

(Color-)magnetic flux tubes in dense matter

A. Haber, A. Schmitt, PRD 95, 116016 (2017); EPJ Web Conf. 137, 09003 (2017)

A. Haber, A. Schmitt, JPG 45, 065001 (2018)



Theoretical motivation

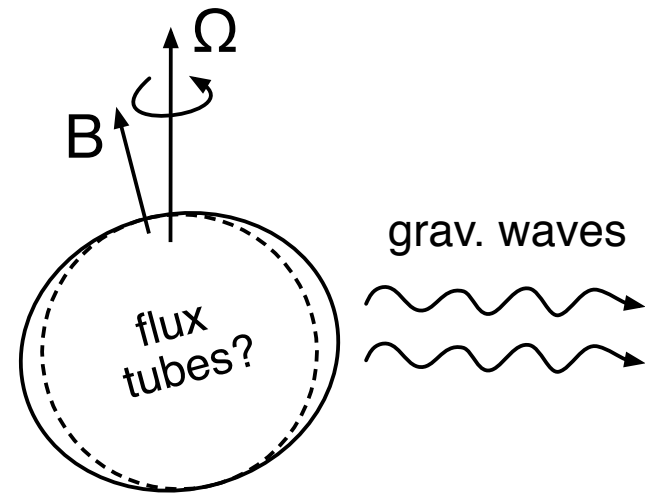
- What's the behavior of a **multi-component superconductor** regarding type-I/type-II?
- What (color-)flux tubes are there in a **non-abelian superconductor** like color-flavor locked (CFL) quark matter?

Phenomenological motivation

- Do (color-)flux tubes support **ellipticities of neutron stars**
→ **gravitational waves**?

K. Glampedakis, D. I. Jones and L. Samuelsson,
PRL 109, 081103 (2012)

(besides magnetic forces, see also **crystalline structures**)

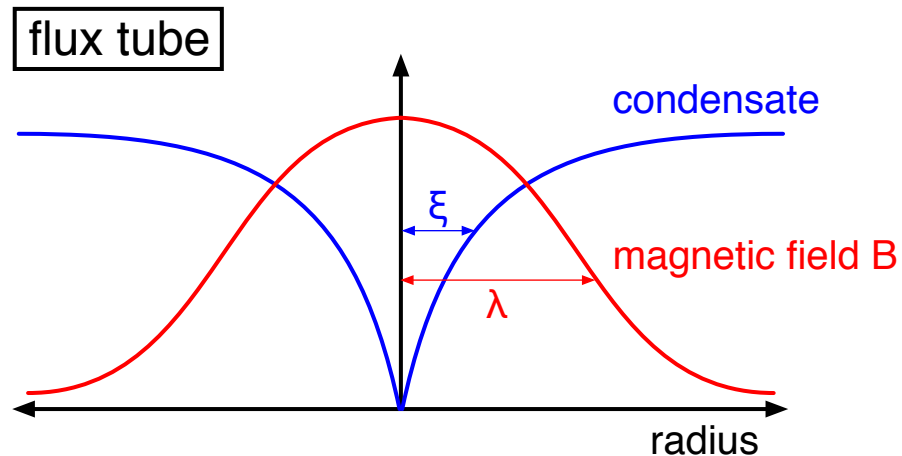


Single-component superconductor

- Ginzburg-Landau potential U for complex field ϕ with charge q coupled to gauge field A^μ

$$U = \frac{\mathbf{B}^2}{2} + |(\nabla + iq\mathbf{A})\phi|^2 - \mu^2|\phi|^2 + \lambda|\phi|^4$$

