-asassn14jb-corr	0.02645	56139.120893	-16.300129	44.165477	-52.273437	19.087617	7.501353e-15

Customization is easy

**rate**, **redshift** range, absolute **magnitude** etc. when calling [from\_]draw()

# **Pre-Built Transients**

Transient template is built on top of sncosmo

Any **sncosmo** built-in source is available: https://sncosmo.readthedocs.io/en/stable/source-list.html

Type Ia Supernovae snia = skysurvey.SNeIa()

**Transient Time Serie** 

sn\_07pk = skysurvey.TSTransient("v19-2007pk-corr") sn\_07pk = skysurvey.TSTransient.from\_draw(1000, template="v19-2007pk-corr")

kilonova = skysurvey.Kilonova() Kilonova

Supernova Types

"Collection Transient Time Series"

snii = skysurvey.SNeII.from\_draw(1000) | SNIIn | SNIb | SNIc | SNIcBL SNIIb SNII 23 12 12 7 6 6 Vincenzi+2019

See **skysurvey** doc for how to build a transient blackbody



# pip install skysurvey

### targets

### survey

import pandas
from shapely import geometry

X

```
# footprint | 2-deg radius circle
footprint = geometry.Point(0,0).buffer(2)
```

# stored observing logs | ra, dec, zp, skynoise, gain, filter ...
data = pandas.read\_parquet("observing\_logs.parquet")

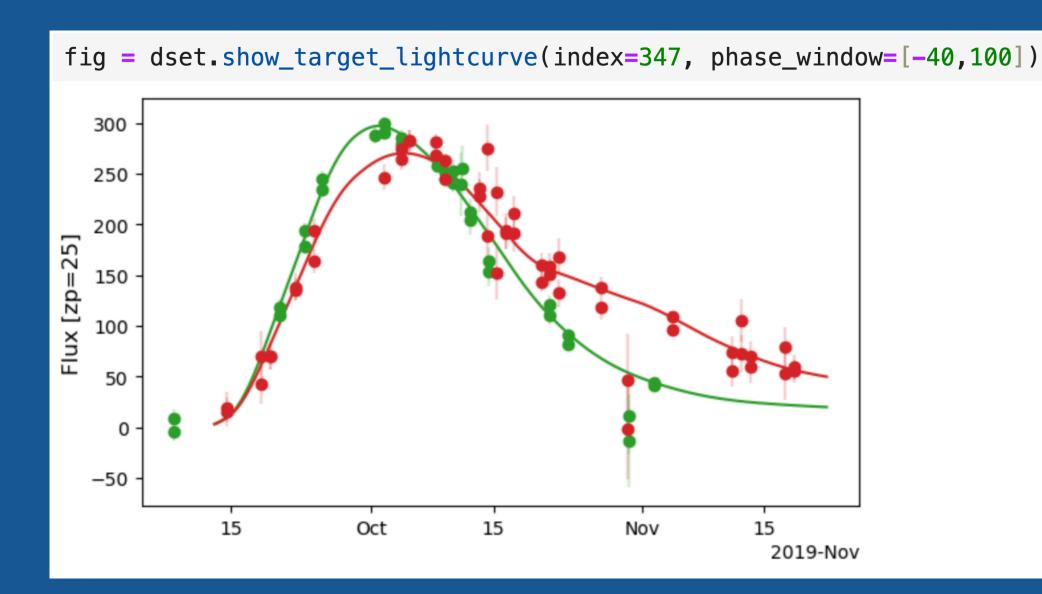
survey = skysurvey.Survey.from\_pointings(data, footprint=footprint)

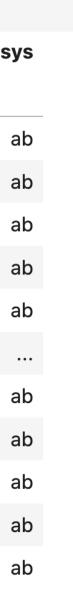
### dataset

dset = skysurvey.DataSet.from\_targets\_and\_survey(snia, ztf)

	dset	.data								
targets _ .index	→index	survey .index	fieldid	rcid	time	band	flux	fluxerr	zp	zpsy
macx	5821	26009040	1	3	59104.390625	ztfr	-39.302373	36.143482	25.233383	a
		26054356	1	3	59106.328125	ztfg	8.518026	14.717361	24.559206	a
		26331019	1	3	59112.300781	ztfg	24.598649	20.259558	26.025703	a
		26338533	1	3	59112.359375	ztfr	-3.302169	36.260182	26.127666	a
		27121580	1	3	59128.250000	ztfr	-20.931607	46.604773	25.959711	a
		•••								
	9727	30211656	1408	16	59203.179688	ztfr	14.887892	60.705860	26.204699	a
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		30841192	1408	16	59221.167969	ztfr	34.804501	22.752532	26.142845	â
		36571545	1408	16	59392.414062	ztfr	-9.636968	41.250546	26.149637	e
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<sup>8429351</sup> rows × 8 columns