Reminder: type-I/type-II superconductivity

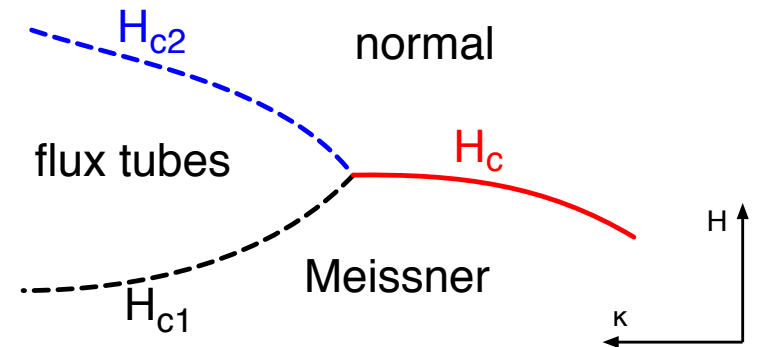


- Ginzburg-Landau parameter

$$\kappa = \frac{\lambda}{\xi}$$

- type-II superconductivity
for $\kappa > 1/\sqrt{2}$: flux tube lattice
for $H_{c1} < H < H_{c2}$

A.A. Abrikosov, Soviet Physics JETP 5, 1174 (1957)



Multi-component superconductors (page 1/2)

- two fields with charges q_1, q_2 (neutron/proton: $q_1 = 2e, q_2 = 0$)

$$U = \frac{\mathbf{B}^2}{2} + \sum_{i=1,2} \left[|(\nabla + iq_i \mathbf{A})\phi_i|^2 - \mu_i^2 |\phi_i|^2 + \lambda_i |\phi_i|^4 \right] + 2h |\phi_1|^2 |\phi_2|^2$$

- fields are coupled indirectly via gauge field (if both $q_1, q_2 \neq 0$) and directly with coupling h (neutron/proton system: additional derivative coupling [A. Haber, A. Schmitt, PRD 95, 116016 \(2017\)](#))

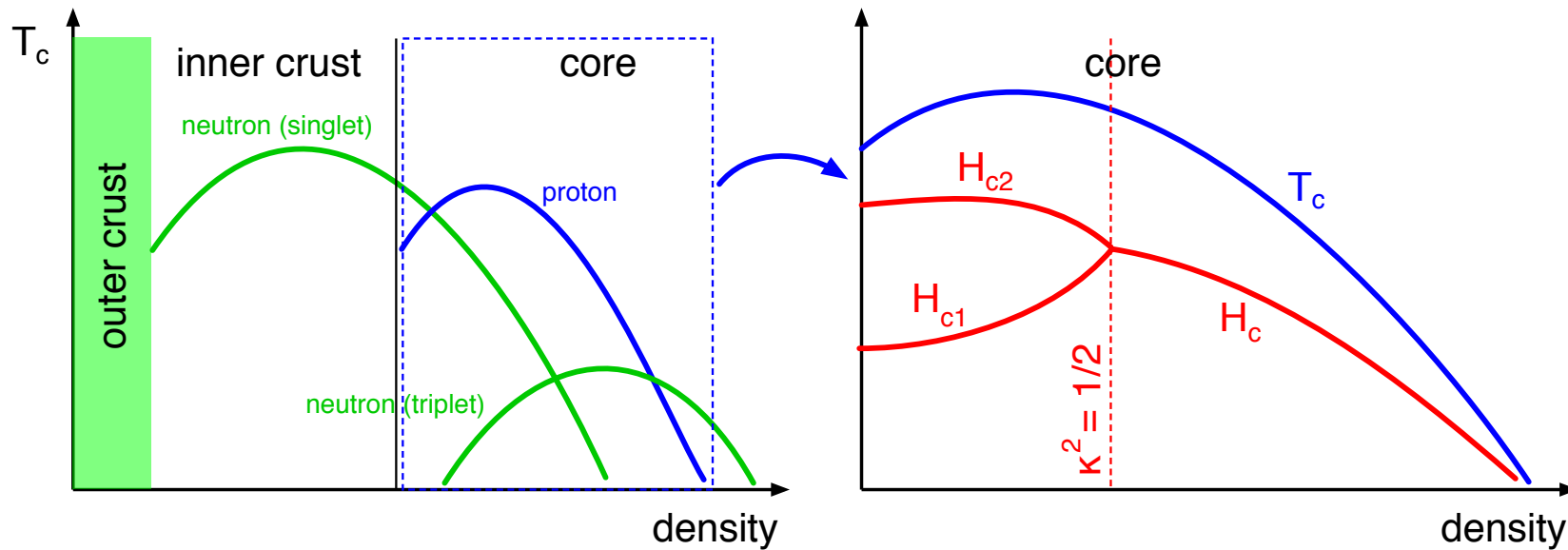
Multi-component superconductors (page 2/2)

- more fields and multiple (color-)gauge fields

$$U = \frac{\mathbf{B}_1^2}{2} + \frac{\mathbf{B}_2^2}{2} + \sum_{i=1}^3 \left[|(\nabla + iq_{i1}\mathbf{A}_1 + iq_{i2}\mathbf{A}_2)\phi_i|^2 - \mu^2|\phi_i|^2 + \lambda|\phi_i|^4 \right]$$
$$- 2h(|\phi_1|^2|\phi_2|^2 + |\phi_1|^2|\phi_3|^2 + |\phi_2|^2|\phi_3|^2)$$

- color superconductor: 3 scalar components
and 3 gauge fields: 1 electromagnetic and 2 color fields

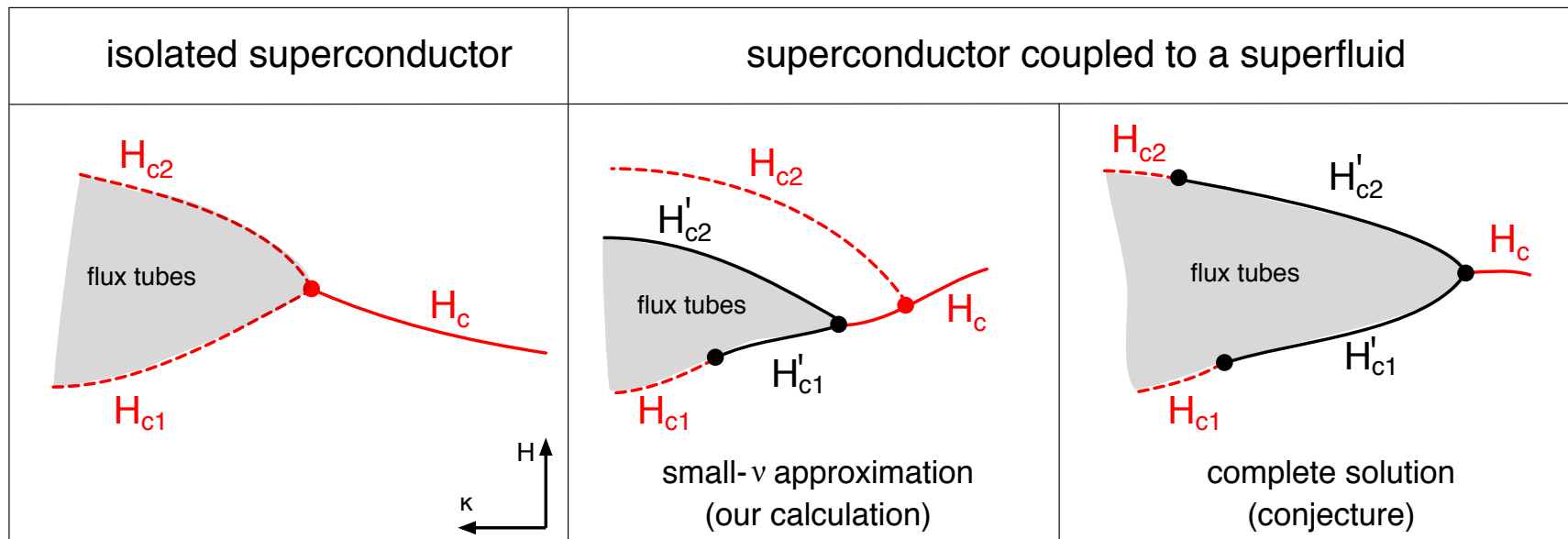
Two-component superconductors in neutron star cores



- density-dependent κ
- type-I/type-II transition in the core?
- effect of coupling to superfluid on type-I/type-II transition?