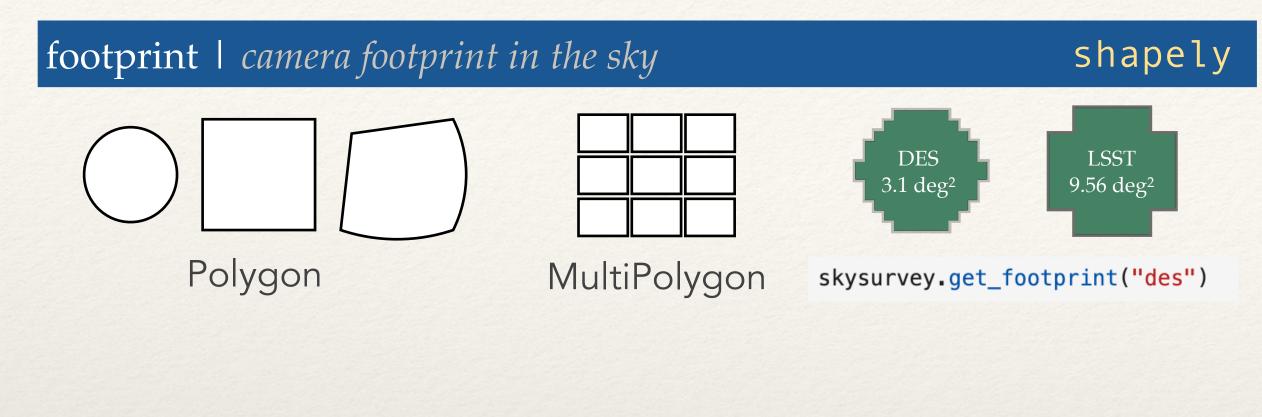








# Build your own Survey



### data | what was observed when, under which conditions

pandas

- *ra, dec* Camera pointings
- *mjd* time in Modified Julian Date
- *band* name of the filter used
- *skynoise* level of the sky noise (unit of zp)
- zp depth of the image
- gain camera gain (e-/adu), used for noise

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### Will be used by sncosmo for lightcurve realization

```
import numpy as np
from shapely import geometry
from skysurvey.tools import utils
# footprint | 2-deg radius circle
footprint = geometry.Point(0,0).buffer(2)
# fake observations
size = 100_000 # n-pointings
data = {}
data["gain"] = 1
data["zp"] = 30
data["skynoise"] = np.random.normal(size=size, loc=200, scale=20)
data["mjd"] = np.random.uniform(58_800, 59_600, size=size)
data["band"] = np.random.choice(["desg","desr","desi"], size=size)
data["ra"], data["dec"] = utils.random_radec(size=size,
                                             ra_range=[10,350],
                                             dec_range=[-50,10])
```

```
# Load the survey
survey = skysurvey.Survey.from_pointings(data, footprint=footprint)
```

```
# footprint | 2-deg radius circle
footprint = geometry.Point(0,0).buffer(2)
# fake observations
data = pandas.read_parquet("observing_logs.parquet")
```

# Load the survey survey = skysurvey.Survey.from\_pointings(data, footprint=footprint)





# Build your own GridSurvey

### footprint | *camera footprint in the sky* shapely LSST DES 9.56 deg<sup>2</sup> 3.1 deg<sup>2</sup> Polygon MultiPolygon skysurvey.get\_footprint("des")

fields | *Coordinates of the grid center* 

{fieldid: {ra: value, dec: value}}

### data | what was observed when, under which conditions

- *ra, dec* Camera pointings
- *mjd* time in Modified Julian Date
- *band* name of the filter used
- *skynoise* level of the sky noise (unit of zp)
- zp depth of the image
- gain camera gain (e-/adu), used for noise

Mickael RIGAULT

### Will be used by sncosmo for lightcurve realization

pandas

```
import numpy as np
import skysurvey
from shapely import geometry
# footprint
footprint = geometry.Point(0,0).buffer(2)
# fields
fields_radec = { 'C1': {'dec': -27.11161, 'ra': 54.274292+180},
                 'C2': {'dec': -29.08839, 'ra': 54.274292+180},
                 'C3': {'dec': -28.10000, 'ra': 52.648417+180},
                 'E1': {'dec': -43.00961, 'ra': 7.8744167+180},
                 'E2': {'dec': -43.99800, 'ra': 9.5000000+180},
                 'S1': {'dec': 0.00000, 'ra': 42.820000+180},
                 'S2': {'dec': -0.988389, 'ra': 41.194417+180},
                 'X1': {'dec': -4.929500, 'ra': 34.475708+180},
                 'X2': {'dec': -6.412111, 'ra': 35.664500+180},
                 'X3': {'dec': -4.600000, 'ra': 36.450000+180}
# observing logs
size = 100_000
data = {}
data["gain"] = 1
data["zp"] = 30
data["skynoise"] = np.random.normal(size=size, loc=200, scale=20)
data["mjd"] = np.random.uniform(58_800, 59_600, size=size)
data["band"] = np.random.choice(["desg","desr","desi"],
                                size=size)
# - no ra, dec - #
data["fieldid"] = np.random.choice( list(fields_radec.keys()),
                                   size=size)
# Load a GridSurvey
survey = skysurvey.GridSurvey.from_pointings(data, fields_radec,
                                             footprint=footprint)
```

pandas