Critical magnetic fields in a two-component system

- compute flux tube profiles and flux tube interaction
→ attractive long-distance interaction in type-II regime



numerical calculation supports conjecture

A. Haber, D. Müller, preliminary results

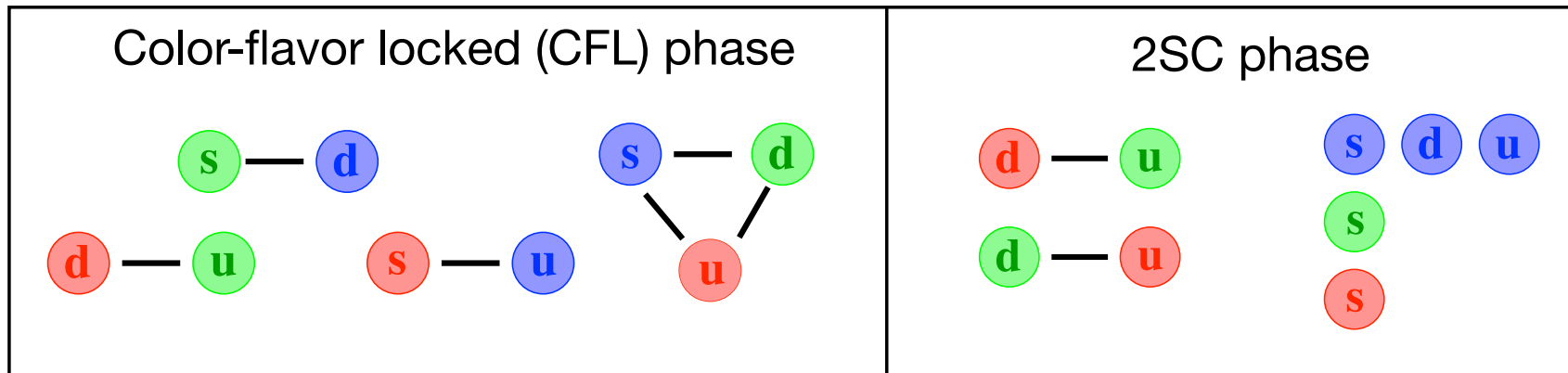
- first-order phase transition allows for flux tube clusters
see also "type-1.5" superconductors J. Carlström, J. Garaud, E. Babaev, PRB 84, 134515 (2011)

Ginzburg-Landau potential for color superconductors

$$U = -\text{Tr}[(D_\mu \Phi)^\dagger (D^\mu \Phi)] + a \text{Tr}[\Phi^\dagger \Phi] + b \text{Tr}[(\Phi^\dagger \Phi)^2] + c (\text{Tr}[\Phi^\dagger \Phi])^2 + \frac{1}{4} F_{\mu\nu}^a F_a^{\mu\nu} + \frac{1}{4} F_{\mu\nu} F^{\mu\nu}$$

$$\Phi = \begin{pmatrix} \phi & 0 & 0 \\ 0 & \phi & 0 \\ 0 & 0 & \phi \end{pmatrix} \quad \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & \phi \end{pmatrix} \quad \begin{pmatrix} \phi_1(\vec{r}) & 0 & 0 \\ 0 & \phi_2(\vec{r}) & 0 \\ 0 & 0 & \phi_3(\vec{r}) \end{pmatrix}$$

CFL 2SC ansatz for flux tubes



Mixing of gluons and photons in CFL

- symmetry breaking pattern of CFL

$$[SU(3)]_c \times \underbrace{SU(3)_L \times SU(3)_R}_{\supset [U(1)]_Q} \times U(1)_B \rightarrow \underbrace{SU(3)_{c+L+R}}_{\supset [U(1)]_{\tilde{Q}}} \times \mathbb{Z}_2$$

- CFL is a superfluid \rightarrow rotational vortices
- Meissner effect for gluons T_1, \dots, T_7 ("color superconductor")
- all Cooper pairs neutral under $\tilde{Q} = Q + \frac{2}{\sqrt{3}}T_8$
and (differently) charged under orthogonal combination \tilde{T}_8
 - \tilde{Q} -magnetic field penetrates CFL
 - Meissner effect for \tilde{T}_8 -magnetic field

(Analogous to gauge field mixing in standard model, $[SU(2)] \times [U(1)] \rightarrow [U(1)]_Q$)

Ginzburg-Landau potential with rotated gauge fields

$$\begin{aligned}
 U = & \frac{\tilde{\mathbf{B}}^2}{2} + \frac{\mathbf{B}_3^2}{2} + \frac{\tilde{\mathbf{B}}_8^2}{2} + \left| \left(\nabla + i\frac{g}{2}\mathbf{A}_3 + i\tilde{g}_8\tilde{\mathbf{A}}_8 \right) \phi_1 \right|^2 + \left| \left(\nabla - i\frac{g}{2}\mathbf{A}_3 + i\tilde{g}_8\tilde{\mathbf{A}}_8 \right) \phi_2 \right|^2 \\
 & + \left| \left(\nabla - 2i\tilde{g}_8\tilde{\mathbf{A}}_8 \right) \phi_3 \right|^2 - \mu^2(|\phi_1|^2 + |\phi_2|^2 + |\phi_3|^2) + \lambda(|\phi_1|^4 + |\phi_2|^4 + |\phi_3|^4) \\
 & - 2h(|\phi_1|^2|\phi_2|^2 + |\phi_1|^2|\phi_3|^2 + |\phi_2|^2|\phi_3|^2)
 \end{aligned}$$

- approximations
 - massless quarks, $m_s = m_d = m_u = 0$
 - purely bosonic approach \rightarrow neglect effects of magnetic field on Cooper pair constituents
 - Ginzburg-Landau parameters μ, λ, h : (mostly) use perturbative results and extrapolate down in density (= to large couplings)

Ginzburg-Landau potential with rotated gauge fields

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 \end{aligned}$$

- ansatz for flux tube solutions

$$\phi_i(r, \theta) = \rho_i(r) e^{in_i\theta}$$

with winding numbers n_1, n_2, n_3

- solve equations of motion for $\rho_1, \rho_2, \rho_3, \mathbf{A}_3, \tilde{\mathbf{A}}_8$
- boundary conditions: homogeneous CFL far away from flux tube, $\rho_i = 0$ in center if winding n_i is nonzero

• usually: vortex \rightarrow baryon circulation $\Gamma = \oint d\ell \cdot \mathbf{v}$

flux tube \rightarrow magnetic flux $\Phi = \oint d\ell \cdot \mathbf{A}$

• CFL line defects can have both!

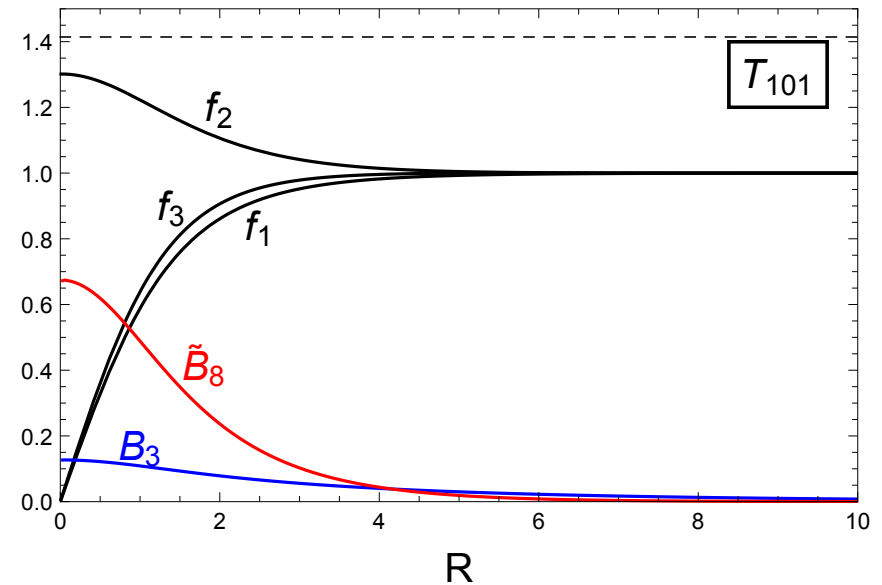
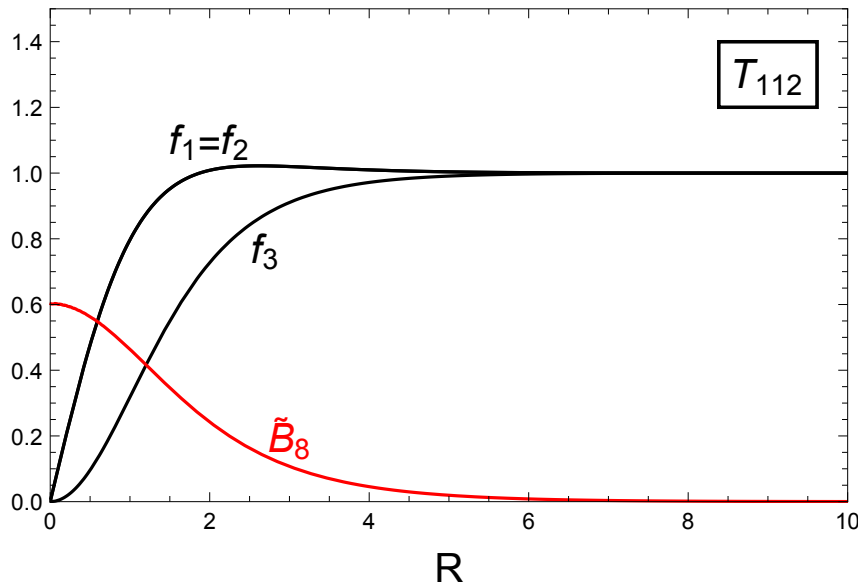
$$\Gamma \propto n_1 + n_2 + n_3, \quad \Phi_3 \propto n_1 - n_2, \quad \tilde{\Phi}_8 \propto n_1 + n_2 - 2n_3$$

CFL line defects	(n_1, n_2, n_3)	$\Gamma [\pi/3\mu_q]$	$\Phi_3 [\pi/g]$	$\tilde{\Phi}_8 [\pi/\tilde{g}_8]$
Global vortex Forbes, Zhitnitsky (2002)	(n, n, n)	$-n$	0	0
”Semi-superfluid” vortex Balachandran, Digal, Matsuura (2006)	$(0, 0, n)$	$-\frac{n}{3}$	0	$\frac{2n}{3}$
Magnetic flux tube T_{112} Iida (2005)	$(n, n, -2n)$	0	0	$-2n$
Magnetic flux tube T_{101} Haber, Schmitt (2018)	$(n, 0, -n)$	0	$-n$	$-n$

- vortices ($\Gamma \neq 0$): "topological" since $\pi_1[U(1)] = \mathbb{Z}$
 - global vortex decays into 3 semi-superfluid vortices
M. G. Alford, S. K. Mallavarapu, T. Vachaspati and A. Windisch, PRC 93, 045801 (2016)
- flux tubes ($\Gamma = 0$): "non-topological" since $\pi_1[SU(3)] = 0$
 - stabilized through external magnetic field \rightarrow see rest of the talk

CFL line defects	(n_1, n_2, n_3)	$\Gamma [\pi/3\mu_q]$	$\Phi_3 [\pi/g]$	$\tilde{\Phi}_8 [\pi/\tilde{g}_8]$
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Magnetic flux tube T_{112} Iida (2005)	$(n, n, -2n)$	0	0	$-2n$
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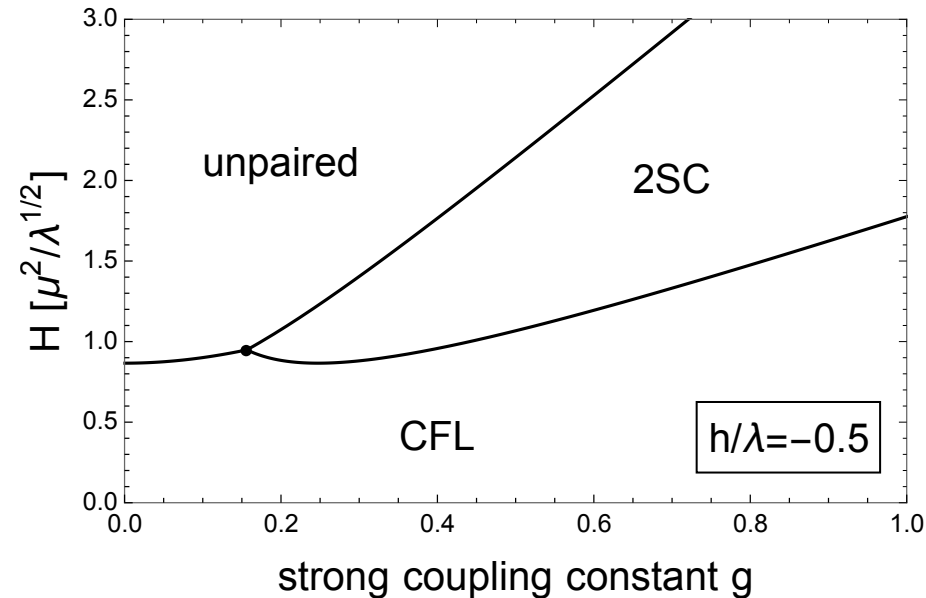
Flux tube profiles



- flux tube with "unpaired core"
K. Iida, PRD 71, 054011 (2005)

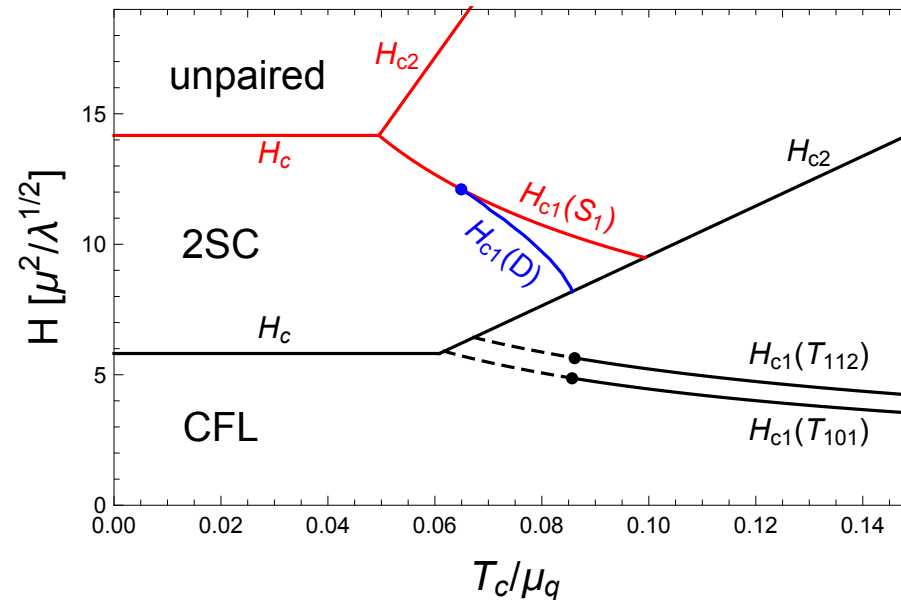
- flux tube with "2SC core"
A. Haber, A. Schmitt, JPG 45, 065001 (2018)
- additional B_3 field
(cost in free energy)
- non-vanishing condensate in
core (gain in free energy)

Phase structure without flux tubes



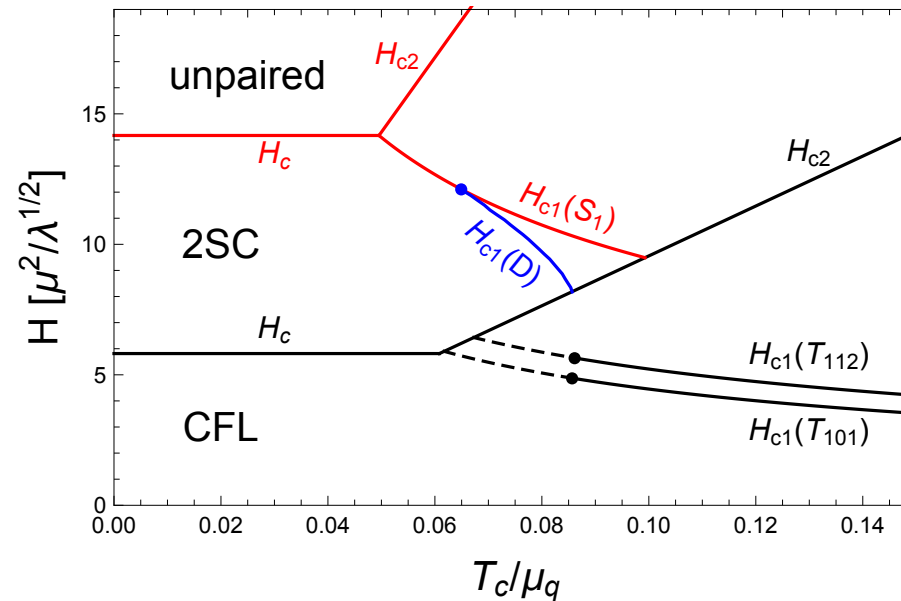
- in weak coupling: $h/\lambda = -0.5$
- CFL superseded by 2SC except for small values of strong coupling constant g (even though here $m_s \simeq 0!$)
- in neutron stars $\mu_q \simeq 400 \text{ MeV} \Rightarrow g \simeq 3.5$

Phase diagram including flux tubes



- type-II regime for sufficiently large T_c/μ_q
(neutron stars: $T_c/\mu_q \sim (10 - 100)/(400 - 500) \sim 0.1$)
- type-I/type-II transition complicated (multi-component structure!)

Phase diagram including flux tubes



- CFL flux tubes with 2SC core (T_{101}) preferred
- critical fields $H \sim 10^{19} \text{ G} \gg H_{\text{NS}}$, however:
creation of flux tubes through cooling into superconducting phase?
- 2SC domain walls (D) preferred over ordinary 2SC flux tubes (T_1)
for sufficiently large T_c/μ_q

Summary

- multi-component superconductors have a nontrivial type-I/type-II transition
- dense quark matter is a multi-component superconductor and has various possible line defects
- CFL flux tubes (without baryon circulation) are not protected by topology, but can be stabilized by a magnetic field
- we have found new magnetic defects in CFL and 2SC:
CFL tubes with 2SC core and 2SC domain walls
- defects in superconducting/superfluid nuclear and quark matter are relevant for neutron star observables, e.g., gravitational waves